



Infrared divergences in QED revisited

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Recently, it has been shown that the vacuum state in QED is infinitely degenerate. Moreover, a transition among the degenerate vacua is induced in any nontrivial scattering process and determined from the associated soft factor. Conventional computations of scattering amplitudes in QED do not account for this vacuum degeneracy and therefore always give zero. This vanishing of all conventional QED amplitudes is usually attributed to infrared divergences. Here, we show that if these vacuum transitions are properly accounted for, the resulting amplitudes are nonzero and infrared finite. Our construction of finite amplitudes is mathematically equivalent to, and amounts to a physical reinterpretation of, the 1970 construction of Faddeev and Kulish.

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I. INTRODUCTION

Recently, it has been shown [1–5] (see Ref. [6] for a review) that the IR sector of all Abelian gauge theories, including QED, is governed by an infinite-dimensional symmetry group. The symmetry group is generated by large gauge transformations that approach angle-dependent constants at null infinity. The soft photon theorem is the matrix element of the associated conservation laws. This large gauge symmetry is spontaneously broken, resulting in an infinite vacuum degeneracy.

QED has been tested to 16 decimal places and is the most accurate theory in the history of human thought. The preceding statements have no mathematically new content within QED and certainly do not imply errors in any previous QED calculations. However, as emphasized herein, they do perhaps provide a physically illuminating new way of describing the IR structure. Moreover, generalizations of this perspective to other contexts have led to a variety of truly new mathematical relations in both gauge theory and gravity [6].

One of the puzzling features of the IR structure of QED is the appearance of IR divergences.¹ These divergences set all conventional Fock-basis $\mathcal{S}$ -matrix elements to zero. Often, they are dealt with by restricting to inclusive cross sections in which physically unmeasurable photons below some IR cutoff are traced over. The trace gives a divergence which offsets the zero and yields a finite result for the physical measurement [9–12]. While this is adequate for most experimental applications, for many purposes, it is

nice to have an $\mathcal{S}$ -matrix.² For example, precise discussions of unitarity or symmetries require an $\mathcal{S}$ -matrix.

It is natural to ask if the newly discovered IR symmetries are related to the IR divergences of the $\mathcal{S}$ -matrix. We will see that the answer is yes. The conservation laws imply that every nontrivial scattering process is necessarily accompanied by a transition among the degenerate vacua. Conventional QED $\mathcal{S}$ -matrix analyses tend to assume that the vacuum is unique and hence that the initial and final vacua are the same. Since this violates the conservation laws, the Feynman diagrammatics give a vanishing result. This is usually attributed to “IR divergences,” but we feel that this phrase is something of a misnomer. Rather, zero is the correct physical answer. The vanishing of the amplitudes is a penalty for not accounting for the required vacuum transition. In this paper, we allow for vacuum transitions to occur and find that the resulting amplitudes are perfectly IR finite and generically nonvanishing when the conservation laws are obeyed.

Although we have phrased this result in a way that sounds new, the mathematics behind it is not new. We have merely rediscovered the 1970 formulas [16–20] of Faddeev and Kulish (FK) and others, who showed that certain dressings of charges by clouds of soft photons yield IR finite scattering amplitudes. The FK dressings implicitly generate precisely the required shift between degenerate vacua.

While our formulas are not new,³ our physical interpretation is new. One may hope that the new physical insight will enable a construction of IR finite $\mathcal{S}$ -matrices for unconfined

¹Originally, these were found by looking at the spectrum and number of photons produced by particles undergoing acceleration. Mott [7] looked at corrections to Rutherford scattering as a result of the emission of photons. Bloch and Nordsieck [8] examined the spectrum of photons produced by a small change in the velocity of an electron.

²It may also be challenging to describe experimental measurements of the electromagnetic memory effect [13–15] in a theory with a finite IR cutoff.

³Except for, in Sec. VI, a conjectured generalization of the FK IR divergence cancellation mechanism to amplitudes involving some undressed charges but still obeying the conservation laws. An example given there is e^+e^- scattering with no incoming radiation.

non-Abelian gauge theory and also have useful applications to gravity.

Related discussions of the FK construction in the context of large gauge symmetry have appeared in Refs. [21–24].

II. VACUUM SELECTION RULES

In this section, we review the derivation of and formulas for the vacuum transitions induced by the scattering of charged massless particles. We refer the reader to Refs. [1–3,25] for further details. The conceptually similar massive case is treated in Sec. V. Incoming states are best described in advanced coordinates in Minkowski space,

$$ds^2 = -dv^2 + 2dvdr + 2r^2\gamma_{z\bar{z}}dzd\bar{z}, \quad (2.1)$$

while outgoing states employ retarded coordinates,

$$ds^2 = -du^2 - 2dudr + 2r^2\gamma_{z\bar{z}}dzd\bar{z}. \quad (2.2)$$

Here, $\gamma_{z\bar{z}} = \frac{2}{(1+z\bar{z})^2}$ is the unit round metric on S^2 , and $u = t - r$ ($v = t + r$) is the retarded (advanced) time. The z coordinates used in the advanced and retarded coordinate systems differ by an antipodal map on S^2 . In and out states are characterized by the charges⁴

$$\begin{aligned} Q_\varepsilon^- &= \frac{1}{e^2} \int_{\mathcal{I}_+^-} d^2w \gamma_{w\bar{w}} \varepsilon F_{rv}, \\ Q_\varepsilon^+ &= \frac{1}{e^2} \int_{\mathcal{I}_+^+} d^2w \gamma_{w\bar{w}} \varepsilon F_{ru}, \end{aligned} \quad (2.3)$$

where F is the electromagnetic field strength, $\mathcal{I}_+^-$ is the future boundary of $\mathcal{I}^-$, $\mathcal{I}_+^+$ is the past boundary of $\mathcal{I}^+$, and ε is any function on S^2 . The conservation law for these charges

$$\langle \text{out} | (Q_\varepsilon^+ \mathcal{S} - \mathcal{S} Q_\varepsilon^-) | \text{in} \rangle = 0 \quad (2.4)$$

is implied by the soft photon theorem. In and out soft photon modes are defined as integrals of the radiative part of F over the null generators of past and future null infinity ($\mathcal{I}^\pm$) according to

$$\int_{-\infty}^{\infty} du F_{uz} \equiv N_z^+, \quad (2.5)$$

$$\int_{-\infty}^{\infty} dv F_{vz} \equiv N_z^-. \quad (2.6)$$

Choosing $\varepsilon(w, \bar{w}) = \frac{1}{z-w}$, Eq. (2.4) can be written in the form

