# Algebraic bounds on the Rayleigh-Bénard attractor 

Yu Cao ${ }^{1}$, Michael S Jolly ${ }^{1}{ }^{\text {© }}$, Edriss S Titi ${ }^{2,3,4, *}$ (1) and Jared P Whitehead ${ }^{5}$<br>${ }^{1}$ Department of Mathematics, Indiana University Bloomington, IN 47405, United States of America<br>${ }^{2}$ Department of Mathematics, Texas A \& M University, 3368 TAMU, College<br>Station, TX 77843-3368, United States of America<br>${ }^{3}$ Department of Computer Science and Applied Mathematics, Weizmann Institute of Science, Rehovot 76100, Israel<br>${ }^{4}$ Department of Applied Mathematics and Theoretical Physics, University of Cambridge, Cambridge CB3 0WA, United Kingdom<br>${ }^{5}$ Department of Mathematics Brigham Young University Provo, UT 84602, United States of America<br>E-mail: cao20@iu.edu, msjolly@indiana.edu, titi@math.tamu.edu, Edriss.Titi@damtp.cam.ac.uk and whitehead@mathematics.byu.edu

Received 3 May 2019, revised 10 April 2020
Accepted for publication 24 August 2020
Published 22 January 2021


## Abstract

The Rayleigh-Bénard system with stress-free boundary conditions is shown to have a global attractor in each affine space where velocity has fixed spatial average. The physical problem is shown to be equivalent to one with periodic boundary conditions and certain symmetries. This enables a Gronwall estimate on enstrophy. That estimate is then used to bound the $L^{2}$ norm of the temperature gradient on the global attractor, which, in turn, is used to find a bounding region for the attractor in the enstrophy-palinstrophy plane. All final bounds are algebraic in the viscosity and thermal diffusivity, a significant improvement over previously established estimates. The sharpness of the bounds are tested with numerical simulations.

Keywords: Rayleigh-Bénard convection, global attractor, synchronization Mathematics Subject Classification numbers: 35Q35, 76E06, 76F35, 34D06.
(Some figures may appear in colour only in the online journal)
*Author to whom any correspondence should be addressed.
Original content from this work may be used under the terms of the Creative Commons Attribution 3.0 licence. Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI.

## 1. Introduction

The long-time behavior of the Rayleigh-Bénard (RB) problem was analyzed in [11, 18] for several types of boundary conditions. In that work the authors derived explicit estimates for enstrophy and the ( $L^{2}$-norm of the) temperature gradient on the global attractor for the case of no-slip boundary conditions in space dimension two. They also outlined the functional setting for the case of stress-free velocity boundary conditions (see (2.2a) and (2.2b)), and mentioned that corresponding estimates can be carried out in a similar fashion. In this paper we revisit the 2D, stress-free boundary conditions case, and as in the case of rigorous bounds on the time averaged heat transport [19], we find estimates on the global attractor which are dramatically reduced from those in the no-slip boundary conditions case. We also derive estimates for the palinstrophy and $H^{2}$-norm of the temperature.

One marked difference between no-slip and stress-free boundary conditions is that in the latter case, the system is not dissipative for general initial velocity data. This is due to the existence of steady states with arbitrarily large $L^{2}$-norms, namely velocity of the form $u(x, t)=(c, 0)$, (where $c$ is a constant) with zero temperature $\theta(x, t)=0$ such as the shear-dominated flow investigated in [13]. Since, however, the spatial average is conserved for these flows, the system is dissipative within each invariant affine space of fixed horizontal velocity average. This wrinkle does not influence the estimates on the temperature or higher Sobolev norm estimates on the velocity.

The a priori estimates are carried out in section 4 . The key to finding sharper bounds in the stress-free case is to extend the physical domain, as done in [9], to one that is fully periodic and twice the height of the original. This makes the trilinear term vanish from the enstrophy balance, giving an easy bound that is $\mathcal{O}\left(\nu^{-2}\right)$ in terms of the kinematic viscosity $\nu$. Though the trilinear term persists when estimating the temperature gradient, we are able to avoid the exponential bound that resulted from using a uniform Gronwall lemma in [11], by using the algebraic bound on the enstrophy. We find that on the global attractor the ( $L^{2}$-norm of the) temperature gradient satisfies a bound that is $\mathcal{O}\left(\mathrm{Ra}^{2}\right)$, for $\operatorname{Pr} \sim 1$, where Ra is the Rayleigh number and $\operatorname{Pr}$ is the Prandtl number.

We then follow the approach in [6] for the Navier-Stokes equations (NSE) to obtain an estimate for the palinstrophy, with the temperature playing the role of the body force in the NSE. This leads to curves which bound the attractor in the enstrophy-palinstrophy plane, with an overall bound on palinstrophy that is $\mathcal{O}\left(\mathrm{Ra}^{3}\right)$ for $\operatorname{Pr} \sim 1$. Using this palinstrophy bound, we then follow a similar procedure to find a bounding region for temperature $\theta$ in the $\|\nabla \theta\|_{L^{2}}^{2}-\|\Delta \theta\|_{L^{2}}^{2}$ plane.

In section 5 we recall from [9] how all of these bounds impact the practicality of data assimilation by nudging with just the horizontal component of velocity of the stress-free RB system. The sharpness of our rigorous bounds are tested with numerical simulations over a range of Rayleigh numbers in section 6 . Simulations are also presented there to demonstrate that the nudging algorithm works for data with much lower resolution than the analysis requires. This is actually what suggested we might improve on the exponential bounds in [11, 18]. All the bounds here on the attractor are algebraic in the physical parameters.

## 2. Preliminaries

The RB problem on the domain $\Omega_{0}=(0, L) \times(0,1)$ can be written in dimensionless form as (see, e.g., [11])

$$
\begin{align*}
& \frac{\partial u}{\partial t}-\nu \Delta u+(u \cdot \nabla) u+\nabla p=\theta e_{2}  \tag{2.1a}\\
& \frac{\partial \theta}{\partial t}-\kappa \Delta \theta+(u \cdot \nabla) \theta=u \cdot e_{2}  \tag{2.1b}\\
& \nabla \cdot u=0  \tag{2.1c}\\
& u(0 ; x)=u_{0}(x), \quad \theta(0 ; x)=\theta_{0}(x) \tag{2.1d}
\end{align*}
$$

where $e_{2}=(0,1)$ and $\kappa$ is the thermal diffusivity. In this paper, we consider the following set of boundary conditions that are stress-free on the velocity:

$$
\begin{array}{rlll}
\text { inthe } x_{2} \text {-direction: } u_{2}, \theta=0 & \text { at } x_{2}=0 & \text { and } & x_{2}=1 \\
\qquad \frac{\partial u_{1}}{\partial x_{2}}=0 & \text { at } x_{2}=0 & \text { and } & x_{2}=1 \tag{2.2b}
\end{array}
$$

inthe $x_{1}$-direction: $u, \theta, p$ are of period $L$,
where the indices 1 and 2 refer the horizontal and vertical components, respectively.
Following [9], in the rest of this paper, we consider the equivalent formulation of problem (2.1) subject to the fully periodic boundary conditions on the extended domain $\Omega=(0, L) \times$ $(-1,1)$ with the following special spatial symmetries:

$$
\begin{gathered}
u_{1}\left(x_{1}, x_{2}\right)=u_{1}\left(x_{1},-x_{2}\right), \quad u_{2}\left(x_{1}, x_{2}\right)=-u_{2}\left(x_{1},-x_{2}\right), \\
p\left(x_{1}, x_{2}\right)=p\left(x_{1},-x_{2}\right), \quad \theta\left(x_{1}, x_{2}\right)=-\theta\left(x_{1},-x_{2}\right),
\end{gathered}
$$

for $\left(x_{1}, x_{2}\right) \in \Omega$. As a result of this symmetry, we observe that smooth enough functions satisfy

$$
\begin{equation*}
u_{2}, \theta, \frac{\partial u_{1}}{\partial x_{2}}=0, \quad \text { for } x_{2}=-1,0,1 \tag{2.3}
\end{equation*}
$$

### 2.1. Function spaces

We will use the same notation indiscriminately for both scalar and vector Lebesgue and Sobolev spaces, which should not be a source of confusion. We denote

$$
\begin{aligned}
(u, v) & :=\int_{\Omega} u \cdot v, \quad \text { for } u, v \in L^{2}(\Omega) \\
((u, v)) & :=\sum_{i, j=1}^{2} \int_{\Omega} \frac{\partial u_{i}}{\partial x_{j}} \frac{\partial v_{i}}{\partial x_{j}}, \quad \text { for } u, v \in H^{1}(\Omega),
\end{aligned}
$$

and

$$
|u|:=(u, u)^{1 / 2}, \quad\|u\|:=((u, u))^{1 / 2} .
$$

Note that $\|\cdot\|$ is not a norm, but will form part of one in (2.4). We define function spaces corresponding to the relevant physical boundary conditions as in [9], where
$\mathcal{F}_{1}$ is the set of trigonometric polynomials in $\left(x_{1}, x_{2}\right)$, with period $L$ in the $x_{1}$-variable, that are even, with period 2 in the $x_{2}$-variable,
and
$\mathcal{F}_{2}$ is the set of trigonometric polynomials in $\left(x_{1}, x_{2}\right)$, with period $L$ in the $x_{1}$-variable, that are odd, with period 2 in the $x_{2}$-variable.
The space of smooth vector-valued functions which incorporates the divergence-free condition shall be denoted by

$$
\mathcal{V}:=\left\{u \in \mathcal{F}_{1} \times \mathcal{F}_{2}: \nabla \cdot u=0\right\} .
$$

