

1 **The emergence of shallow easterly jets within QBO westerlies**

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ABSTRACT

13 A configuration of an idealized general circulation model has been obtained
14 in which a deep, stratospheric, equatorial, westerly jet is established that is
15 spontaneously and quasi-periodically disrupted by shallow easterly jets. Sim-
16 ilar to the disruption of the quasibiennial oscillation (QBO) observed in early
17 2016, meridional fluxes of wave activity are found to play a central role. The
18 possible relevance of two feedback mechanisms to these disruptions are con-
19 sidered. The first involves the secondary circulation produced in the shear
20 zones on the upper and lower flanks of the easterly jet. This is found to play
21 a role in maintaining the aspect ratio of the emerging easterly jet. The second
22 involves the organization of the eddy fluxes by the mean flow: the presence of
23 a weak easterly anomaly within a tall, tropical, westerly jet is demonstrated
24 to produce enhanced and highly focused wave-activity fluxes which reinforce
25 and strengthen the easterly anomalies. The eddies appear to be organized by
26 the formation of strong potential vorticity gradients on the subtropical flanks
27 of the easterly anomaly. Similar wave activity and potential vorticity struc-
28 tures are found in the ERA Interim reanalysis for the observed QBO disrup-
29 tion, indicating this second feedback was active then.

30 **1. Introduction**

31 The quasibiennial oscillation (QBO) is the dominant pattern of variability in the tropical lower
32 stratosphere, characterized by alternating descent of easterly and westerly zonal jets with a period
33 of roughly 28 months. In early 2016, a shallow, broad easterly jet emerged in the tropical lower
34 stratosphere, disrupting the QBO by splitting a descending westerly jet roughly in half. This
35 occurrence was unprecedented in more than 50 years of observations (Newman et al. 2016; Osprey
36 et al. 2016).

37 We report here on a series of integrations with an idealized general circulation model (a ‘dry dy-
38 namical core’) which produce events that share dynamical features with the observed disruption.
39 These integrations were originally performed to study extratropical processes, not the tropical cir-
40 culations that resulted. They, along with the further dynamical analysis they prompted, nonetheless
41 suggest some valuable insights into the observed disruption.

42 In these integrations, the model forms a deep, narrow westerly jet, confined to within 15° degrees
43 of the equator, which is quasi-periodically disrupted by shallow, broad easterly jets (Fig. 1a).
44 These form spontaneously just above the base of the westerly jet, then migrate upwards. Figure 1a
45 also shows the meridional component of the wave-induced forcing in the model simulation. In
46 contrast to the standard picture of QBO dynamics, and as is thought to be the case in the observed
47 disruption (Osprey et al. 2016), these meridional fluxes play a central role in the emergence of the
48 easterly jets.

49 One feature common to the observed event and the disruptions in the idealized model is the
50 shallow vertical length scale of the easterly jets. In the idealized model, the tropical upwelling
51 is substantially modified within the easterly jets—enough so that there is net downwelling in the
52 westerly shear zone on their upper flanks. This ‘secondary circulation’ is well-known to produce

53 asymmetries between the descent of easterly and westerly phases of the QBO (Reed 1964; Plumb
54 and Bell 1982; Dunkerton 1991). Indeed Wallace (1967) concluded that this causes westerly shears
55 to propagate downward more rapidly than easterly shears and that this “causes easterly regimes
56 to decrease in vertical extent as they move downwards.” We consider here, through explicit cal-
57 culations with a one-dimensional advective model, whether this secondary circulation could be
58 determining the vertical scale of the easterly jets. This leads to the identification of a threshold
59 forcing strength, above which the advective effects of secondary circulation play an important role
60 in determining the aspect ratio of the jet.

61 Further consideration of Fig. 1 supports the idea of a threshold, which, once passed, triggers a
62 feedback process which then leads to the full development of the easterly jet. The time interval
63 between successive disruptions seen in Fig. 1a is variable, ranging from 3500 to nearly 5000
64 days. However, the evolution of the easterly jet over the period of a few hundred days around the
65 time of the reversal of the winds is quite similar from event to event. This timescale is still long
66 compared to typical timescales of fluctuations in the equatorial wave driving (Fig. 1b), suggesting
67 that the evolution of the jet during this period is not determined by a single extreme wave driving
68 event. Moreover once the jet is established there is consistently enhanced wave driving focused
69 on the easterly jet, suggesting the waves are being systematically organized by the mean flow. We
70 demonstrate explicitly that this is indeed the case in the idealized model, and present evidence that
71 a similar feedback was active during the observed event. This process is distinct from the threshold
72 described in the previous paragraph.