⁴In these and the following equations, $(F_{rv}, F_{ru}, j_v, j_u)$ denote the coefficient of the leading $\mathcal{O}(\frac{1}{r^2})$ term of the large r field expansions, while (F_{vz}, F_{uz}) denote the leading $\mathcal{O}(r^0)$ terms.

$$\langle \text{out} | (N_z^+ \mathcal{S} - \mathcal{S} N_z^-) | \text{in} \rangle = \Omega_z^{\text{soft}} \langle \text{out} | \mathcal{S} | \text{in} \rangle, \quad (2.7)$$

where the soft factor is

$$\Omega_z^{\text{soft}} = \Omega_z^{\text{soft}-} - \Omega_z^{\text{soft}+}, \quad (2.8)$$

$$\Omega_z^{\text{soft}-} = \frac{e^2}{4\pi} \sum_{k \in \text{in}} \frac{Q_k}{z - z_k}, \quad \Omega_z^{\text{soft}+} = \frac{e^2}{4\pi} \sum_{k \in \text{out}} \frac{Q_k}{z - z_k}. \quad (2.9)$$

Here, Q_k and z_k denote the charges of the asymptotic particles and the angles at which they enter or exit at $\mathcal{I}^\pm$. Degenerate incoming vacua can be characterized by their N_z^- eigenvalue:

$$N_z^-(z, \bar{z}) | N_z^{\text{in}} \rangle = N_z^{\text{in}}(z, \bar{z}) | N_z^{\text{in}} \rangle. \quad (2.10)$$

Let us consider special states denoted $|\text{in}; N_z^{\text{in}}\rangle$ comprised of finite numbers of noninteracting incoming charged particles and hard photons built by acting with asymptotic creation operators on eigenstates (2.10) of N_z^- .⁵ Such hard particles do not affect the zero modes and hence obey

$$N_z^- |\text{in}; N_z^{\text{in}}\rangle = N_z^{\text{in}} |\text{in}; N_z^{\text{in}}\rangle. \quad (2.11)$$

Adopting a similar notation for out states, Eq. (2.7) becomes

$$(N_z^{\text{out}} - N_z^{\text{in}}) \langle \text{out}; N_z^{\text{out}} | \mathcal{S} | N_z^{\text{in}}; \text{in} \rangle = \Omega_z^{\text{soft}} \langle \text{out}; N_z^{\text{out}} | \mathcal{S} | N_z^{\text{in}}; \text{in} \rangle. \quad (2.12)$$

We conclude that either

$$\langle \text{out}; N_z^{\text{out}} | \mathcal{S} | N_z^{\text{in}}; \text{in} \rangle = 0 \quad (2.13)$$

or

$$N_z^{\text{out}} - N_z^{\text{in}} = \Omega_z^{\text{soft}}. \quad (2.14)$$

The second relation (2.14) expresses conservation of the charges—one for each point on the sphere—associated to large gauge symmetries. The first states that any amplitude violating the conservation law must vanish.

In conventional formulations of QED, the vacuum is presumed to be unique.⁶ In that case, Eq. (2.14) is not an option, and we conclude that, according to (2.13), all $\mathcal{S}$ -matrix elements vanish. In fact, this result is well known and attributed to IR divergences. We see here that the IR divergences which set all such amplitudes to zero can be

⁵It is assumed here that the Fourier coefficients of the photon creation operators are finite as the frequency $\omega \rightarrow 0$.

⁶Since N_z manifestly carries zero energy, if the vacuum is assumed to be unique, it would have to be an N_z eigenstate.

understood as a penalty for neglecting the fact that the in and out vacua differ for every nontrivial scattering process. Armed with this insight, we will construct a natural and IR finite set of scattering amplitudes.

The necessity for vacuum transitions in any scattering process follows from the constraint equations on $\mathcal{I}^+$,

$$\partial_u F_{ru} + D^z F_{uz} + D^{\bar{z}} F_{u\bar{z}} + e^2 j_u = 0, \quad (2.15)$$

and $\mathcal{I}^-$,

$$\partial_v F_{rv} - D^z F_{vz} - D^{\bar{z}} F_{v\bar{z}} - e^2 j_v = 0. \quad (2.16)$$

Assuming that the electric field vanishes in the far past and far future and using the matching conditions⁷

$$F_{ru}|_{\mathcal{I}^+} = F_{rv}|_{\mathcal{I}^+}, \quad (2.17)$$

$$A_z|_{\mathcal{I}^+} = A_z|_{\mathcal{I}^+}, \quad (2.18)$$

the divergence of (2.14) (and its complex conjugate) is the sum of the integrals of (2.15) and (2.16).

Let us examine the classical electromagnetic field configuration needed to satisfy the constraints. A single charge Q_0 particle incoming at $(v_0, z_0, \bar{z}_0)$ corresponds to

$$j_v = Q_0 \delta(v - v_0) \gamma^{z\bar{z}} \delta^{(2)}(z - z_0). \quad (2.19)$$

We write the state consisting of one such particle in the $N_z^{\text{in}} = 0$ vacuum as

$$|z_0; 0\rangle, \quad N_z^- |z_0; 0\rangle = 0. \quad (2.20)$$

We can solve the constraints for finite z either using the Coulombic modes with⁸

$$F_{rv} = Q_0 e^2 \theta(v - v_0) \gamma^{z\bar{z}} \delta^{(2)}(z - z_0) \quad (2.21)$$

or with the radiative modes⁹

$$\begin{aligned} A_z &= -\frac{Q_0 e^2}{4\pi} \partial_z G(z, z_0) \theta(v - v_0), \\ F_{vz} &= -\frac{Q_0 e^2}{4\pi} \partial_z G(z, z_0) \delta(v - v_0), \end{aligned} \quad (2.22)$$

where

$$\partial_z \partial_{\bar{z}} G(z, w) = 2\pi \delta^{(2)}(z - w). \quad (2.23)$$

⁷Here, we consider theories with no magnetic charges so that $F_{z\bar{z}}|_{\mathcal{I}^+} = 0 = F_{z\bar{z}}|_{\mathcal{I}^-}$.