We denote the closures of $\mathcal{V}$ and $\mathcal{F}_{2}$ in $L^{2}(\Omega)$ by $H_{0}$ and $H_{1}$, respectively, which are endowed with the usual inner products

$$
(u, v)_{H_{0}}:=(u, v), \quad(\psi, \phi)_{H_{1}}:=(\psi, \phi)
$$

and the associated norms

$$
\|u\|_{H_{0}}:=(u, u)^{1 / 2}, \quad\|\psi\|_{H_{1}}:=(\psi, \psi)^{1 / 2}
$$

We define for $k=1,2$

$$
H_{\mathrm{per}}^{k}(\Omega)=\left\{\phi \in H^{k} \mid \phi \text { has period } L \text { in } x_{1}, \text { period } 2 \text { in } x_{2}\right\}
$$

Finally, we denote the closures of $\mathcal{V}$ and $\mathcal{F}_{2}$ in $H_{\text {per }}^{1}(\Omega)$ by $V_{0}$ and $V_{1}$ respectively, endowed with the inner products

$$
((u, v))_{V_{0}}:=\frac{1}{|\Omega|}(u, v)+((u, v)), \quad((\psi, \phi))_{V_{1}}:=((\psi, \phi)),
$$

and associated norms

$$
\begin{equation*}
\|u\|_{V_{0}}:=\left(\frac{1}{|\Omega|}|u|^{2}+\|u\|^{2}\right)^{1 / 2}, \quad\|\phi\|_{V_{1}}:=\|\phi\| \tag{2.4}
\end{equation*}
$$

where $|\Omega|=2 L$ is the volume of $\Omega$.

### 2.2. The linear operators $A_{i}$

Let $D\left(A_{0}\right)=V_{0} \cap H_{\text {per }}^{2}(\Omega)$ and $D\left(A_{1}\right)=V_{1} \cap H_{\text {per }}^{2}(\Omega)$. Let $A_{i}: D\left(A_{i}\right) \rightarrow H_{i}(i=0,1)$ be the unbounded linear operators defined by

$$
\left(A_{i} \phi, \psi\right)_{H_{i}}=((\phi, \psi)), \quad \phi, \psi \in D\left(A_{i}\right) .
$$

Due to periodic boundary conditions, we have $A_{i}=-\Delta$. The operator $A_{0}$ is a nonnegative operator and possesses a sequence of eigenvalues with

$$
0=\lambda_{0,1} \leqslant \lambda_{0,2} \leqslant \cdots \leqslant \lambda_{0, m} \leqslant \cdots,
$$

associated with an orthonormal basis $\left\{w_{0, m}\right\}_{m \in \mathbb{N}}$ of $H_{0}$. The operator $A_{1}$ is a positive self-adjoint operator and possesses a sequence of eigenvalues with

$$
0<\lambda_{1,1} \leqslant \lambda_{1,2} \leqslant \cdots \leqslant \lambda_{1, m} \leqslant \cdots,
$$

associated with an orthonormal basis $\left\{w_{1, m}\right\}_{m \in \mathbb{N}}$ of $H_{1}$. Observe that we have the Poincaré inequality for temperature:

$$
|\theta|^{2} \leqslant \lambda_{1}^{-1}\|\theta\|^{2}, \quad \forall \theta \in V_{1}
$$

$$
|\theta|^{2} \leqslant \lambda_{1}^{-1}\left\|A_{1} \theta\right\|^{2}, \quad \forall \theta \in D\left(A_{1}\right),
$$

where $\lambda_{1}=\lambda_{1,1}=\pi^{2} \min \left(1 / 4, L^{-2}\right)$.

### 2.3. The bilinear maps $B_{i}$

Denote the dual space of $V_{i}$ by $V_{i}^{\prime}(i=0,1)$. Define the bounded trilinear map $b_{0}: V_{0} \times V_{0} \times$ $V_{0} \rightarrow \mathbb{R}$ by the continuous extension (see $[5,18]$ ) of

$$
b_{0}(u, v, w):=((u \cdot \nabla) v, w)=-((u \cdot \nabla) w, v) \quad u, v, w \in \mathcal{V}
$$

The boundedness of the trilinear map $b_{0}$ implies that there is a unique bounded bilinear map $B_{0}: V_{0} \times V_{0} \rightarrow V_{0}^{\prime}$ such that

$$
\left\langle B_{0}(u, v), w\right\rangle_{V_{0}^{\prime}}:=b_{0}(u, v, w)
$$

Define the scalar analogue trilinear bounded map $b_{1}: V_{0} \times V_{1} \times V_{1} \rightarrow \mathbb{R}$ by the continuous extension of

$$
b_{1}(u, \theta, \phi):=((u \cdot \nabla) \theta, \phi)=-((u \cdot \nabla) \phi, \theta), \quad u \in \mathcal{V}, \quad \theta, \phi \in \mathcal{F}_{2} .
$$

As before the boundedness of the trilinear map $b_{1}$ in turn implies the existence of a unique bounded bilinear $B_{1}: V_{0} \times V_{1} \rightarrow V_{1}^{\prime}$ such that

$$
\left\langle B_{1}(u, \theta), \phi\right\rangle_{V_{1}^{\prime}}:=b_{1}(u, \theta, \phi)
$$

From the above, the trilinear maps $b_{i}$ (and the bilinear maps $B_{i}$ ), $i=0,1$, enjoy the following orthogonality property:

$$
\begin{equation*}
b_{0}(u, v, v)=0, \quad b_{1}(u, \theta, \theta)=0, \quad u, v \in V_{0}, \theta \in V_{1} . \tag{2.5}
\end{equation*}
$$

Furthermore, due to periodicity on $\Omega$, i.e., since $A_{0}=-\Delta$, we have

$$
\begin{equation*}
b_{0}\left(u, u, A_{0} u\right)=0, \quad \forall u \in D\left(A_{0}\right) \tag{2.6}
\end{equation*}
$$

as well as

$$
\begin{equation*}
b_{0}\left(v, v, A_{0} w\right)+b_{0}\left(v, w, A_{0} v\right)+b_{0}\left(w, v, A_{0} v\right)=0, \quad \forall v, w \in D\left(A_{0}\right), \tag{2.7}
\end{equation*}
$$

(see, e.g., [18] for (2.6), [10] for (2.7)).

### 2.4. Functional setting

Following [11], we have the functional form of the RB problem (2.1):

$$
\begin{align*}
& \frac{\mathrm{d} u}{\mathrm{~d} t}+\nu A_{0} u+B_{0}(u, u)=\mathbb{P}_{\sigma}\left(\theta e_{2}\right),  \tag{2.8a}\\
& \frac{\mathrm{d} \theta}{\mathrm{~d} t}+\kappa A_{1} \theta+B_{1}(u, \theta)=u \cdot e_{2}  \tag{2.8b}\\
& u(0 ; x)=u_{0}(x), \quad \theta(0 ; x)=\theta_{0}(x), \tag{2.8c}
\end{align*}
$$

where $\mathbb{P}_{\sigma}$ denotes the Leray projector.

## 3. Statement of result

Theorem 3.1. The RB problem (2.1) with stress-free boundary conditions (2.3) has a global attractor $\mathcal{A}_{\alpha}$ within the invariant affine space

$$
W_{\alpha}=\left\{(u, \theta) \in V_{0} \times V_{1}: \int_{\Omega} u_{1}(x, t) \mathrm{d} x=\alpha\right\} .
$$

The elements in $\mathcal{A}_{\alpha}$ satisfy

$$
\begin{align*}
& |u|^{2} \leqslant \frac{|\Omega|}{\nu^{2} \lambda_{1}^{2}}+\alpha^{2}|\Omega|  \tag{3.1}\\
& \|u\|^{2} \leqslant z_{\max }:=\frac{|\Omega|}{\nu^{2} \lambda_{1}}  \tag{3.2}\\
& \|\theta\|^{2} \lesssim \vartheta_{\max }:=|\Omega| z_{\max } R a P r+\left(\frac{|\Omega|}{\lambda_{1}} z_{\max } \operatorname{RaPr}\right)^{1 / 2},  \tag{3.3}\\
& \left|A_{0} u\right|^{2} \leqslant f\left(\|u\|^{2}\right) \lesssim q_{\max }:=\frac{z_{\max }^{2}}{\nu^{2}}+\frac{z_{\max }^{1 / 2}}{\nu} \vartheta_{\max }^{1 / 2}  \tag{3.4}\\
& \left|A_{1} \theta\right|^{2} \leqslant g\left(\|\theta\|^{2}\right) \lesssim \eta_{\max }:=\frac{z_{\max } \vartheta_{\max }}{\kappa^{2}}+\frac{q_{\max }^{2 / 3} \vartheta_{\max }}{\kappa^{4 / 3} \lambda_{1}^{1 / 3}+\frac{z_{\max }}{\kappa^{2} \lambda_{1}}} \tag{3.5}
\end{align*}
$$

where the functions $f, g$ are defined below in (4.28) and (4.35), respectively, $\operatorname{Pr}$ is the Prandtl number $\nu / \kappa, \operatorname{Ra}=1 /(\nu \kappa)$ is the Rayleigh number, and $Q_{1} \lesssim Q_{2}$ means $Q_{1} \leqslant c Q_{2}$ for a nondimensional universal constant $c$ that is independent of the physical parameters.