73 The leading order influence of extratropical wave fluxes on QBO winds during both the disrup-
74 tion in early 2016 and disruptions in the dry-dynamical core is interesting in light of early research
75 on the mechanisms behind the QBO. Many researchers (e.g. Wallace 1967) sought to explain the
76 QBO by assuming that the dominant wave-driving is due to horizontal eddy momentum fluxes (of

77 extratropical origin). But convincing model experiments by Wallace and Holton (1968) showed
78 that this mechanism was simply not viable without also assuming this wave-driving also moves
79 downward. This led the way to the Lindzen and Holton (1968) formulation of a model in which
80 the dominant wave forcing came from within the tropics and naturally moved downward with the
81 QBO winds, thereby producing a realistic QBO.

82 The outline of this paper is as follows. Section 2 describes in detail the base configuration of
83 the dry dynamical core. Section 3 discusses the phenomenology of the disruptions in this config-
84 uration, focusing on the structure of the wave driving and of the secondary circulation during the
85 disruptions. Section 4 considers the role of the secondary circulation in setting the structure and
86 evolution of the easterly jet, by considering the response of a one-dimensional advective model
87 to an imposed force. Section 5 demonstrates the feedback between the waves and the mean flow
88 in the dry-dynamical core through two additional sets of integrations. The first set considers the
89 response of the tall westerly equatorial jet to a fixed, imposed forcing of various strengths. For suf-
90 ficiently strong forcing the wave forcing becomes highly organized by the mean flow, amplifying
91 the imposed force. The second set considers the response of the waves to a fixed equatorial zonal
92 wind structure, clarifying the structure of the feedback. Section 6 then discusses the observed
93 event as captured by the ERA-Interim reanalysis (Dee et al. 2011) in light of these results. The
94 structure of anomalous Eliassen-Palm (E-P) fluxes are found to closely resemble those associated
95 with the wave, mean-flow feedback identified in the idealized model, suggesting the same feed-
96 back is relevant for the observed event. Finally, conclusions are given in Section 7, with discussion
97 in particular of the implication of these results for efforts to model and forecast the QBO.

98 2. Model and data

99 *a. Model configuration*

100 The dry dynamical core used is a version of the Reading Intermediate General Circulation Model
101 (IGCM) which solves the hydrostatic primitive equations following Hoskins and Simmons (1975).
102 All integrations are performed using the ‘jagged’ triangular truncation T42 (see Hoskins and Sim-
103 mons 1975) on $N = 60$ hybrid pressure levels spanning from the surface to a log-pressure height of
104 $z_T = 60$ km. The angular-momentum conserving vertical discretization of Simmons and Burridge
105 (1981) is used; this is not a standard feature of the IGCM. The hybrid half-levels are specified by

$$\eta_{i+1/2} = \exp \left\{ -\frac{z_T}{H} \left(\frac{i}{N} \right)^\xi \right\}, \quad i = 0, \dots, N \quad \xi = 1.2 \quad H = 7 \text{ km}, \quad (1)$$

106 and the pressure is specified following Laprise and Girard (1990, their eq (5.1)) as

$$p(\eta) = A(\eta)p_0 + B(\eta)p_s \quad (2)$$

$$A(\eta) = \eta - B(\eta) \quad (3)$$

$$B(\eta) = \left(\frac{\eta - \eta_T}{1 - \eta_T} \right)^r, \quad (4)$$

107 with $\eta_T = \eta_{N+1/2}$ and $r = 1.5$. The full levels are given by $\eta_i = (\eta_{i+1/2} + \eta_{i-1/2})/2$. This grid
108 has a vertical resolution of about 1 km in the lower stratosphere and a horizontal resolution of 4.3°
109 or 480 km.

110 Explicit sixth-order horizontal hyper-diffusion is used to avoid build-up of enstrophy at small
111 scales. The coefficient is set to $5.27 \times 10^{26} \text{ m}^6 \text{ s}^{-1}$, corresponding to a damping timescale of 0.25
112 days for the highest resolved wave numbers. A Robert time filter with parameter 0.02 is also used.