⁸ $\theta(v) = 1$ for $v > 0$ and vanishes otherwise.

⁹Note that the “soft charge” N_z^- vanishes for the Coulombic dressing, while for the radiative dressing, all Q_ε^- reduce to multiples of the global charge for which $\varepsilon = 1$.

For finite z , the choice

$$G(z, w) = \ln |z - w|^2 \quad (2.24)$$

gives simply

$$A_z = -\frac{Q_0 e^2}{4\pi(z - z_0)} \theta(v - v_0). \quad (2.25)$$

The purely Coulombic choice will violate the matching conditions (2.17) unless the outgoing state also has Coulomb fields at $z = z_0$, where there may not even be any particles on $\mathcal{I}^+$. We first consider the radiative dressing (2.25). This potential is pure gauge except at advanced time $v = v_0$ where a radiative shock wave emerges. There is a shift in the flat gauge connection between the boundaries $\mathcal{I}_+^-$ and $\mathcal{I}_-^-$ of $\mathcal{I}^-$ given by $N_z^- = -\frac{Q_0 e^2}{4\pi(z - z_0)}$.

Of course, more general solutions of the constraints, which do involve Coulomb fields, are possible and can be obtained by adding to (2.25) any solution of the source-free equation. Indeed, the difference between (2.21) and (2.25) is such a solution. We will return to the more general case in Secs. V and VI.

The Green function G in (2.24) leads to image charges at $z = \infty$. Since there are no physical charges presumed at this point and we must preserve the constraints, delta function “wires” of nonzero F_{vr} are added, connecting the images at various values of v_k where particles enter. One such wire with net integral $\sum_{k \in \text{in}} Q_k$ will cross to $\mathcal{I}^+$. Overall charge conservation guarantees, if a similar construction is used to satisfy the $\mathcal{I}^+$ constraints, that this will match with the $z = \infty$ wire on $\mathcal{I}^+$.

Of course, these wires can be smoothed out by adding source-free solutions of the free Maxwell equation.¹⁰ For example, we can use

$$G(z, w) = \ln[|z - w|^2 (1 + z\bar{z})^{-1} (1 + w\bar{w})^{-1}], \quad (2.26)$$

which obeys $2\partial_z \partial_{\bar{z}} G = 4\pi \delta^{(2)}(z - w) - \gamma_{z\bar{z}}$. This effectively spreads the image charges, and along with them the F_{vr} flux wires required for their cancellation, evenly over the sphere.

For our purposes, we are primarily interested in the structure near $z = z_0$ which has the same singularity for all the G ’s. The choice of G will not be central, and we focus on the simplest one (2.24).

III. DRESSED QUANTUM STATES

The story of dressed charges began with Dirac [27], who realized that part of the problem with the formulation of quantum electrodynamics was that conventional states for charged particles were not gauge invariant. Suppose one is

¹⁰There may be a preferred Lorentz covariant dressing if the charged particles are taken as conformal primaries rather than plane waves as in Ref. [26].

considering the Dirac field for an electron, $\psi(x)$. Then, one usually thinks of the operator $\psi(x)$ as creating an electron at the point x . In classical physics, if one makes a gauge transformation,

$$A_\mu \rightarrow A_\mu + \partial_\mu \varepsilon(x), \quad (3.1)$$

then the gauge transformation of a field with charge Q_0 is

$$\psi(x) \rightarrow e^{iQ_0 \varepsilon(x)} \psi(x). \quad (3.2)$$

Dirac made a field invariant under gauge transformations which die at infinity by introducing a dressing of the charged particle. One replaces $\psi(x)$ by the gauge invariant $\psi^*(x)$ defined by

$$\psi^*(x) = \psi(x) e^{iQ_0 \int A^\mu(x') C_\mu(x') d^4 x'}. \quad (3.3)$$

$C_\mu(x)$ is then required to obey the equation

$$\partial_\mu C^\mu = \delta^{(4)}(x - x'). \quad (3.4)$$

Solutions to this equation are of course not unique, since we can add to it any solution of the homogeneous equation. Thus, Dirac's prescription is not unique and may involve either radiative or Coulomb modes depending on how C^μ is chosen.

In quantum field theory in the Schrödinger picture, operators are time independent, and so one would replace these expressions by the corresponding noncovariant forms in which only the spatial components of A_μ and C^μ are used. The integral is then taken over a three-dimensional spatial section of spacetime, and the four-dimensional delta function is replaced by the three-dimensional delta function. Thus,

$$\psi^*(x) = \psi(x) e^{iQ_0 \int A^i(x') C_i(x') d^3 x'}, \quad (3.5)$$

and

$$\partial_i C^i = \delta^{(3)}(x - x'). \quad (3.6)$$

This prescription replaces the bare electron by an electron together with an electromagnetic cloud. It is important to note that the operator (3.3) is only invariant under small gauge transformations vanishing sufficiently quickly at infinity, since an integration by parts is needed in order to demonstrate invariance. Under large gauge transformations, Dirac's operators transform with a phase, as charged operators should. Dirac provided an illuminating example of a C_i that satisfies $\partial_i C^i = \delta^{(3)}(x - x')$. This is just

$$C_i = -\partial_i \left\{ \frac{1}{4\pi|x - x'|} \right\}. \quad (3.7)$$

C_i is then just the electric Coulomb field of a point charge.

In the full interacting quantum theory, the radiative modes of the electromagnetic field obey the exact $\mathcal{I}^-$ commutator

$$[A_w(v, w, \bar{w}), F_{v'\bar{z}}(v', z, \bar{z})] = \frac{ie^2}{2} \delta(v - v') \delta^{(2)}(w - z). \quad (3.8)$$

The commutators of Coulombic modes are then, according to Dirac, whatever they must be in order that the constraints (2.16) are satisfied. That is, the operator F_{rv} is defined by

$$F_{rv} \equiv \int_{-\infty}^v dv' (D^z F_{v'z} + D^{\bar{z}} F_{v'\bar{z}} + e^2 j_{v'}), \quad (3.9)$$

where the constant of integration is set (for massless charges only) by demanding that the Coulomb field vanishes in the far past. Its commutators are then computed using (3.8) along with those for the matter fields appearing in j_v .