Regions that bound the global attractor in the enstrophy-palinstrophy and $\|\theta\|^{2}-\left|A_{1} \theta\right|^{2}$ planes are depicted in figures 1 and 2, below.

## 4. A priori estimates

Global existence and uniqueness follows by the standard Galerkin procedure based on the trigonometric basis functions in the definitions of $\mathcal{F}_{1}$ and $\mathcal{F}_{2}$. We thus proceed with a priori estimates.

## 4.1. $L^{2}$ bound on temperature

We have the following maximum principle from lemma 2.1 in [11]

$$
\begin{equation*}
|\theta(t)| \leqslant|\Omega|^{1 / 2}+\Theta_{0} \mathrm{e}^{-\kappa t} \tag{4.1}
\end{equation*}
$$

where $|\Omega|$ is the volume of $\Omega$,

$$
\Theta_{0}=\left|(\theta(0)-1)_{+}\right|+\left|(\theta(0)+1)_{-}\right|,
$$

and

$$
M_{+}=\max \{M, 0\}, \quad M_{-}=\max \{-M, 0\} \quad \forall M \in \mathbb{R}
$$



Figure 1. Qualitative sketches of the curves bounding $\mathcal{A}_{\alpha}$.

While the proof in [11] was done for no-slip boundary conditions, the only place the velocity $u$ enters is the orthogonality property $b_{1}(u, \theta, \theta)=0$. The proof carries over verbatim to the stress-free case by (2.5). Consequently, we have (4.1) for each strong solution $(u, \theta)$ of (2.8).

## 4.2. $L^{2}$ bounds on velocity

We denote the space average of the horizontal velocity over the extended domain by

$$
\alpha(t)=\frac{1}{|\Omega|} \int_{\Omega} u_{1}(x, t) \mathrm{d} x .
$$

From (2.1) and the periodic boundary conditions on $\Omega$, we find that the spatial average of the horizontal velocity is conserved, i.e., $\mathrm{d} \alpha / \mathrm{d} t=0$. It follows that $u_{\alpha}=u-\alpha e_{1}$ satisfies

$$
\frac{\mathrm{d} u_{\alpha}}{\mathrm{d} t}+\nu A_{0} u_{\alpha}+B_{0}\left(u_{\alpha}+\alpha e_{1}, u_{\alpha}\right)=\theta e_{2} .
$$

Since $u_{\alpha}$ has zero average, it satisfies the Poincaré inequality

$$
\begin{equation*}
\lambda_{1}\left|u_{\alpha}\right|^{2} \leqslant\left\|u_{\alpha}\right\|^{2} . \tag{4.2}
\end{equation*}
$$

Note that since $u_{2}$ has zero mean, it satisfies a Poincaré inequality

$$
\begin{equation*}
\lambda_{1}\left|u_{2}\right|^{2} \leqslant\left\|u_{2}\right\|^{2}, \tag{4.3}
\end{equation*}
$$



Figure 2. Bounding region in the $\|\theta\|^{2}-\left|A_{1} \theta\right|^{2}$ plane.
even though $u_{1}$ does not. Taking the scalar product with $u_{\alpha}$, and applying (2.5), the Cauchy-Schwarz and Young inequalities as well as (4.2), we get

$$
\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\left|u_{\alpha}\right|^{2}+\nu\left\|u_{\alpha}\right\|^{2} \leqslant \frac{1}{2 \nu \lambda_{1}}|\theta|^{2}+\frac{\nu \lambda_{1}}{2}\left|u_{\alpha}\right|^{2} \leqslant \frac{|\theta|^{2}}{2 \nu \lambda_{1}}+\frac{\nu}{2}\left\|u_{\alpha}\right\|^{2} .
$$

Applying (4.2) once again, together with (4.1) and Young's inequality, we have

$$
\frac{\mathrm{d}}{\mathrm{~d} t}\left|u_{\alpha}\right|^{2}+\nu \lambda_{1}\left|u_{\alpha}\right|^{2} \leqslant \frac{1}{\nu \lambda_{1}}\left(|\Omega|+\Theta_{0}^{2} \mathrm{e}^{-2 \kappa t}\right),
$$

so that

$$
\begin{equation*}
\left|u_{\alpha}(t)\right|^{2} \leqslant \mathrm{e}^{-\nu \lambda_{1} t}\left|u_{\alpha}(0)\right|^{2}+\frac{1}{\nu \lambda_{1}} \int_{0}^{t}\left(|\Omega|+\Theta_{0}^{2} \mathrm{e}^{-2 \kappa s}\right) \mathrm{e}^{\nu \lambda_{1}(s-t)} \mathrm{d} s, \tag{4.4}
\end{equation*}
$$

and thus,

$$
\begin{equation*}
\limsup _{t \rightarrow \infty}|u(t)|^{2} \leqslant \frac{|\Omega|}{\nu^{2} \lambda_{1}^{2}}+\alpha^{2}|\Omega| . \tag{4.5}
\end{equation*}
$$

### 4.3. An enstrophy bound

We note that $\nabla u$ has zero average over $\Omega$ by the periodicity of $u$. As a consequence, we have the Poincaré inequality

$$
\begin{equation*}
\lambda_{1}\|u\|^{2} \leqslant\left|A_{0} u\right|^{2} . \tag{4.6}
\end{equation*}
$$

Taking the scalar product of (2.8a) with $A_{0} u$, we have by the orthogonality property (2.6)

$$
\begin{align*}
\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\|u\|^{2}+\nu\left|A_{0} u\right|^{2} & \leqslant\left|\left(\theta e_{2}, A_{0} u\right)\right|  \tag{4.7}\\
& \leqslant|\theta|\left|A_{0} u\right| \leqslant \frac{1}{2 \nu}|\theta|^{2}+\frac{\nu}{2}\left|A_{0} u\right|^{2},
\end{align*}
$$

hence, by (4.1) and (4.6) we have

$$
\frac{\mathrm{d}}{\mathrm{~d} t}\|u\|^{2}+\nu \lambda_{1}\|u\|^{2} \leqslant \frac{1}{\nu}\left(|\Omega|+\Theta_{0}^{2} \mathrm{e}^{-2 \kappa t}\right),
$$

and thanks to the Gronwall inequality we obtain

$$
\begin{equation*}
\limsup _{t \rightarrow \infty}\|u(t)\|^{2} \leqslant z_{\max }:=\frac{|\Omega|}{\nu^{2} \lambda_{1}} \tag{4.8}
\end{equation*}
$$

Similar to the no-slip case analyzed in $[11,18]$, if $\|u(0)\| \leqslant M_{1},\|\theta(0)\| \leqslant M_{2}$, and $\varepsilon>0$ we have from (4.1), (4.5) and (4.8) that there exists $t_{0}=t_{0}\left(M_{1}, M_{2}, \varepsilon\right)$ such that

$$
\begin{align*}
&|\theta(t)|^{2} \leqslant|\Omega|+\varepsilon, \quad \forall t \geqslant t_{0}  \tag{4.9}\\
&|u(t)|^{2} \leqslant \frac{|\Omega|}{\nu^{2} \lambda_{1}^{2}}+\alpha^{2}|\Omega|+\varepsilon, \quad \forall t \geqslant t_{0}  \tag{4.10}\\
&\|u(t)\|^{2} \leqslant \frac{|\Omega|}{\nu^{2} \lambda_{1}}+\varepsilon, \quad \forall t \geqslant t_{0} \tag{4.11}
\end{align*}
$$

### 4.4. Bound on the temperature gradient

We start by taking the scalar product of (2.8b) with $A_{1} \theta=-\Delta \theta$, integrating by parts and applying the Cauchy-Schwarz and Young inequalities

$$
\begin{equation*}
\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\|\theta\|^{2}+\kappa\left|A_{1} \theta\right|^{2} \leqslant\left|\left(B_{1}(u, \theta), A_{1} \theta\right)\right|+\frac{\left|u_{2}\right|^{2}}{\kappa}+\frac{\kappa}{4}\left|A_{1} \theta\right|^{2} \tag{4.12}
\end{equation*}
$$

We apply integration by parts to rewrite the trilinear term as

$$
\begin{aligned}
\left(B_{1}(u, \theta), A_{1} \theta\right) & =-\sum_{i, j=1}^{2} \int_{\Omega} u_{i} \partial_{i} \theta \partial_{j}^{2} \theta \mathrm{~d} x \\
& =\sum_{i, j=1}^{2} \int_{\Omega} u_{i} \partial_{i j} \theta \partial_{j} \theta \mathrm{~d} x+\sum_{i, j=1}^{2} \int_{\Omega} \partial_{j} u_{i} \partial_{i} \theta \partial_{j} \theta \mathrm{~d} x .
\end{aligned}
$$