113 The diabatic processes are specified following Polvani and Kushner (2002), which produces an
114 extratropical circulation analogous to a perpetual winter configuration, including a stratospheric

115 polar vortex, taken to be in the Northern Hemisphere in the present work. The parameter γ , which
116 determines the strength of the polar vortex, is set to 1 K km^{-1} .

117 A quasi-stationary wave field is produced by specifying a Gaussian surface topography

$$h_s = h_0 \exp \left\{ - \left(\frac{\phi - \phi_h}{\Delta\phi_h} \right)^2 - \left(\frac{\lambda}{\Delta\lambda_h} \right)^2 \right\}, \quad (5)$$

118 centered in the Northern Hemisphere at $\phi_h = 45^\circ \text{ N}$ with $\Delta\phi_h = \Delta\lambda_h = 15^\circ$. The height h_0 of the
119 mountain is 3 km.

120 Other parameters (including the surface and sponge layer friction) are set identically following
121 Polvani and Kushner (2002), with the exception of κ (0.286 is used here instead of $2/7$) and the
122 hemispheric asymmetry parameter ε , set here to 5 K.

123 As has been found by other authors, the structure of tropical variability in such configurations
124 is sensitive to subtle details, e.g., to the choice of dynamical core (Yao and Jablonowski 2015).
125 This is found to be the case here as well; for instance varying ε by just a few degrees is enough to
126 substantially change the character of the easterly jets. We proceed for now assuming that this sen-
127 sitivity does not imply that the processes involved in the disruption events themselves are similarly
128 sensitive, and return to this question briefly in the conclusions.

129 The base run has been integrated for 25000 days, with instantaneous output every 6 hours. A
130 brief description of the behavior was given in Section 1. Further detailed description and inter-
131 pretation is given in Section 3. Further integrations of the model, with changes in configuration
132 to examine interaction between waves and mean flow, are described in Section 5; these produced
133 instantaneous output on a daily basis. All quantities shown are based on daily averages of the 6h
134 output in the case of the base run or the daily instantaneous output of the further integrations.

135 *b. Reanalysis data*

136 We make use of six-hourly model level data output on a 1 degree grid from the ERA Interim
 137 reanalysis (Dee et al. 2011). Quantities shown on pressure levels are interpolated first to the
 138 pressure levels closest to the hybridized model levels.

139 *c. Derived fields*

140 The forcing of the mean flow by the waves is diagnosed using the Transformed Eulerian Mean
 141 framework (Andrews et al. 1987). The wave forcing is quantified by the Eliassen-Palm (E-P)
 142 flux, while the meridional circulation is estimated by the residual velocities and streamfunction, as
 143 defined on log-pressure coordinates following Andrews et al. (1987). The meridional gradient of
 144 quasi-geostrophic potential vorticity is computed on pressure levels also following Andrews et al.
 145 (1987) as

$$\frac{\partial_\phi \bar{q}}{a} = \frac{2\Omega}{a} \cos \phi - \partial_\phi \left[\frac{\partial_\phi (\cos \phi \bar{u})}{a^2 \cos \phi} \right] - \frac{1}{\rho_0} \partial_z \left(\rho_0 \frac{f^2}{N^2} \partial_z \bar{u} \right). \quad (6)$$

146 *d. Time filtering*

147 Fields are in some cases smoothed in time by convolving time series with an exponential filter
 148 of time scale τ of the form

$$f(t; \tau) = e^{-|t|/\tau}. \quad (7)$$

149 The wave forcing is smoothed using a causal version of this filter;

$$f_c(t; \tau) = e^{-t/\tau} \text{ if } t \geq 0, 0 \text{ otherwise.} \quad (8)$$

150 This is motivated by the fact that only wave forcing that precedes a given time contributes to the
 151 structure of the circulation at that time. In all cases a finite number of weights are used; this is

152 chosen to be large enough that the results are not sensitive to further increases. The choice of
153 time-scale τ is given in the figure captions.