The coherent quantum state corresponding to (2.25) is, up to a large gauge transformation,

$$|z_0; 0\rangle_{\text{dressed}} \equiv e^{iR_0} |z_0; 0\rangle, \quad R_0 \equiv \frac{Q_0}{2\pi} \int d^2 w \gamma_{w\bar{w}} G(z_0, w) D \cdot A(v_0, w, \bar{w}), \quad (3.10)$$

where $D \cdot A \equiv D^w A_w + D^{\bar{w}} A_{\bar{w}}$. We may describe this as a charged particle surrounded by a cloud of soft photons. It easily follows from (3.8) and (3.9) that $|z_0; 0\rangle_{\text{dressed}}$ obeys the constraints without any Coulombic F_{vr} wires extending out of the charge, and the matching condition (2.17) is trivially satisfied. The dressing shifts the action of F_{vz} on states by

$$[F_{vz}(v, z, \bar{z}), iR_0] = -\frac{Q_0 e^2 \delta(v - v_0)}{4\pi(z - z_0)}. \quad (3.11)$$

We see explicitly that the early and late vacua on $\mathcal{I}^-$ differ by a large gauge transformation and

$$N_z^- |z_0; 0\rangle_{\text{dressed}} = -\frac{Q_0 e^2}{4\pi(z - z_0)} |z_0; 0\rangle_{\text{dressed}}. \quad (3.12)$$

Had we started with a vacuum state with nonzero N_z^{in} , the dressing would have simply shifted the eigenvalue.

The dressed single particle state (3.10) is easily generalized to a multiparticle state,

$$|{\text{in}}; 0\rangle_{\text{dressed}} \equiv e^{iR} |{\text{in}}; 0\rangle, \quad R \equiv \frac{1}{2\pi} \int dv d^2 w \gamma_{w\bar{w}} d^2 z \gamma_{z\bar{z}} j_v(v, z, \bar{z}) \times G(z, w) D \cdot A(v, w, \bar{w}). \quad (3.13)$$

Also, for outgoing states,

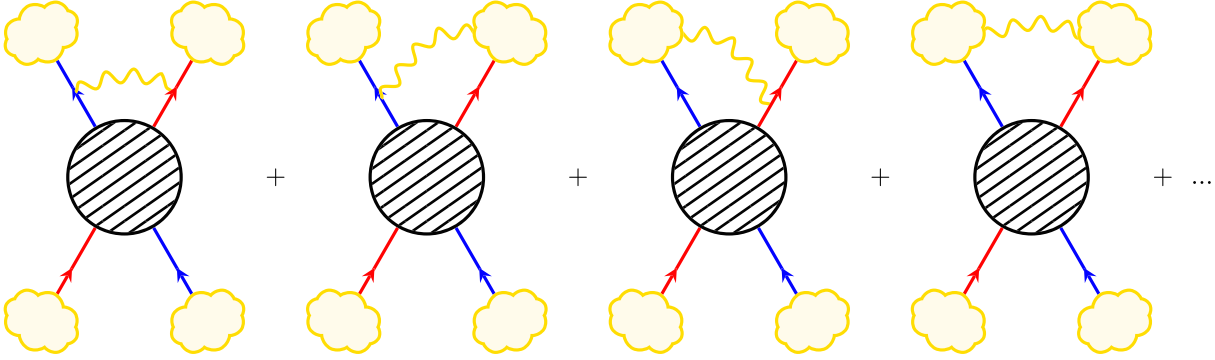


FIG. 1. IR divergences arise from soft photon exchange between pairs of external charges. When the charges are dressed with appropriately correlated clouds of soft photons, these divergences are pairwise cancelled by exchanges involving the soft clouds.

$$\langle \text{out}; 0 |_{\text{dressed}} \equiv \langle \text{out}; 0 | e^{-iR}. \quad (3.14)$$

The dressed states accompany the charges with nonzero eigenvalues for the soft photon operator,

$$N_z^- |\text{in}; 0\rangle_{\text{dressed}} = - \sum_{k \in \text{in}} \frac{Q_k e^2}{4\pi(z - z_k)} |\text{in}; 0\rangle_{\text{dressed}},$$

$$\langle \text{out}; 0 |_{\text{dressed}} N_z^+ = - \langle \text{out}; 0 |_{\text{dressed}} \sum_{k \in \text{out}} \frac{Q_k e^2}{4\pi(z - z_k)}. \quad (3.15)$$

In particular, the eigenvalues automatically obey the selection rule (2.14)

$$N_z^+ - N_z^- = \Omega_z^{\text{soft}}, \quad (3.16)$$

so that generically

$$\langle \text{out}; 0 |_{\text{dressed}} \mathcal{S} | \text{in}; 0 \rangle_{\text{dressed}} \neq 0. \quad (3.17)$$

In fact, the dressed amplitudes are free of IR divergences altogether. The basic mechanism is illustrated in Fig. 1. IR divergences arise from the exchange of soft photons between pairs of external legs. These exponentiate in such a way to cause ordinary Fock-basis amplitudes to vanish. However, when the charged particles are dressed by soft photon clouds, further divergences arise when a soft photon is exchanged between one external leg and the soft cloud surrounding the second external leg or between the pair of soft clouds. FK [20] showed that, by a judicious choice of such a soft cloud, one can arrange for the IR divergences to cancel and obtain an IR finite $\mathcal{S}$ -matrix. In the next section, we show that our dressed states differ from those of FK only by terms which are subleading in the IR and therefore effect the same IR cancellations.