We then use the chain rule to rewrite the first sum, again apply integration by parts, and then incompressibility to find

$$
\begin{align*}
\sum_{i, j=1}^{2} \int_{\Omega} u_{i} \partial_{i j} \theta \partial_{j} \theta \mathrm{~d} x & =\frac{1}{2} \sum_{i, j=1}^{2} \int_{\Omega} u_{i} \partial_{i}\left(\partial_{j} \theta\right)^{2} \mathrm{~d} x \\
& =-\frac{1}{2} \int_{\Omega}\left(\partial_{1} u_{1}+\partial_{2} u_{2}\right)\left[\left(\partial_{1} \theta\right)^{2}+\left(\partial_{2} \theta\right)^{2}\right] \mathrm{d} x=0 \tag{4.13}
\end{align*}
$$

Applying the Hölder, Ladyzhenskaya and Young inequalities to each of the remaining four terms, we obtain

$$
\begin{align*}
\left|\left(B_{1}(u, \theta), A_{1} \theta\right)\right| & \leqslant 4\|u\|\|\nabla \theta\|_{L^{4}}^{2} \\
& \leqslant c_{1}\|u\|\|\theta\|\left|A_{1} \theta\right| \\
& \left.\leqslant \frac{c_{1}^{2}}{\kappa}\|u\|\|\theta\|\right)^{2}+\frac{\kappa}{4}\left|A_{1} \theta\right|^{2} . \tag{4.14}
\end{align*}
$$

Now combine (4.12) and (4.14) and the Poincaré inequality (4.3) so that

$$
\frac{\mathrm{d}}{\mathrm{~d} t}\|\theta\|^{2}+\kappa\left|A_{1} \theta\right|^{2} \leqslant \frac{2 c_{1}^{2}}{\kappa}\|u\|^{2}\|\theta\|^{2}+\frac{2\left|u_{2}\right|^{2}}{\kappa} \leqslant \frac{2 c_{1}^{2}}{\kappa}\|u\|^{2}\|\theta\|^{2}+\frac{2\|u\|^{2}}{\kappa \lambda_{1}} .
$$

We note that by the Cauchy-Schwarz inequality and (4.1),

$$
\|\theta\|^{2} \leqslant|\theta|\left|A_{1} \theta\right| \leqslant\left(|\Omega|^{1 / 2}+\Theta_{0} \mathrm{e}^{\kappa t}\right)\left|A_{1} \theta\right|,
$$

so that

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} t}\|\theta\|^{2} \leqslant-\frac{\kappa}{\left(|\Omega|^{1 / 2}+\varepsilon\right)^{2}}\|\theta\|^{4}+\frac{2 c_{1}^{2}}{\kappa}\|u\|^{2}\|\theta\|^{2}+\frac{2\|u\|^{2}}{\kappa \lambda_{1}} . \tag{4.15}
\end{equation*}
$$

Let $R_{1}=\left(|\Omega|^{1 / 2}+\varepsilon\right)^{2}, R_{2}=z_{\max }+\varepsilon$, for $\varepsilon$ as in (4.9)-(4.11). From (4.9), (4.11) and (4.15) and Young's inequality, we have for all $t \geqslant t_{0}$

$$
\begin{aligned}
\frac{\mathrm{d}}{\mathrm{~d} t}\|\theta\|^{2} & \leqslant-\frac{\kappa}{R_{1}}\|\theta\|^{4}+\frac{2 c_{1}^{2}}{\kappa}\|\theta\|^{2} R_{2}+\frac{2 R_{2}}{\kappa \lambda_{1}} \\
& \leqslant-\frac{\kappa}{2 R_{1}}\|\theta\|^{4}+\frac{2 c_{1}^{4}}{\kappa^{3}} R_{1} R_{2}^{2}+\frac{2 R_{2}}{\kappa \lambda_{1}} \\
& \leqslant-\frac{\kappa}{2 R_{1}}\left(\|\theta\|^{4}-K^{4}\right)
\end{aligned}
$$

where

$$
K^{4}=\frac{2 R_{1}}{\kappa}\left[\frac{2 c_{1}^{4}}{\kappa^{3}} R_{1} R_{2}^{2}+\frac{2 R_{2}}{\kappa \lambda_{1}}\right] .
$$

We claim that

$$
\begin{equation*}
\limsup _{t \rightarrow \infty}\|\theta(t)\|^{2} \leqslant 2\left[\frac{c_{1}^{4}}{\kappa^{4}}|\Omega|^{2} z_{\max }^{2}+\frac{|\Omega|}{\kappa^{2} \lambda_{1}} z_{\max }\right]^{1 / 2} \tag{4.16}
\end{equation*}
$$

To prove this we take $\varepsilon>0$, as above, and consider two possibilities.
Case I: If $\|\theta(t)\|^{2} \leqslant(1+4 \varepsilon)^{1 / 2} K^{2}$, for all $t \geqslant t_{0}$, then clearly

$$
\begin{equation*}
\limsup _{t \rightarrow \infty}\|\theta(t)\|^{2} \leqslant(1+4 \varepsilon)^{1 / 2} K^{2}, \quad \forall \varepsilon>0 \tag{4.17}
\end{equation*}
$$

Case II: Suppose there exists $t_{*} \geqslant t_{0}$ such that $\left\|\theta\left(t_{*}\right)\right\|^{2} \geqslant(1+4 \varepsilon)^{1 / 2} K^{2}$. We would then have that

$$
\frac{\mathrm{d}}{\mathrm{~d} t}\|\theta\|^{2} \leqslant-\frac{\kappa \varepsilon}{R_{1}} K^{4}, \quad \forall t \geqslant t_{*} \text { such that }\|\theta(t)\|^{2} \geqslant(1+2 \varepsilon)^{1 / 2} K^{2}
$$

We conclude that $\|\theta(t)\|^{2}$ is strictly decreasing at a rate faster than $-\kappa \varepsilon K^{4} /\left(2 R_{1}\right)$ for all $t \geqslant t_{*}$ such that $\|\theta(t)\|^{2} \geqslant(1+2 \varepsilon)^{1 / 2} K^{2}$. In particular, there exists $t_{* *}$, with $t_{*}<t_{* *}<\infty$ such that $\left\|\theta\left(t_{* *}\right)\right\|^{2}=(1+2 \varepsilon)^{1 / 2} K^{2}$. Moreover, for all $t>t_{* *}$ we have $\|\theta(t)\|^{2}<(1+2 \varepsilon)^{1 / 2} K^{2}$. As a result, we again obtain (4.17).

In either case we may now take $\varepsilon \rightarrow 0^{+}$to conclude (4.16). Introducing the Rayleigh and Prandtl numbers in (4.16) and using the concavity of the square root function, we arrive at the bounding expression

$$
\limsup _{t \rightarrow \infty}\|\theta(t)\|^{2} \lesssim \vartheta_{\max }:=|\Omega| z_{\max } \operatorname{RaPr}+\left(\frac{|\Omega|}{\lambda_{1}} z_{\max } \operatorname{RaPr}\right)^{1 / 2}
$$

Thus, the ball $\mathcal{B}_{\alpha}(\varepsilon) \subset V_{0} \times V_{1}$, defined by

$$
\mathcal{B}_{\alpha}(\varepsilon):=\left\{(u, \theta):\|u\|_{H^{1}}^{2} \leqslant \frac{1+\lambda_{1}}{\nu^{2} \lambda_{1}^{2}}|\Omega|+\alpha^{2}|\Omega|+2 \varepsilon, \quad\|\theta\|^{2} \leqslant \vartheta_{\max }\right\}
$$

is absorbing. This gives for each $\alpha$ the existence of a global attractor $\mathcal{A}_{\alpha}$, within the invariant subspace of solutions $(u, \theta)$ where the spatial average of velocity is fixed at $\alpha$. The global attractor is contained in $\mathcal{B}_{\alpha}(0)$.

### 4.5. Palinstrophy bound

To estimate palinstrophy on $\mathcal{A}_{\alpha}$ we follow [6] almost verbatim except that the effect of time independent forcing of the NSE is played by the bound $\|\theta\|^{2} \leqslant \vartheta_{\text {max }}$. The other difference is that our velocity is not normalized as in [6]. For completeness, and in order to arrive at an overall bound in terms of $\nu, \kappa$, we distill the essential argument here.