154 **3. Phenomenology of the disruptions**

155 We present in this section further quantitative details of the tropical circulation obtained in the
156 base run, emphasizing the dynamics of the disruptions. Tropical upwelling in this model con-
157 figuration is of the order of $5 \times 10^{-5} \text{ m s}^{-1}$ at 30 hPa, or about 1 km every 230 days. This is
158 substantially weaker than estimates of observed tropical upwelling which is roughly $3 \times 10^{-4} \text{ m}$
159 s^{-1} at 70 hPa, or about 1 km every 40 days. Nonetheless the upwelling plays an important role in
160 the overall structure of the equatorial winds. The westerly equatorial jet arises from the ascent of
161 air through a region of momentum flux convergence; this is primarily due to vertically propagating
162 waves (not shown) but includes a weak contribution from horizontal momentum fluxes (seen in
163 Fig. 1b). The vertically propagating waves are likely Kelvin waves forced non-linearly through
164 extratropical variability given the absence of any convection (parameterized or otherwise). The
165 momentum flux convergence between 40 hPa and 20 hPa is weak and winds are therefore approx-
166 imately uniform with height. Above 20 hPa there is a further region of positive momentum flux
167 convergence, again arising both from horizontally and vertically propagating waves. The easterly
168 jets form within the layer of uniform winds after an extended period of variable but systemati-
169 cally easterly forcing, then migrate upwards, with weak westerlies then being restored by upward
170 advection from the westerly shear layer at 50 hPa.

171 The structure of the jet and the meridional component of the E-P flux in the meridional plane is
172 shown in Fig. 2a-c for three periods: prior, during and after the disruption highlighted in Fig. 1b.
173 The central dates of these periods are indicated by vertical dashed lines in Fig. 1a. The westerly
174 winds at the equator are seen to be part of a relatively narrow jet, generally confined to within 15°

175 of the equator. In all cases there is southwards cross-equatorial E-P flux throughout the tropical
176 stratosphere, consistent with the presence of westerly winds and stationary wave source in the
177 Northern Hemisphere and easterly winds in the Southern Hemisphere, though the convergence of
178 these fluxes is weak. At day 10500 (Fig. 2a), the easterly jet from previous disruption around day
179 8000 has reached the upper stratosphere, and the westerly jet that is reforming below does not
180 yet show a strong second shear zone above 30 hPa. The cross-equatorial meridional E-P fluxes
181 are somewhat stronger towards the base of the jet. By day 13500 (Fig. 2b) the top westerly jet
182 has reached nearly 1 hPa. Although the easterly jet has not yet emerged, the jet has narrowed
183 significantly just below 20 hPa. The cross-equatorial fluxes within 10° of the equator are strongest
184 at this level. The winds at 20 hPa reverse at about day 13650, and by day 14000 (Fig. 2c) the
185 easterly jet has fully formed. In contrast to the tall, narrow westerly jet, the easterly jet is shallow
186 and broad. The shear zones above and below the easterly jet are stronger than the shear zone
187 at 50 hPa, reaching magnitudes greater than 0.01 s^{-1} (shear zones associated with observed QBO
188 approach but rarely exceed this value). Remarkably, the equatorward fluxes are stronger at the level
189 of the easterly jet than they are through the westerly winds above and below, despite the presence
190 of a zero wind line. The same feature can be seen near 5 hPa at the level of the easterly jet around
191 day 10500 (Fig. 2a). The fluxes that are focused on the easterly jet are strongly absorbed by the
192 jet in contrast to the fluxes through the westerly jets that cross the equator relatively unchanged. It
193 is worth noting that most of the patterns of divergence and convergence seen in Fig. 1b arise from
194 quite subtle features in these cross equatorial fluxes.

195 Figures 2d-f show the anomalous residual mass stream function (defined with respect to the
196 time average over days 5000 to 25000). This highlights the presence of secondary circulation cells
197 with vertical convergence and meridional outflow over the equator at the level of the easterly jet,
198 with return flow broadly centered on the westerly jet. The circulations can be understood through

199 the well-established arguments that have previously been applied to the QBO (e.g. Plumb and Bell
200 1982), that they are maintained by the radiative damping of temperature anomalies associated with
201 vertical shear at the equator, implying relative descent in westerly shear zones and relative ascent
202 in easterly shear zones. Here the structure of the circulations are consistent with a tendency to
203 make westerly jets tall and narrow and easterly jets shallow and broad.