IV. FK STATES

Faddeev and Kulish [20], building on Dirac and others [16–19], developed a scheme for dressing charged particles, which eliminates IR divergences. Their starting point was to argue that the Lehmann, Symanzik, Zimmermann

(LSZ) procedure for identifying asymptotic states is inapplicable in quantum electrodynamics; since the electromagnetic interaction has infinite range, there can be no isolated interaction region. They resolved this by observing that the action for charged particles contains a term,

$$\int J^\mu A_\mu d^4x, \quad (4.1)$$

where J^μ is the electromagnetic current. Since this current is conserved, the action is gauge invariant provided appropriate boundary conditions hold for the gauge transformations. This is a special case of Dirac's treatment, which leads to a collection of soft photons accompanying any charged particle.

If one studies the state for a single electron of three-momentum p^i , then in the eikonal approximation, the current is the classical current of a single charged particle located at $x^i = p^i t/m$. FK dressed a single particle charged state $|\vec{p}\rangle$ with the associated soft cloud

$$|\vec{p}\rangle_{FK} = \exp \left[- \frac{eQ_0}{(2\pi)^3} \int \frac{d^3q}{2q_0} (f^\mu a_\mu^\dagger(\vec{q}) - f^{*\mu} a_\mu(\vec{q})) \right] |\vec{p}\rangle,$$

$$f^\mu = \left[\frac{p^\mu}{p \cdot q} - c^\mu \right] e^{i \frac{p \cdot k}{p_0 t}}, \quad (4.2)$$

where c^μ satisfies $c \cdot q = 1, c^2 = 0$. In fact, they demonstrate that, in order to cancel infrared divergences, it is sufficient to choose an arbitrary dressing,

$$f^\mu = \left[\frac{p^\mu}{p \cdot q} - c^\mu \right] \psi(p, q), \quad (4.3)$$

with the condition that $\psi(p, q) = 1$ in a neighborhood of $q = 0$. For a multiparticle state with zero net total charge and minimal dressing $\psi = 1$, the c^μ terms cancel out of the dressing function, and we can deal solely with the dressing factor

$$f^\mu = \frac{p^\mu}{p \cdot q}. \quad (4.4)$$

We wish to rewrite (4.2) in terms of asymptotic fields at $\mathcal{I}$. The plane-wave expansion of the radiative modes of the electromagnetic potential is given by

$$A_\nu(x) = e \sum_{\alpha=\pm} \int \frac{d^3q}{(2\pi)^3} \frac{1}{2\omega} [\epsilon_\nu^{*\alpha}(\vec{q}) a_\alpha^{\text{in}}(\vec{q}) e^{iq \cdot x} + \epsilon_\nu^\alpha(\vec{q}) a_\alpha^{\text{in}}(\vec{q})^\dagger e^{-iq \cdot x}], \quad (4.5)$$

where $q^2 = 0$, the two polarization vectors satisfy a normalization condition $\epsilon_\alpha^\nu \epsilon_{\beta\nu}^* = \delta_{\alpha\beta}$, and

$$[a_\alpha^{\text{in}}(\vec{q}), a_\beta^{\text{in}}(\vec{q}')^\dagger] = \delta_{\alpha\beta} (2\pi)^3 2\omega \delta^{(3)}(\vec{q} - \vec{q}'). \quad (4.6)$$

Null momenta can be characterized by a point on the asymptotic S^2 and an energy ω ,

$$q^\mu = \frac{\omega}{1 + z\bar{z}} (1 + z\bar{z}, z + \bar{z}, -i(z - \bar{z}), 1 - z\bar{z}) \\ = (\omega, q^1, q^2, q^3). \quad (4.7)$$

Now, use the expansion for the spatial part of the plane-wave wave functions,

$$e^{i\vec{q} \cdot \vec{x}} = \sum_l i^l (2l+1) j_l(qr) P_l(\cos \gamma), \quad (4.8)$$

where j_l are the spherical Bessel functions, γ is the angle between $\vec{q}$ and $\vec{x}$, and $q = |\vec{q}|$. The asymptotic form of $j_l(qr)$ for $qr \gg 1$ is

$$j_l(qr) \sim \frac{1}{qr} \sin\left(qr - \frac{1}{2}\pi l\right), \quad (4.9)$$

which yields an approximation that localizes the momenta of the gauge field in the optical direction [6],

$$A_z^{(0)}(v, z, \bar{z}) = \lim_{r \rightarrow \infty} A_z(v, r, z, \bar{z}) \\ = -\frac{i}{8\pi^2} \frac{\sqrt{2}e}{1 + z\bar{z}} \int_0^\infty d\omega [a_+^{\text{in}}(\omega \hat{x}) e^{-i\omega v} - a_-^{\text{in}}(\omega \hat{x})^\dagger e^{i\omega v}], \quad \hat{x} = \hat{x}(z, \bar{z}). \quad (4.10)$$

One then finds

$$|\vec{p}\rangle_{FK} = \exp\left[\frac{i}{2\pi} \int dv d^2w \gamma_{w\bar{w}} d^2z \gamma_{z\bar{z}} j_v(v, z, \bar{z}) \right. \\ \left. \times G(z, w) D \cdot A(0, w, \bar{w})\right] |\vec{p}\rangle. \quad (4.11)$$

This is almost the same as our dressed state (3.10), except that the soft cloud in an FK state is always at $v = 0$, whereas in (3.10), it appears at the same advanced

time v_0 as the charged particle.¹¹ This difference is subleading in ω , since leading-order quantities at $\omega = 0$ are insensitive to null separations. In fact, with the natural choice

$$\psi = e^{i\omega v_0}, \quad (4.12)$$

which clearly satisfies $\psi = 1$ near $\omega = 0$, one obtains precisely the dressing in Sec. III. Since the two dressings differ only by terms higher order in ω , both implement the all-orders cancellation of IR divergences established by FK.