Returning to (4.7), we integrate by parts, and then apply the Cauchy-Schwarz inequality to get

$$
-\|u\| \sqrt{\vartheta_{\max }} \leqslant \frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\|u\|^{2}+\nu\left|A_{0} u\right|^{2} \leqslant\|u\| \sqrt{\vartheta_{\max }}, \quad \forall(u, \theta) \in \mathcal{A}_{\alpha}
$$

We denote

$$
\begin{equation*}
z=\|u\|^{2}, \quad q=\left|A_{0} u\right|^{2}, \quad \zeta=\left|A_{0}^{3 / 2} u\right|^{2}, \quad \vartheta=\|\theta\|^{2} \tag{4.18}
\end{equation*}
$$

Then whenever

$$
\begin{equation*}
\|u\| \sqrt{\vartheta_{\max }} \leqslant \frac{\nu}{2}\left|A_{0} u\right|^{2}, \quad \text { equivalently } q \geqslant \frac{2}{\nu} \sqrt{z \vartheta_{\max }} \tag{4.19}
\end{equation*}
$$

we have

$$
\begin{equation*}
-3 \nu q \leqslant \frac{\mathrm{~d} z}{\mathrm{~d} t} \leqslant-\nu q \tag{4.20}
\end{equation*}
$$

Setting $w=A_{0} u$ in (2.7) and applying Agmon's inequality, we have

$$
\left|\left(B_{0}(u, u), A_{0}^{2} u\right)\right|=\left|\left(B_{0}\left(A_{0} u, u\right), A_{0} u\right)\right| \leqslant c_{2}\left|A_{0} u\right|^{2}\|u\|^{1 / 2}\left|A_{0}^{3 / 2} u\right|^{1 / 2}
$$

We next take the scalar product of (2.8a) with $A_{0}^{2} u$, and integrate by parts to obtain

$$
\begin{align*}
\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\left|A_{0} u\right|^{2}+\nu\left|A_{0}^{3 / 2} u\right|^{2} & \leqslant\left|\left(\theta, A_{0}^{2} u\right)\right|+\left|\left(B_{0}(u, u), A_{0}^{2} u\right)\right|  \tag{4.21}\\
& \leqslant\|\theta\|\left|A_{0}^{3 / 2} u\right|+c_{2}\left|A_{0} u\right|^{2}\|u\|^{1 / 2}\left|A_{0}^{3 / 2} u\right|^{1 / 2}
\end{align*}
$$

Note that by the Cauchy-Schwarz inequality

$$
\begin{equation*}
\zeta:=\left|A_{0}^{3 / 2} u\right|^{2} \geqslant \frac{\left|A_{0} u\right|^{4}}{\|u\|^{2}}=\frac{q^{2}}{z} \geqslant \frac{4}{\nu^{2}} \vartheta_{\max } \tag{4.22}
\end{equation*}
$$

in the region

$$
\begin{equation*}
\mathcal{R}:=\left\{(z, q) \left\lvert\, q \geqslant \frac{2}{\nu} \sqrt{z \vartheta_{\max }}\right.\right\} . \tag{4.23}
\end{equation*}
$$

It follows that

$$
\|\theta\|\left|A_{0}^{3 / 2} u\right|=\vartheta^{1 / 2} \zeta^{1 / 2} \leqslant \frac{\nu}{2} \zeta \quad \forall(z, q) \in \mathcal{R},
$$

and hence, as in [6],

$$
\begin{equation*}
\frac{\mathrm{d} q}{\mathrm{~d} t} \leqslant \psi(\zeta):=-\nu \zeta+2 c_{2} q z^{1 / 4} \zeta^{1 / 4} \tag{4.24}
\end{equation*}
$$

To close the system (eliminate $\zeta$ ) we find that the maximum of $\psi$ is achieved at

$$
\zeta_{\max }:=\left(\frac{c_{2}}{2 \nu} q z^{1 / 4}\right)^{4 / 3} \quad \text { with a value } \psi_{\max }=3 \nu \zeta_{\max }
$$

We note that

$$
\frac{q^{2}}{z} \geqslant \zeta_{\max } \quad \text { if and only if } \quad q \geqslant\left(\frac{c_{2}}{2 \nu} z\right)^{2}
$$

so that by (4.22)

$$
\begin{equation*}
\frac{\mathrm{d} q}{\mathrm{~d} t} \leqslant \psi_{\max }=\frac{3}{\nu^{1 / 3}}\left(\frac{c_{2}}{2} q z^{1 / 4}\right)^{4 / 3} \quad \text { if } q \leqslant\left(\frac{c_{2}}{2 \nu} z\right)^{2} \tag{4.25}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\mathrm{d} q}{\mathrm{~d} t} \leqslant \psi\left(q^{2} / z\right)=-\nu \frac{q^{2}}{z}+2 c_{2} q^{3 / 2} \quad \text { if } q \geqslant\left(\frac{c_{2}}{2 \nu} z\right)^{2} \tag{4.26}
\end{equation*}
$$

We see that

$$
\begin{equation*}
\frac{\mathrm{d} q}{\mathrm{~d} t} \leqslant 0 \quad \text { if } q \geqslant\left(\frac{2 c_{2}}{\nu} z\right)^{2} \quad \text { and } \quad q \geqslant \frac{2}{\nu} \sqrt{z \vartheta_{\max }} \tag{4.27}
\end{equation*}
$$

By considering the steepest descent possible below

$$
q=\left(\frac{2 c_{2}}{\nu} z\right)^{2}
$$

and the most shallow ascent possible above this parabola, we find three bounding curves $q=f_{j}(z), j=1,2,3$, after solving, in order, three final value problems. The first combines the (positive) bound in (4.25) with the upper bound in (4.20)

$$
\begin{aligned}
& \frac{\mathrm{d} q}{\mathrm{~d} z}=-3\left(\frac{c_{2}}{2 \nu}\right)^{4 / 3}(q z)^{1 / 3}, \quad \text { for } z_{1} \leqslant z \leqslant z_{0}=z_{\max }=\frac{|\Omega|}{\nu^{2} \lambda_{1}} \\
& q\left(z_{0}\right)=q_{0}:=\frac{2}{\nu^{2}} z_{\max }^{1 / 2} \vartheta_{\max }^{1 / 2} .
\end{aligned}
$$

The second picks up where the first leaves off and combines the (positive) bound in (4.26) with the upper bound in (4.20)

$$
\begin{aligned}
& \frac{\mathrm{d} q}{\mathrm{~d} z}=\frac{q}{z}-\frac{2 c_{2}}{\nu} q^{1 / 2}, \quad \text { for } z_{2} \leqslant z \leqslant z_{1} \\
& q\left(z_{1}\right)=q_{1}
\end{aligned}
$$

while the third combines the (negative) bound in (4.26) with the lower bound in (4.20)

$$
\begin{aligned}
& \frac{\mathrm{d} q}{\mathrm{~d} z}=\frac{q}{3 z}-\frac{2 c_{2}}{3 \nu} q^{1 / 2}, \quad \text { for } 0 \leqslant z \leqslant z_{2} \\
& q\left(z_{2}\right)=q_{2}
\end{aligned}
$$

where $q_{1}, q_{2}$ are determined by the intersections of $f_{1}$ and $f_{2}$ (defined below) with the parabolas

$$
q=\left(\frac{c_{2}}{2 \nu} z\right)^{2} \quad \text { and } \quad q=\left(\frac{2 c_{2}}{\nu} z\right)^{2}
$$

respectively. This results in a convex function in $z$

$$
f_{1}(z):=\left[\frac{3}{2}\left(\frac{c_{2}}{2 \nu}\right)^{4 / 3}\left(z_{0}^{4 / 3}-z^{4 / 3}\right)+q_{0}^{2 / 3}\right]^{3 / 2}
$$

and concave functions in $z$

$$
\begin{aligned}
& f_{2}(z):=\frac{1}{\nu^{2}}\left[-2 c_{2} z+\left(\nu q_{1}^{1 / 2}+2 c_{2} z_{1}\right)\left(\frac{z}{z_{1}}\right)^{1 / 2}\right]^{2} \\
& f_{3}(z):=\frac{1}{25 \nu^{2}}\left[-6 c_{2} z+\left(5 \nu q_{2}^{1 / 2}+6 c_{2} z_{2}\right)\left(\frac{z}{z_{2}}\right)^{1 / 6}\right]^{2}
\end{aligned}
$$

A qualitative sketch of these three curves is shown in figure 1. It is shown in [6] that the curve $q=f_{3}(z)$ does not intersect the curve $q=2 \sqrt{z \vartheta_{\max }} / \nu$. Let

$$
f(z):= \begin{cases}f_{1}(z) & \text { if } z_{1} \leqslant z \leqslant z_{\max }  \tag{4.28}\\ f_{2}(z) & \text { if } z_{2} \leqslant z<z_{1} \\ f_{3}(z) & \text { if } 0 \leqslant z \leqslant z_{2}\end{cases}
$$

To prove (3.4), suppose there is an element in $\mathcal{A}_{\alpha}$ such that $q(0)>f(z(0))$. The solution through any element in $\mathcal{A}_{\alpha}$ exists for all negative time. If $q(t)>f(z(t))$ for all $t<0$, since $q(t)$ increases with negative time, as long as $z(t)<z_{2}$, we have $q(t)>\min \left\{q(0), q_{0}\right\}$. By the upper bound in (4.20), $z(t)$ would then exceed $z_{2}$ in finite negative time. Thus, we must have $q(t) \leqslant f(z(t))$ at some $t<0$. But forward in time, the region $q \leqslant f(z)$ is invariant, contradicting the assumption that the initial condition satisfied $q(0)>f(z(0))$.