204 The evolution of equatorial winds, the full E-P flux divergence, and the vertical velocity com-
205 posited over five disruption events are shown in further detail in Fig. 3. The central date of the
206 disruptions are defined by the date at which the zonal wind first turns easterly at 20 hPa. By 500
207 days prior to the disruption, the net wave driving (Fig. 3a) is weak but systematically easterly over
208 the layer of uniform westerly winds. The wave forcing is dominated by meridional fluxes but its
209 structure is modified by vertical fluxes as can be inferred by comparing with earlier figures. The
210 layer of easterly forcing is quite shallow, and roughly commensurate with the depth of the east-
211 erly jet that emerges at the central date. The beginning of a well-defined easterly anomaly arises
212 in the composite 500 days prior to the central date, centered somewhat below the level at which
213 the winds first reverse. The composite wave driving strengthens somewhat over these 500 days,
214 though the wave driving in individual events is still quite intermittent as can be seen in Fig. 1b.
215 Around the central date the wave driving strengthens substantially as the easterly jet strengthens
216 and begins to migrate upwards, consistent with the focusing seen in Fig. 2c.

217 Prior to the reversal, there is ascent throughout the depth of the tropical stratosphere (Fig. 3b).
218 As the westerly shear strengthens, the secondary circulation first counter acts then ultimately over-
219 whelms the background upwelling, resulting in net downwelling on the upper flank of the easterly
220 jet. Conversely the ascent strengthens in the easterly shear zone. Consistent with the discussion
221 of the stream function anomalies shown in Fig. 2, the secondary circulation acts in the vertical to
222 confine the easterly jet and extend the westerly jet.

223 We have deliberately avoided presenting the momentum budget during this period, in part to
 224 avoid a lengthy digression on the technical details, and in part because closing the budget accu-
 225 rately in this region is quite difficult given the delicate balances and short vertical length scales
 226 relative to the model grid spacing. It has been confirmed, however, that the easterly wave forcing
 227 associated with the meridional E-P flux convergence is the dominant easterly force and is more
 228 than sufficient to explain the net acceleration over the period shown in Fig. 3.

229 **4. Response of a vertical advection model**

230 To better understand this phenomenology, we consider first the role of advection by the sec-
 231 ondary circulation in the response of the westerly jet to an applied force. The basic model we will
 232 use is a one-dimensional model of the vertical profile of equatorial zonal mean zonal wind $u(z, t)$,
 233 considering explicitly the role of vertical advection

$$\partial_t u + (w_0 - \Gamma \partial_z u) \partial_z u = \mathcal{F}. \quad (9)$$

234 We include an applied force \mathcal{F} to represent the easterly wave force over the 500 days or so prior to
 235 the onset of the easterly winds; we will focus mostly on the case where \mathcal{F} is negative and constant.
 236 We return to the substantial enhancement of the wave forcing after this period in the next section.

237 The zonal wind is advected by the vertical wind that consists of a constant background upwelling
 238 $w_0 > 0$ modulated by the secondary circulation produced by radiative relaxation ($-\Gamma \partial_z u$). This
 239 sensitivity of the vertical winds to the vertical shear and its role in the descent of QBO winds in
 240 the meridional plane is discussed by Dunkerton (1991). We restrict our attention here to a single
 241 (spatial) dimension for simplicity, though this has an apparent price: (9) does not conserve total
 242 momentum $\int u dz$. This is in fact consistent with considering this to be a model of the equatorial
 243 region under some simple assumptions as justified in Appendix A; the lack of conservation can

244 be associated with an implied meridional transport and is in fact be a useful feature of (9) as will
 245 become apparent. The essential mechanisms discussed here have also been confirmed in a zonally
 246 symmetric model of the meridional plane (not shown).

247 After Dunkerton (1991, his section 4), Γ can be related to other known parameters. Consider a
 248 local temperature anomaly T' with meridional length scale L over the equator

$$T' = T_0 \left(1 - \frac{y^2}{2L^2}\right). \quad (10)$$

249 Assuming thermal wind balance $\beta y \partial_z u = -R \partial_y T' / H$, and that the relaxational radiative heating
 250 is in quasi-steady balance with the adiabatic heating $N^2 w' = -\alpha R T' / H$, then at the equator $\Gamma =$
 251 $\alpha \beta L^2 / N^2$. Here α is the radiative relaxation rate, β is the meridional gradient of the Coriolis
 252 parameter at the Equator, N is the buoyancy frequency, R is the dry gas constant, and H is a
 253 density scale height.