V. MASSIVE PARTICLES

In this section, the discussion is generalized to massive particles. The soft factor Ω_z^{soft} , which determines the change in the vacuum state, is given for massless particles in (2.8) in terms of the points z_k at which they exit or enter the celestial sphere. Such a formula cannot exist for massive particles in eigenstates with momentum p_k^μ as they never reach null infinity. Instead, a massive particle is characterized by a point on the unit three-dimensional hyperboloid H_3 , which may be parametrized by

$$\hat{p}^\mu = \frac{p^\mu}{m}, \quad \hat{p}^2 = -1. \quad (5.1)$$

As described in Refs. [4,28], the contribution to Ω_z^{soft} from such a particle is proportional to

$$G_z(\hat{p}_k) = \int d^2w G(w, \bar{w}; \hat{p}_k) \frac{1}{w - z}, \quad (5.2)$$

where G here is the bulk-to-boundary propagator on H_3 obeying $\square G = 0$. If we infinitely boost $\hat{p}$, G reduces to a boundary delta function. Hence, in analogy with (3.10), in order to prevent IR divergences from setting the amplitudes to zero, we should dress such massive particle states as

$$|\hat{p}_1, \dots; 0\rangle_{\text{dressed}} \equiv e^{iR_m} |\hat{p}_1, \dots; 0\rangle, \\ R_m \equiv \sum_{k \in \text{in}} \frac{Q_k}{2\pi} \int d^2z (G_z(\hat{p}_k) A_{\bar{z}}(0, z, \bar{z}) + \text{H.c.}), \quad (5.3)$$

where, to avoid a separate discussion of the zero mode, we restrict to the special case $\sum_{k \in \text{in}} Q_k = 0$ with zero net charge. It is straightforward to show that this is precisely the FK state given by (4.4) and therefore has IR finite scattering amplitudes.

Unlike the massless case studied above, this construction gives Coulomb fields. One finds

¹¹The FK states can still satisfy the constraints, at the price of Coulomb fields extending from $v = 0$ to the locations of the charged particles.

$$[D^z F_{vz} + D^{\bar{z}} F_{v\bar{z}}, iR_m] = -\delta(v) \gamma^{z\bar{z}} \sum_{k \in \text{in}} Q_k e^2 G(z, \bar{z}, \hat{p}_k). \quad (5.4)$$

This is a radiative shock wave coming out at $v = 0$. As there are no charged particles incoming at $v = 0$, the constraints then imply that $F_{rv}(z, \bar{z})$ must shift by $-\gamma^{z\bar{z}} \sum_{k \in \text{in}} Q_k e^2 G(z, \bar{z}, \hat{p}_k)$ at $v = 0$. This is precisely the (negative of the) asymptotic incoming Coulomb field in the absence of any radiation, associated with a collection of incoming massive point particles with momentum $m\hat{p}_k$. The constant part of F_{rv} is fixed by demanding that near $\mathcal{I}^-$ it equals the Coulomb field sourced by the massive charges entering through past timelike infinity i^- . This gives

$$F_{rv} = \theta(-v) \gamma^{z\bar{z}} \sum_{k \in \text{in}} Q_k e^2 G(z, \bar{z}, \hat{p}_k). \quad (5.5)$$

The effect of the radiative shock wave is to set to zero the Coulomb fields after $v = 0$ (note we are considering zero net charge). Similarly, fixing the integration function for F_{ru} with a boundary condition at $\mathcal{I}^+$, the corresponding out state has no Coulomb fields before $u = 0$. This is illustrated in Fig. 2. Since we then have

$$F_{ru}|_{\mathcal{I}^+} = 0 = F_{rv}|_{\mathcal{I}^+}, \quad (5.6)$$

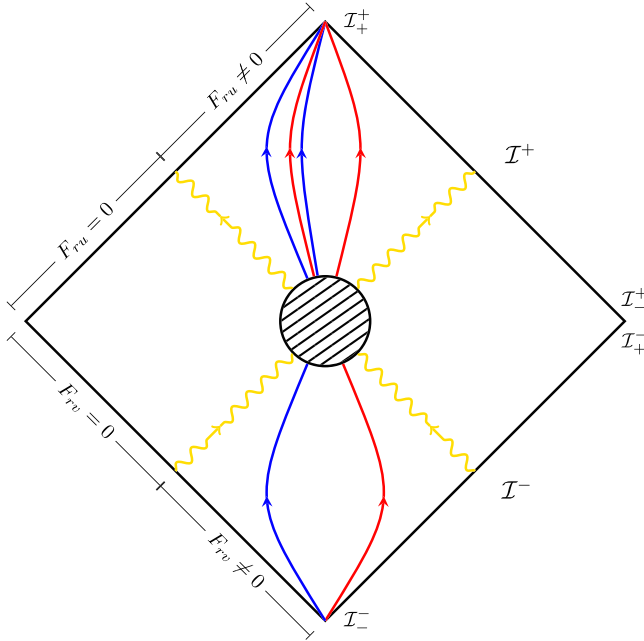


FIG. 2. Dressed massive particles of zero net charge come in from i^- , scatter, and go out to i^+ . The Faddeev-Kulish dressing introduces radiative shock waves at $v = 0$ and $u = 0$, which cancel the asymptotic Coulomb fields of the particles for $v > 0$ and $u < 0$ respectively. For neutral scattering states, the Coulomb field will vanish near spatial infinity, while for charged ones, it will be an angle-independent constant.

this construction guarantees that the matching condition (2.17) is trivially satisfied and the amplitude need not vanish. Had we not restricted to the zero-charge sector, the boundary field strengths (5.6) would be angle-independent constants.

Note that for massive charged particles in plane-wave states, the soft photon cloud can never be “on top of” the particle. Massive particles go to timelike infinity, while radiative photons always disperse to null infinity. The simplest FK states have them coming out at $u = 0$, but exactly when they come out or the fact that they come out before the charges themselves is unimportant since only the leading IR behavior of the cloud is relevant to the cancellation of divergences.

VI. CHARGED STATES

In this section, we consider a more general class of physical states in which Coulomb fields persist to $\mathcal{I}^+$ and $\mathcal{I}^-$, the generic charges are nonzero, and the matching conditions are nontrivially satisfied.