We now find an overall bound on palinstrophy in $\mathcal{A}_{\alpha}$. A straightforward calculation shows that substituting

$$
q_{1}=\left(\frac{c_{2} z_{1}}{2 \nu}\right)^{2} \quad \text { into } q_{2}=f_{2}\left(z_{2}\right)=\left(\frac{2 c_{2} z_{2}}{\nu}\right)^{2}
$$

reduces to

$$
z_{2}=\frac{25}{64} z_{1}
$$

Similarly, using

$$
q_{0}=\frac{2}{\nu} \sqrt{z_{0} \vartheta_{\max }} \quad \text { in } q_{1}=f_{1}\left(z_{1}\right)=\left(\frac{c_{2} z_{1}}{2 \nu}\right)^{2}
$$

leads to

$$
z_{1} \leqslant\left[\frac{3}{2} z_{0}^{4 / 3}+\frac{4 \nu^{2 / 3}}{c_{2}^{4 / 3}} z_{0}^{1 / 3} \vartheta_{\max }^{1 / 3}\right]^{3 / 4} \lesssim z_{\max }+\nu^{1 / 2} z_{\max }^{1 / 4} \vartheta_{\max }^{1 / 4}
$$

so that

$$
\begin{equation*}
\left|A_{0} u\right|^{2} \leqslant q_{2} \lesssim q_{\max }:=\frac{z_{\max }^{2}}{\nu^{2}}+\frac{z_{\max }^{1 / 2}}{\nu} \vartheta_{\max }^{1 / 2} . \tag{4.29}
\end{equation*}
$$

### 4.6. A bound on $\left|A_{1} \theta\right|$

From (4.12) and (4.14) we have

$$
-\frac{c_{3}}{\kappa}\|u\|^{2}\|\theta\|^{2}-\frac{2\|u\|^{2}}{\kappa \lambda_{1}} \leqslant \frac{\mathrm{~d}}{\mathrm{~d} t}\|\theta\|^{2}+\kappa\left|A_{1} \theta\right|^{2} \leqslant \frac{c_{3}}{\kappa}\|u\|^{2}\|\theta\|^{2}+\frac{2\|u\|^{2}}{\kappa \lambda_{1}} .
$$

Thus, if

$$
\left|A_{1} \theta\right|^{2} \geqslant \frac{2}{\kappa^{2}}\left(c_{3}\|u\|^{2}\|\theta\|^{2}+\frac{2\|u\|^{2}}{\lambda_{1}}\right)
$$

it follows that

$$
\begin{equation*}
-\frac{3}{2} \kappa\left|A_{1} \theta\right|^{2} \leqslant \frac{\mathrm{~d}}{\mathrm{~d} t}\|\theta\|^{2} \leqslant-\frac{1}{2} \kappa\left|A_{1} \theta\right|^{2} \tag{4.30}
\end{equation*}
$$

We next take the scalar product of the temperature equation with $A_{1}^{2} \theta=\Delta^{2} \theta$ and using the fact that $A_{0} u=-\Delta u$, write

$$
\begin{equation*}
\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\left|A_{1} \theta\right|^{2}+\kappa\left|A_{1}^{3 / 2} \theta\right|^{2} \leqslant\left|A_{0} u\right|\left|A_{1} \theta\right|+\left|\left(B_{1}(u, \theta), A_{1}^{2} \theta\right)\right| \tag{4.31}
\end{equation*}
$$

We need to move two derivatives in the trilinear term in order to ultimately obtain a bound for it in which the highest order norm is $\left|A_{1}^{3 / 2} \theta\right|$. We integrate by parts to write

$$
\begin{aligned}
\left(B_{1}(u, \theta), A_{1}^{2} \theta\right) & =\sum_{i, j, k=1}^{2} \int_{\Omega} u_{i} \partial_{i} \theta \partial_{j}^{2} \partial_{k}^{2} \theta \mathrm{~d} x \\
& =-\sum_{i, j, k=1}^{2} \int_{\Omega} u_{i} \partial_{i j} \theta \partial_{j} \partial_{k}^{2} \theta \mathrm{~d} x-\sum_{i, j, k=1}^{2} \int_{\Omega} \partial_{j} u_{i} \partial_{i} \theta \partial_{j} \partial_{k}^{2} \theta \mathrm{~d} x=I+I I .
\end{aligned}
$$

We then integrate the first summation by parts

$$
I=\sum_{i, j, k=1}^{2} \int_{\Omega} u_{i} \partial_{i} \partial_{j}^{2} \theta \partial_{k}^{2} \theta \mathrm{~d} x+\sum_{i, j, k=1}^{2} \int_{\Omega} \partial_{j} u_{i} \partial_{i, j} \theta \partial_{k}^{2} \theta \mathrm{~d} x=I_{a}+I_{b}
$$

and split the resulting first summation as

$$
I_{a}=\sum_{i, j=1}^{2} \int_{\Omega} u_{i} \partial_{i} \partial_{j}^{2} \theta \partial_{j}^{2} \theta \mathrm{~d} x+\sum_{i, j \neq k=1}^{2} \int_{\Omega} u_{i} \partial_{i} \partial_{j}^{2} \theta \partial_{k}^{2} \theta \mathrm{~d} x=I_{a_{1}}+I_{a_{2}}
$$

Proceeding as in (4.13), we find that $I_{a_{1}}=0$. Integrating by parts again, we have

$$
I_{a_{2}}=-\sum_{i, j \neq k=1}^{2} \int_{\Omega} \partial_{i} u_{i} \partial_{j}^{2} \theta \partial_{k}^{2} \theta \mathrm{~d} x-\sum_{i, j \neq k=1}^{2} \int_{\Omega} u_{i} \partial_{j}^{2} \theta \partial_{i} \partial_{k}^{2} \theta \mathrm{~d} x
$$

Since the first sum is zero by incompressibility, we have by symmetry that $I_{a_{2}}=-I_{a_{2}}$, and thus $I_{a_{2}}=0$. Integrating by parts one more time, we have

$$
I I=\sum_{i, j, k=1}^{2} \int_{\Omega} \partial_{j}^{2} u_{i} \partial_{i} \theta \partial_{k}^{2} \theta \mathrm{~d} x+I_{b}
$$

After gathering what remains, we use Agmon's and Ladyzhenskaya's inequalities to estimate the trilinear term as

$$
\begin{aligned}
\left|\left(B_{1}(u, \theta), A_{1}^{2} \theta\right)\right|= & \left|\sum_{i, j, k=1}^{2} \int_{\Omega} \partial_{j}^{2} u_{i} \partial_{i} \theta \partial_{k}^{2} \theta \mathrm{~d} x+2 \sum_{i, j, k=1}^{2} \int_{\Omega} \partial_{j} u_{i} \partial_{i, j} \theta \partial_{k}^{2} \theta \mathrm{~d} x\right| \\
\leqslant & c\left|A_{0} u\right|\|\theta\|^{1 / 2}\left|A_{1}^{3 / 2} \theta\right|^{1 / 2}\left|A_{1} \theta\right| \\
& +\left.c\|u\|^{1 / 2}\left|A_{0} u\right|^{1 / 2}\left|\theta \|_{H^{2}}\right| A_{1} \theta\right|^{1 / 2}\left|A_{1}^{3 / 2} \theta\right|^{1 / 2} \\
\leqslant & c_{4}\left|A_{0} u\right| \frac{\left|A_{1} \theta\right|^{3 / 2}}{\lambda_{1}^{1 / 4}}\left|A_{1}^{3 / 2} \theta\right|^{1 / 2}=\frac{c_{4}}{\lambda_{1}^{1 / 4}} q^{1 / 2} \eta^{3 / 4} \xi^{1 / 4}
\end{aligned}
$$

where $\eta=\left|A_{1} \theta\right|^{2}, \xi=\left|A_{1}^{3 / 2} \theta\right|^{2}$ and for convenience in what follows, we take $c_{4}=2 \max \left(c, c_{3}\right)$.
Using this in (4.31), we find

$$
\begin{aligned}
\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\left|A_{1} \theta\right|^{2}+\kappa\left|A_{1}^{3 / 2} \theta\right|^{2} & \leqslant\left|A_{0} u\right|\left|A_{1} \theta\right|+\frac{c_{4}}{\lambda_{1}^{1 / 4}}\left|A_{0} u \| A_{1} \theta\right|^{3 / 2}\left|A_{1}^{3 / 2} \theta\right|^{1 / 2} \\
& \leqslant \frac{2 c_{4}}{\lambda_{1}^{1 / 4}}\left|A_{0} u \| A_{1} \theta\right|^{3 / 2}\left|A_{1}^{3 / 2} \theta\right|^{1 / 2}
\end{aligned}
$$

Thus, invoking our palinstrophy bound $q_{\max }$, we have

$$
\frac{\mathrm{d}}{\mathrm{~d} t} \eta \leqslant \Phi(\xi):=-2 \kappa \xi+\frac{4 c_{4}}{\lambda_{1}^{1 / 4}} q_{\max }^{1 / 2} \eta^{3 / 4} \xi^{1 / 4}
$$

We find that

$$
\Phi(\xi) \leqslant \Phi_{\max }=\frac{2}{\kappa^{1 / 3}}\left(\frac{c_{4}}{2 \lambda_{1}^{1 / 4}}\right)^{4 / 3} q_{\max }^{2 / 3} \eta
$$

and that

$$
\Phi(\xi) \leqslant 0 \quad \forall \xi \geqslant \xi^{*}:=\gamma \eta, \quad \text { where } \gamma:=\left(\frac{2 c_{4}}{\kappa \lambda_{1}^{1 / 4}}\right)^{4 / 3} q_{\max }^{2 / 3} .
$$

In terms of $z_{0}$, our enstrophy bound on the attractor, (4.30) holds for

$$
\begin{equation*}
\eta \geqslant g_{3}(\vartheta):=\frac{z_{\max }}{\kappa^{2}}\left(c_{4} \vartheta+\frac{4}{\lambda_{1}}\right) . \tag{4.32}
\end{equation*}
$$

Once again, by the Cauchy-Schwarz inequality, we have

$$
\left|A_{1}^{3 / 2} \theta\right| \geqslant \frac{\left|A_{1} \theta\right|^{2}}{\|\theta\|}, \quad \text { i.e., } \xi \geqslant \frac{\eta^{2}}{\vartheta} .
$$