254 *a. Steady-state response*

255 Returning to (9), it is useful to consider first the steady-state solution to a fixed imposed forcing
 256 $\mathcal{F} = f(z)$. There are two solutions for the shear as a result of the quadratic non-linearity

$$\partial_z u = \frac{w_0}{2\Gamma} \left(1 \pm \sqrt{1 - \frac{4\Gamma f(z)}{w_0^2}}\right). \quad (11)$$

257 If the secondary circulation is weak compared to the background upwelling, the appropriate solu-
 258 tion is the negative root. For small values of the non-dimensional forcing $\Gamma f(z) / w_0^2$, this solution
 259 can be written

$$\partial_z u = \frac{f}{w_0} + \frac{\Gamma f^2}{w_0^3} + O\left(\left(\frac{\Gamma f(z)}{w_0^2}\right)^2\right) \quad (12)$$

$$u(z) \approx u(z_0) + \int_0^z \frac{f(z')}{w_0} + \frac{\Gamma f(z')^2}{w_0^3} dz', \quad (13)$$

260 where (13) follows if the origin is taken to be below a localized forcing. For a single-signed
261 forcing, the largest response is above the forcing region where the ascending parcels have been
262 subject to the largest time-integrated force. In the presence of stronger upwelling, parcels will be
263 subject to the forcing for a shorter time, and therefore the net zonal wind response will be weaker.
264 Steady state is achieved by advecting the anomalous momentum upwards away from the region of
265 the forcing.

266 The secondary circulation introduces an asymmetry between westerly and easterly forces. The
267 ascent of parcels through an easterly force increases, shortening their residence time and weaken-
268 ing the wind response. In this case (11) remains valid, though for large forcings the shear depends
269 on the square root of the forcing (instead of linearly in the case with $\Gamma = 0$). In contrast, the ascent
270 of parcels through a westerly force will slow, lengthening their residence time within the forcing
271 region and resulting in an amplified wind response at the top of the jet. If the forcing exceeds the
272 threshold $w_0^2/4\Gamma$ at some height z , the steady-state solution (11) is no longer valid.

273 *b. Response to a switch on forcing*

274 More direct insight comes from analysis of the transient problem in which u is initially zero.
275 We consider again a localized force, with vertical length scale $D = z_f$, but assume in this subsec-
276 tion that it is abruptly switched on then held fixed. On timescales short relative to the advective
277 timescale $T = z_f/w_0$, easterly shear will develop where the easterly forcing amplitude increases
278 with height, and westerly shear where the easterly forcing amplitude decreases with height. The
279 ascent of parcels within the westerly shear zone will slow, and for sufficiently strong forcing, the
280 induced secondary circulation will produce net downwelling. It is shown in Appendix B that if the

281 easterly forcing is stronger than

$$F_c = \frac{w_0^2}{4\Gamma} = \frac{w_0^2 N^2}{4\alpha\beta L^2}, \quad (14)$$

282 this process leads in (9) to the formation of a localized easterly jet with a discontinuity in the shear
283 at the jet maximum. From the discussion of the steady-state solution in the previous section, one
284 might assume that the arresting of parcel ascent would lead to the build up of easterly momentum
285 within the forcing region; that this does not occur is a result of the meridional transport implied
286 by (9).

287 It is worth noting that this threshold does not depend on the vertical length scale of the force;
288 though for fixed f_c , w_0 , and Γ , the timescale on which this localized maximum emerges, and the
289 magnitude of the associated wind anomalies, do.

290 This is illustrated in Fig. 4, which shows numerical solutions to (9) for a forcing

$$f(z) = -\frac{f_0}{z_f^2}(2z_f - z)z \quad \text{if } 0 < z < 2z_f; \quad 0 \text{ otherwise}, \quad (15)$$

291 with three values of f_0 . The flow has been non-dimensionalized using the advective timescale T ,
292 the vertical scale D of the forcing, and a velocity scale $U = w_0 z_f / 2\Gamma$. The last can be thought of as
293 the wind anomaly associated with a shear layer strong enough for the secondary circulation to be
294 comparable to the background upwelling, with a factor of 2 included for analytical convenience.
295 The solution is determined by the single non-dimensional parameter $F = f_0 / F_c$. Weak vertical
296 diffusion has been added to keep the solutions regular, but it has been verified that the character of
297 the solutions is very weakly sensitive to the value chosen.

298 Figure 4a shows the solution for $F = 0.5$. The shading shows the secondary vertical winds,
299 normalized by the background upwelling w_0 . The upwelling is only slightly enhanced through the
300 forcing layer where forcing produces easterly shear, and the region of transient westerly shear is
301 advected away as the easterly anomaly spreads upwards.