The condition (5.6), which is obeyed by all FK states (with zero net charge), implies that the in and out charges $Q_e^\pm$ vanish. Explicitly, one finds

$$Q^-|_{\epsilon=\frac{1}{z-w}}|\text{in}; 0\rangle_{\text{dressed}} = \left(\frac{4\pi}{e^2} N_z^- - \sum_{k \in \text{in}} Q_k G_z(\hat{p}_k) \right) |\text{in}; 0\rangle_{\text{dressed}} = 0, \quad (6.1)$$

$$\langle \text{out}; 0 |_{\text{dressed}} Q^+|_{\epsilon=\frac{1}{z-w}} = \langle \text{out}; 0 |_{\text{dressed}} \times \left(\frac{4\pi}{e^2} N_z^+ - \sum_{k \in \text{out}} Q_k G_z(\hat{p}_k) \right) = 0. \quad (6.2)$$

That is, the incoming (outgoing) soft photon cloud shields all of the nonzero mode charges of the incoming (outgoing) charged particles. Since the charges all commute with $a_{\pm}^{\text{in}\dagger}(\vec{k})$, adding radiative photons will not change the charges.¹²

In contrast to the situation for FK states, nonzero mode $Q_e^\pm$ charges are generically all nonvanishing in the real world. Consider for example e^+e^- scattering in the center-of-mass frame with incoming velocities $\pm\vec{v}$, unaccompanied by incoming radiation. The coefficient of the $\frac{1}{r^2}$ radial electric field is given by the Liénard-Wiechert formula

$$F_{rv} = \frac{e^2(1-\vec{v}^2)}{4\pi(1+\hat{x}\cdot\vec{v})^2} - \frac{e^2(1-\vec{v}^2)}{4\pi(1-\hat{x}\cdot\vec{v})^2} = -\frac{e^2}{\pi} \frac{\hat{x}\cdot\vec{v}(1-\vec{v}^2)}{(1-(\hat{x}\cdot\vec{v})^2)^2}. \quad (6.3)$$

¹²Assuming the frequency distribution does not have poles or other singularities for $\omega \rightarrow 0$. If it does have such poles, new IR singularities may appear, and the state would not be among those shown by FK to be IR finite.

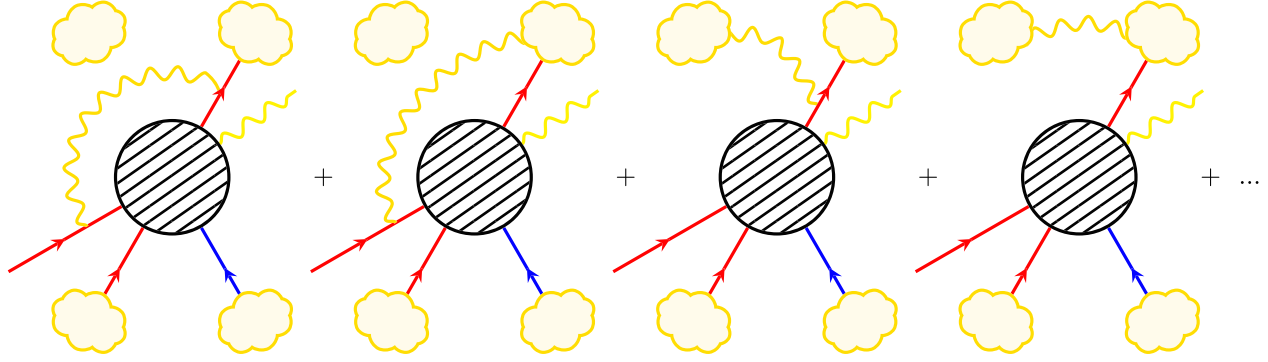


FIG. 3. In this figure, the outgoing positron in Fig. 1 is crossed to an incoming electron, but its associated soft photon cloud remains as part of the out state. The leading IR divergences from the depicted soft photon exchange still cancel, even though the in and out states carry nontrivial $Q_\epsilon^\pm$ charges and are no longer Faddeev-Kulish states.

The charges constructed from this are nonzero,

$$\begin{aligned} Q^-|_{\epsilon=\frac{1}{z-w}} &= - \int_{\mathcal{I}_+^-} d^2w \gamma_{w\bar{w}} \frac{1}{z-w} \frac{\hat{x} \cdot \vec{v}(1-\vec{v}^2)}{\pi(1-(\hat{x} \cdot \vec{v})^2)^2} \\ &= Q^+|_{\epsilon=\frac{1}{z-w}} = G_z(\hat{p}_-) - G_z(\hat{p}_+), \end{aligned} \quad (6.4)$$

where $\hat{p}_-$ ($\hat{p}_+$) is the momentum of the electron (positron). Nonzero $Q_\epsilon^\pm$ charges can be generically sourced by radiative Maxwell fields and arise even in the absence of charged particles. Source-free initial data for the Maxwell equation are given by specifying an arbitrary function $F_{vz}(v, z, \bar{z})$ on $\mathcal{I}^-$. Assuming that F_{rv} vanishes in the far past, one has

$$F_{rv}(z, \bar{z})|_{\mathcal{I}_+^-} = \int_{-\infty}^{\infty} dv (D^z F_{vz}(v, z, \bar{z}) + D^{\bar{z}} F_{v\bar{z}}(v, z, \bar{z})). \quad (6.5)$$

Demanding that the right-hand side vanishes is a nonlocal constraint on the incoming initial data. Such nonlocal constraints are indeed imposed on FK states, which are dressed charged particles plus radiative modes. As mentioned above, the frequency-space coefficients of the field operators A_z are (except for the charge dressings) presumed to be finite for $\omega \rightarrow 0$, which precisely imposes the non-local constraint on the field strength that the integral (6.5) vanishes.

Quantum states which describe these physical situations with nonzero $Q_\epsilon^\pm$ charges certainly exist, even if they are not FK states.¹³ It is natural to ask if such states can ever have IR finite scattering amplitudes. Given our earlier argument that the true role of IR divergences is simply to

enforce conservation of all the charges, one might expect it to be possible. We now propose that this is indeed the case.

The basic idea is that the soft photon clouds and charged particles can be separated without affecting the IR cancellation mechanism of FK, even if we move a particle from incoming to outgoing, leaving its cloud intact. However, moving a charged particle from incoming to outgoing will in general take a zero-charge FK state to one with all charges excited.

Let us consider Bhabha scattering as a specific example,

$$e^+ e^- \rightarrow e^+ e^-, \quad (6.6)$$

where the incoming and outgoing charges are all given FK dressings. Then, there will be an incoming wave of photons shielding the incoming charges at $v = 0$ and an outgoing one at $u = 0$ shielding the outgoing charges. The long range fields will be angle independent, in contrast to (6.3). The scattering is IR finite, with IR divergences from soft photon exchanges between pairs of external charges canceled by divergences from soft photon exchanges between external particles and radiative clouds and pairs of radiative clouds. This was depicted in Fig. 1.