Thus for

$$
\frac{\eta^{2}}{\vartheta} \leqslant \xi^{*}, \quad \text { equivalently } \eta \leqslant \gamma \vartheta
$$

we combine

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} t} \eta \leqslant \Phi_{\max } \quad \text { with } \frac{\mathrm{d}}{\mathrm{~d} t} \vartheta \leqslant-\frac{\kappa}{2} \eta \tag{4.33}
\end{equation*}
$$

and solve

$$
\frac{\mathrm{d} \eta}{\mathrm{~d} \vartheta}=-\gamma_{0}, \quad \eta\left(\vartheta_{\max }\right)=\eta_{0}:=\frac{z_{\max }}{\kappa^{2}}\left(c_{4} \vartheta_{\max }+\frac{4}{\lambda_{1}}\right), \quad \text { where } \gamma_{0}=4^{-1 / 3} \gamma
$$

to find a straight-line solution

$$
\eta=g_{1}(\vartheta):=\eta_{0}-\gamma_{0}\left(\vartheta-\vartheta_{\max }\right) .
$$

We then find the intersection of this line with $\eta=\gamma \vartheta$ to be at $\left(\vartheta_{1}, \eta_{1}\right)$, where

$$
\begin{equation*}
\vartheta_{1}=\frac{c_{4} z_{\max } / \kappa^{2}+\gamma_{0}}{\gamma+\gamma_{0}} \vartheta_{\max }+\frac{4 z_{\max }}{\kappa^{2} \lambda_{1}\left(\gamma+\gamma_{0}\right)}, \quad \text { and } \quad \eta_{1}=\gamma \vartheta_{1} . \tag{4.34}
\end{equation*}
$$

For $\eta \geqslant \gamma \vartheta$ we combine

$$
\frac{\mathrm{d}}{\mathrm{~d} t} \eta \leqslant \Phi\left(\eta^{2} / \vartheta\right)=-2 \kappa \frac{\eta^{2}}{\vartheta}+\frac{4 c_{4}}{\lambda_{1}^{1 / 4}} q_{\max }^{1 / 2} \frac{\eta^{5 / 4}}{\vartheta^{1 / 4}} \quad \text { with } \frac{\mathrm{d}}{\mathrm{~d} t} \vartheta \geqslant-\frac{3}{2} \kappa \eta
$$

and solve

$$
\frac{\mathrm{d} \eta}{\mathrm{~d} \vartheta}=\frac{4}{3 \vartheta} \eta-\frac{8 c_{4}}{3 \lambda_{1}^{1 / 4} \kappa} q_{\max }^{1 / 2} \frac{\eta^{1 / 4}}{\vartheta^{1 / 4}}
$$

to find

$$
\eta=g_{2}(\vartheta):=\left[\left(\frac{\vartheta}{\vartheta_{1}}\right) \eta_{1}^{1 / 4}+\tilde{\gamma}\left(\vartheta^{3 / 4}-\vartheta \vartheta_{1}^{-1 / 4}\right)\right]^{4 / 3},
$$

where

$$
\tilde{\gamma}=\frac{8 c_{4}}{\lambda_{1}^{1 / 4} \kappa} q_{\max }^{1 / 2}
$$

As we argued in section 4.5 , if an element in the global attractor were to project in the $\vartheta-\eta$ plane above

$$
\begin{equation*}
\eta=\max \left\{g_{1}(\vartheta), g_{2}(\vartheta), g_{3}(\vartheta)\right\} \tag{4.35}
\end{equation*}
$$

then by (4.33) the solution through it would, in finite negative time, have to enter the region below the curves in (4.35). Yet, this region is invariant. We conclude from (4.34) and (4.29) that we have an overall bound on the global attractor of

$$
\left|A_{1} \theta\right|^{2} \leqslant \eta_{1} \lesssim \eta_{\max }:=\frac{z_{\max }}{\kappa^{2}} \vartheta_{\max }+\gamma \vartheta_{\max }+\frac{z_{\max }}{\kappa^{2} \lambda_{1}} .
$$

A qualitative sketch of the region bounding the global attractor in this plane is shown in figure 2.

## 5. Implications for data assimilation

Suppose reality is represented by a particular solution to an evolution equation

$$
\frac{\mathrm{d} v}{\mathrm{~d} t}=F(v)
$$

where the initial data $v(0)$ is not known. Instead continuous data of the form $I_{h} v(t)$ is known over an interval, $t \in\left[t_{1}, t_{2}\right]$, for a certain type of interpolating operator $I_{h}$ with spatial resolution $h$. The nudging approach to data assimilation amounts to solving the auxiliary system

$$
\begin{equation*}
\frac{\mathrm{d} \tilde{v}}{\mathrm{~d} t}=F(\tilde{v})-\mu I_{h}(\tilde{v}-v) \tag{5.1}
\end{equation*}
$$

using any initial condition, e.g., $\tilde{v}_{0}=0$. It was shown in $[2,3]$ that if $\mu>0$ is sufficiently large, and correspondingly, $h$ sufficiently small, then $v(t)-\tilde{v}(t) \rightarrow 0$, in some norm, at an exponential rate, as $t \rightarrow \infty$. In fact, computations indicate that this approach works with data that is much more coarse than suggested by rigorous estimates (see [1, 7, 8, 12]). Flexibility in the choice of interpolant is one of the main advantages of injecting the observed data through a feedback nudging term, rather than into terms involving spatial derivatives [2, 15]. Numerical errors are shown to be bounded uniformly in time for semi-discrete [16] and fully discrete schemes [14] for (5.1).

Now consider this approach for the stress-free RB system (2.8) using data from only the horizontal component of velocity. This means solving the auxiliary system

$$
\begin{aligned}
& \frac{\mathrm{d} \tilde{u}}{\mathrm{~d} t}+\nu A_{0} \tilde{u}+B_{0}(\tilde{u}, \tilde{u})=\mathbb{P}_{\sigma}\left(\tilde{\theta} e_{2}\right)-\mu \mathbb{P}_{\sigma} I_{h}\left(\tilde{u}_{1}-u_{1}\right) e_{1} \\
& \frac{\mathrm{~d} \tilde{\theta}}{\mathrm{~d} t}+\kappa A_{1} \tilde{\theta}+B_{1}(\tilde{u}, \tilde{\theta})=\tilde{u} \cdot e_{2} \\
& \tilde{u}(0 ; x)=0, \quad \tilde{\theta}(0 ; x)=0 .
\end{aligned}
$$

It was proved in [9] that if $\mu h^{2} \lesssim \nu$ and

$$
\begin{equation*}
\mu \geqslant K_{1} \sim \frac{1}{\kappa \lambda_{1}}+\frac{1}{\nu \kappa^{2}}+\frac{1}{\kappa}+\frac{\left|A_{0} u\right|^{2}}{\nu} \tag{5.2}
\end{equation*}
$$

then

$$
\|u(t)-\tilde{u}(t)\|+|\theta(t)-\tilde{\theta}(t)| \rightarrow 0 \quad \text { as } t \rightarrow \infty
$$

at an exponential rate. Also shown there was that if

$$
\begin{equation*}
\mu \geqslant K_{2} \sim K_{1}+\frac{1}{\kappa}\|\theta\|^{2}\left|A_{1} \theta\right|^{2}, \tag{5.3}
\end{equation*}
$$



Figure 3. Projections after a transient period.
then the stronger convergence

$$
\|u(t)-\tilde{u}(t)\|+\|\theta(t)-\tilde{\theta}(t)\| \rightarrow 0, \quad \text { as } t \rightarrow \infty
$$

holds at an exponential rate. The bounds in this paper on $\|\theta\|,\left|A_{0} u\right|$ and $\left|A_{1} \theta\right|$ are all algebraic, suggesting that data assimilation by nudging with just the horizontal velocity could be effective for the stress-free RB system. We present computational evidence to this effect in the next section.

## 6. Computational results

The computations presented below were done using Dedalus, an open-source package for solving partial differential equations using pseudo-spectral methods (see [4]). The time stepping is done by a four-stage third order Runge-Kutta method.

We solve (2.1) with $L=2$ in the physical domain $\Omega_{0}=(0, L) \times(0,1)$. The physical parameters of viscosity and thermal diffusivity are related to the Rayleigh and Prandtl numbers


Figure 4. Ratios in (6.1).
through

$$
\nu=\sqrt{\frac{\mathrm{Pr}}{\mathrm{Ra}}}, \quad \kappa=\frac{1}{\sqrt{\mathrm{Ra} \cdot \operatorname{Pr}}} .
$$

We take $\operatorname{Pr}=1$ so that in our dimensionless variables $\operatorname{Ra}:=(\nu \kappa)^{-1}=\nu^{-2}$ and use $n_{\mathrm{F}}$ Fourier modes in the $x_{1}$-direction and $n_{\mathrm{C}}$ Chebyshev modes in the $x_{2}$-direction. The numbers of modes used are $n_{\mathrm{F}} \times n_{\mathrm{C}}=256 \times 128,1024 \times 512$, and $2048 \times 1024$ for runs at $\mathrm{Ra}=10^{6}, 10^{7}, 10^{8}$ respectively.