Now, let us move the outgoing positron to an ingoing electron with the same momentum and add a radiative photon to the out state to conserve energy and momentum:

$$e^+ e^- e^- \rightarrow e^- + \gamma. \quad (6.7)$$

This has no effect on the soft factor, since the motion from out to in and the change in the sign of the charge each contributes a factor of -1 . The same $(-1)^2 = 1$ applies to the leading IR divergence of an attached soft photon and ensures that these soft exchanges will continue to cancel. See Fig. 3.

Since we have also done nothing to $N_z^\pm$, the conservation law (3.16) remains satisfied. However, we have changed the charges. While they were previously zero, the contribution from the FK shield of the outgoing positron is no

¹³We note that all the nonzero mode charges can be shielded, classically or quantum mechanically, by a correlated cloud of very soft radiation with arbitrarily small energy at arbitrarily large radius. In this sense, any state can be approximated by an FK state.

longer cancelled, while the new incoming electron does not have an FK shield. These give equal contributions to the incoming and outgoing charges,

$$Q^\pm(z) = G_z(\hat{p}). \quad (6.8)$$

Ultimately, one might hope to use crossing symmetry to prove IR finiteness in this context. FK IR cancellations occur order by order in Feynman diagram perturbation theory, while crossing symmetry also holds in perturbation theory. The action of crossing a single out particle to an in one

produces general $Q_e^\pm$ charges while changing FK to some more general class of states. We conjecture that scattering amplitudes among these more general charged states are IR finite.

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- [1] A. Strominger, Asymptotic symmetries of Yang-Mills theory, *J. High Energy Phys.* **07** (2014) 151.
 - [2] T. He, P. Mitra, A. P. Porfyriadis, and A. Strominger, New symmetries of massless QED, *J. High Energy Phys.* **10** (2014) 112.
 - [3] T. He, P. Mitra, and A. Strominger, 2D Kac-Moody symmetry of 4D Yang-Mills theory, *J. High Energy Phys.* **10** (2016) 137.
 - [4] M. Campiglia and A. Laddha, Asymptotic symmetries of QED and Weinberg's soft photon theorem, *J. High Energy Phys.* **07** (2015) 115.
 - [5] D. Kapec, M. Pate, and A. Strominger, New Symmetries of QED, [arXiv:1506.02906](#).
 - [6] A. Strominger, Lectures on the Infrared Structure of Gravity and Gauge Theory, [arXiv:1703.05448](#).
 - [7] N. F. Mott, On the influence of radiative forces on the scattering of electrons, *Proc. Cambridge Philos. Soc.* **27**, 255 (1931).
 - [8] F. Bloch and A. Nordsieck, Note on the radiation field of the electron, *Phys. Rev.* **52**, 54 (1937).
 - [9] D. R. Yennie, S. C. Frautschi, and H. Suura, The infrared divergence phenomena and high-energy processes, *Ann. Phys. (N.Y.)* **13**, 379 (1961).
 - [10] T. Kinoshita, Mass singularities of Feynman amplitudes, *J. Math. Phys. (N.Y.)* **3**, 650 (1962).
 - [11] T. D. Lee and M. Nauenberg, Degenerate systems and mass singularities, *Phys. Rev.* **133**, B1549 (1964).
 - [12] S. Weinberg, Infrared photons and gravitons, *Phys. Rev.* **140**, B516 (1965).
 - [13] S. Pasterski, Asymptotic Symmetries and Electromagnetic Memory, [arXiv:1505.00716](#).
 - [14] L. Susskind, Electromagnetic Memory, [arXiv:1507.02584](#).
 - [15] L. Bieri and D. Garfinkle, An electromagnetic analogue of gravitational wave memory, *Classical Quantum Gravity* **30**, 195009 (2013).
 - [16] V. Chung, Infrared divergence in quantum electrodynamics, *Phys. Rev.* **140**, B1110 (1965).
 - [17] T. W. B. Kibble, Coherent soft-photon states and infrared divergences. II. Mass-shell singularities of Green's functions, *Phys. Rev.* **173**, 1527 (1968).
 - [18] T. W. B. Kibble, Coherent soft-photon states and infrared divergences. III. Asymptotic states and reduction formulas, *Phys. Rev.* **174**, 1882 (1968).
 - [19] T. W. B. Kibble, Coherent soft-photon states and infrared divergences. IV. The scattering operator, *Phys. Rev.* **175**, 1624 (1968).
 - [20] P. P. Kulish and L. D. Faddeev, Asymptotic conditions and infrared divergences in quantum electrodynamics, *Teor. Mat. Fiz.* **4**, 153 (1970) [*Theor. Math. Phys.* **4**, 745 (1970)].
 - [21] B. Gabai and A. Sever, Large gauge symmetries and asymptotic states in QED, *J. High Energy Phys.* **12** (2016) 095.
 - [22] M. Mirbabayi and M. Porrati, Dressed Hard States and Black Hole Soft Hair, *Phys. Rev. Lett.* **117**, 211301 (2016).
 - [23] C. Gomez and R. Letschka, Memory and the Infrared, [arXiv:1704.03395](#).
 - [24] M. Panchenko, The infrared triangle in the context of IR safe S matrices, [arXiv:1704.03739](#).
 - [25] A. Mohd, A note on asymptotic symmetries and soft-photon theorem, *J. High Energy Phys.* **02** (2015) 060.
 - [26] S. Pasterski, S. H. Shao, and A. Strominger, Flat Space Amplitudes and Conformal Symmetry of the Celestial Sphere, [arXiv:1701.00049](#).
 - [27] P. A. M. Dirac, Gauge invariant formulation of quantum electrodynamics, *Can. J. Phys.* **33**, 650 (1955).
 - [28] M. Campiglia, Null to time-like infinity Green's functions for asymptotic symmetries in Minkowski spacetime, *J. High Energy Phys.* **11** (2015) 160.