### 6.1. Sharpness

Each plot in figure 3 shows the projection of a solution after a transient phase in a plane spanned by the norms bounded by our analysis. The solutions are plotted over the time period $200 \leqslant t \leqslant$ 1000 for $\mathrm{Ra}=10^{6}, 10^{7}$ and over $200 \leqslant t \leqslant 1485$ for $\mathrm{Ra}=10^{8}$ (time units in the RB system (2.1)). The initial condition in each case is $\left(u_{0}, \theta_{0}\right)=(0,0)$ so the average $\alpha$ of the horizontal velocity is zero.

It is not surprising that our rigorous overall bounds as well as the curves in figures 1 and 2 are orders of magnitude greater than the norms of these solutions. Plotting the bounds and curves together with the solutions is not revealing. Instead, to see a trend in sharpness, we plot in figure 4 the ratios

$$
\begin{equation*}
\frac{z_{\max }}{\max _{\mathcal{A}} z}, \quad \frac{\vartheta_{\max }}{\max _{\mathcal{A}} \vartheta}, \quad \frac{q_{\max }}{\max _{\mathcal{A}} q}, \quad \frac{\eta_{\max }}{\max _{\mathcal{A}} \eta} . \tag{6.1}
\end{equation*}
$$

Using the numerical values for the $z$ ratio, we gauge the (highest) power in (3.2) to be inflated (at least over this range of the Rayleigh number), by an addition of $\beta$, where

$$
100^{\beta}=\frac{3.14 \times 10^{6}}{7.05 \times 10^{4}}, \quad \text { i.e., } \beta=0.824 \text {. }
$$



Figure 5. Data assimilation at $\operatorname{Ra}=10^{6}$ and $\operatorname{Pr}=1$ with $h \approx 0.196$.

A similar calculation for the $\vartheta$ ratio gives $\beta=1.68$. We note that the curves are bending favorably for the ratios for $q=$ palinstrophy and $\eta=\left|A_{1} \theta\right|$.

### 6.2. Data assimilation

Nudging is carried out at $\mathrm{Ra}=10^{6}$ using the interpolant operator $I_{h}$ at every $m$ th nodal value in each direction, i.e.,

$$
h(m)=\max \left\{h_{\mathrm{F}}(m), h_{\mathrm{C}}(m)\right\}, \quad h_{\mathrm{F}}(m)=\frac{m L}{n_{\mathrm{F}}}
$$

where $m$ is a positive integer, and

$$
\begin{aligned}
h_{\mathrm{C}}(m) & =\max \left\{\left|x_{2}^{i m}-x_{2}^{(i+1) m}\right|: i=0,1, \ldots,\left\lfloor n_{\mathrm{C}} / m\right\rfloor-1\right\} \\
& =\frac{m \pi}{2 n_{\mathrm{C}}} \sin (\xi), \quad \text { for some } \quad \xi \in\left(\frac{2 m i-1}{2 n_{\mathrm{C}}} \pi, \frac{2 m(i+1)-1}{2 n_{\mathrm{C}}} \pi\right) \\
& \approx \frac{m \pi}{2 n_{\mathrm{C}}}
\end{aligned}
$$



Figure 6. Data assimilation at $\mathrm{Ra}=10^{6}$ and $\operatorname{Pr}=1$ with $h \approx 0.393$.
where $\left(x_{2}^{j}\right)$ are the Chebyshev grid points in the $x_{2}$-direction of the physical space. For $n_{\mathrm{F}} \times n_{\mathrm{C}}=256 \times 128$, this means $h(16) \approx 0.196$ and $h(32) \approx .0 .393$. The nudging parameter is fixed at $\mu=1$.

Figure 5 shows that at $h=0.196$ the solution to the data assimilation system converges to the reference solution at an exponential rate. At $h=0.393$ the error appears to saturate around $10^{-3}$ during rapid oscillations (see figure 6). We found that at $h=h(64)=0.785$ the nudged solution does not converge to the reference at all (not shown). This demonstrates a critical value of $h$.

Data assimilation by nudging works much more effectively than the rigorous analysis can guarantee. The value of $\mu$ and corresponding resolution $h$ of the data suggested by the conditions (5.2) and (5.3) are based on compounded, conservative estimates derived using general inequalities which are not saturated by 2D convective flows. In addition, as demonstrated in (6.1), our algebraic rigorous estimates for $\|\theta\|,\left|A_{0} u\right|$, and $\left|A_{1} \theta\right|$ in this case of stress-free boundary conditions, though much better than the exponential bounds previously found for the no-slip case in [11], are still somewhat artificially inflated. Numerical nudging tests in [8] for the RB system with no-slip boundary conditions suggest that better bounds on the attractor might hold in that case as well. The key here in the stress-free case was extending the physical domain to be fully periodic, hence there is effectively no boundary. Since, in the no-slip case one is
unable to remove the physical boundary, one should have to resolve the boundary layer scales in order to determine the behavior of the solutions. This is even more pronounced in the estimates of the dimension of the global attractor of the 2D NSE with no-slip boundary conditions in comparison to the case with periodic boundary conditions. Thus, improving the bounds in the no-slip case would require entirely different techniques.

## Acknowledgments

The authors acknowledge the Indiana University Pervasive Technology Institute (see [17]) for providing HPC (Big Red II, Carbonate), storage resources that have contributed to the research results reported within this paper. The work of Y Cao was supported in part by National Science Foundation Grant DMS-1418911, that of M S Jolly by NSF Grant DMS-1818754. The work of E S Titi was supported in part by the Einstein Visiting Fellow Program, and by the John Simon Guggenheim Memorial Foundation. J P Whitehead acknowledges support from the Simons Foundation through award number 586788, and the hospitality of the Department of Mathematics at Indiana University where part of this work was instigated.

## ORCID iDs

Michael S Jolly © https://orcid.org/0000-0002-7158-0933
Edriss S Titi © https://orcid.org/0000-0002-5004-1746

## References

[1] Altaf M U, Titi E S, Gebrael T, Knio O M, Zha L, Mccabe M F and Hoteit I 2017 Downscaling the 2D Bénard convection equations using continuous data assimilation Comput. Geosci. 21 393-410
[2] Azouani A, Olson E and Titi E S 2014 Continuous data assimilation using general interpolant observables J. Nonlinear Sci. 24 277-304
[3] Azouani A and Titi E S 2014 Feedback control of nonlinear dissipative systems by finite determining parameters-a reaction-diffusion paradigm Evol. Equ. Control Theory 3 579-94
[4] Burns K J, Vasil G M, Oishi J S, Lecoanet D and Brown B 2020 Dedalus: a flexible framework for numerical simulations with spectral methods Phys. Rev. Res. 2023068
[5] Constantin P and Foias C 1988 Navier-Stokes Equations (Chicago Lectures in Mathematics) (Chicago, IL: University of Chicago Press)
[6] Dascaliuc R, Foias C and Jolly M S 2010 Estimates on enstrophy, palinstrophy, and invariant measures for 2D turbulence J. Differ. Equ. 248 792-819
[7] DiLeoni P C, Mazzino A and Biferale L 2018 Inferring flow parameters and turbulent configuration with physics-informed data assimilation and spectral nudging Phys. Rev. Fluids 3104604
[8] Farhat A, Johnston H, Jolly M and Titi E S 2018 Assimilation of nearly turbulent Rayleigh-Bénard flow through vorticity or local circulation measurements: a computational study J. Sci. Comput. 77 1519-33
[9] Farhat A, Lunasin E and Titi E S 2017 Continuous data assimilation for a 2D Bénard convection system through horizontal velocity measurements alone J. Nonlinear Sci. 27 1065-87
[10] Foias C, Jolly M S, Manley O P and Rosaosa R 2002 Statistical estimates for the Navier-Stokes equations and the Kraichnan theory of 2D fully developed turbulence J. Stat. Phys. 108 591-645
[11] Foias C, Manley O and Temam R 1987 Attractors for the Bénard problem: existence and physical bounds on their fractal dimension Nonlinear Anal. Theory Methods Appl. 11 939-67
[12] Gesho M, Olson E and Titi E S 2016 A computational study of a data assimilation algorithm for the two-dimensional Navier-Stokes equations Commun. Comput. Phys. 19 1094-110
[13] Goluskin D, Johnston H, Flirrl G R and Spiegel E A 2014 Convectively driven shear and decreased heat flux J. Fluid Mech. 759 360-85
[14] Ibdah H A, Mondaini C F and Titi E S 2020 Fully discrete numerical schemes of a data assimilation algorithm: uniform-in-time error estimates IMA J. Numer. Anal. 40 2584-2625
[15] Jones D A and Titi E S 1992 Determining finite volume elements for the 2D Navier-Stokes equations Physica D 60 165-74
[16] Mondaini C F and Titi E S 2018 Uniform-in-time error estimates for the postprocessing Galerkin method applied to a data assimilation algorithm SIAM J. Numer. Anal. 56 78-110
[17] Stewart C A, Welch V, Plale B, Fox G, Pierce M and Sterling T 2017 Indiana University Pervasive Technology Institute (https://doi.org/10.5967/K8G44NGB)
[18] Temam R 1997 Infinite-dimensional Dynamical Systems in Mechanics and Physics (Applied Mathematical Sciences vol 68) 2nd edn (New York: Springer)
[19] Whitehead J P and Doering C R 2011 Ultimate state of two-dimensional Rayleigh-Bénard convection between free-slip fixed-temperature boundaries Phys. Rev. Lett. 106244501