302 Figure 4b shows the case $F = 2$. In this case the secondary circulation is of the same order
303 as the background upwelling, although no net downwelling is produced (values remain below 1).
304 Nonetheless, the vertical convergence is still sufficient to form a localized easterly jet (with a
305 discontinuity in the shear in the inviscid case, as indicated by the characteristics—see Appendix
306 B for discussion). The jet maximum forms within the forcing layer but above its midpoint then
307 migrates upwards until it reaches the top of the forcing layer, upon which the easterly winds again
308 spread upwards.

309 For stronger values of the forcing (Fig. 4c), net downwelling is produced over a narrow region,
310 as occurs in the composite (Fig. 3b). The jet maximum forms earlier and closer to the midpoint of
311 the forcing layer, and persists within the forcing region for a longer period of time. The magnitude
312 of the westerly shear above the jet core is stronger than the easterly shear below, also consistent
313 with Fig. 3.

314 Despite the fact that parcels are being advected towards the center of the jet from above and
315 below, implying they can remain in the forcing region indefinitely, the easterly winds strengthen
316 only moderately. The convergence of the vertical velocities implies a meridional divergence, and
317 thus that the easterly momentum is being transported off the equator, consistent with the structure
318 of the secondary circulation and shallow, broad aspect ratio of the easterly jets seen in Fig. 2.

319 This simple advective model provides the following predictions. Firstly, it suggests a threshold,
320 $F = 1$, above which an imposed easterly forcing will produce an isolated easterly jet within the
321 forcing layer that spreads meridionally (at least transiently), as opposed to an easterly anomaly
322 that spreads upwards with the largest response above the layer of the forcing. It can be shown that
323 this threshold applies essentially unchanged to an applied force of any given vertical structure, and
324 can be generalized to the case where there is shear in the initial profile; these arguments are given
325 in Appendix B. Secondly, for all values of F , the maximum wind response is above the center of

326 the forcing layer, suggesting that the forcing relevant to the disruption lies below the level at which
327 the jets form.

328 While the secondary circulation is essential for determining the aspect ratio of the jet, the vertical
329 scale is determined by the imposed forcing. This is consistent with the structure of the wave driving
330 shown in Fig. 3a and will be discussed in the context of the observed disruption in Section 6.

331 **5. A positive wave, mean-flow feedback**

332 *a. Sensitivity to the stationary wave field*

333 With these insights from the one-dimensional advective model in hand, we return now to the
334 disruptions in the dry dynamical core. It will prove useful to have a basic state in which the
335 tall westerly jet is not spontaneously disrupted by the internal dynamics of the model; easterly
336 forcings of a given geometry can then be externally imposed to test the behaviour expected from
337 the previous section. This, fortuitously, can be achieved by reducing and ultimately eliminating
338 the surface topography in the Northern Hemisphere, though we note that there is still a substantial
339 extratropical planetary-scale wave field even in the absence of the surface topography, forced by
340 non-linear effects (Scinocca and Haynes 1998).

341 Figure 5 shows panels equivalent to Fig. 1a for four additional runs with the height h_0 of the
342 surface topography (cf. 5) reduced to 1500 m, 1000 m, 500 m and finally 0 m. The number of
343 disruptions in each successive run is reduced from (respectively) four, three, one, to finally zero
344 disruptions within the 25000 day integration after the initial transient period. The period between
345 disruptions remains highly variable; in both the $h_0 = 1500$ m and $h_0 = 1000$ m runs, there are
346 disruptions that occur within 4000 days of each other, comparable to the shortest interval between
347 disruptions seen in the base run. The maximum acceleration attained when the winds do reverse

348 from westerly to easterly is very similar across the reduced topography runs and the base run.
 349 These features again suggest that while there is a strongly stochastic aspect to the initiation of
 350 the disruptions that depends on h_0 , the development of the easterly jet itself is controlled by a
 351 deterministic feedback that does not. The presence of such a feedback will shortly be confirmed.

352 *b. Imposed forcing*

353 The predictions of the previous sections can now be tested directly by spinning off a further set
 354 of 20-member ensembles from the $h_0 = 0$ m integration in which a zonally symmetric force G is
 355 imposed. After some trial and error it was found that considering an ensemble, and separating the
 356 starting dates by 1000 days was necessary to avoid artifacts due to low frequency variability in the
 357 westerly shear zone near 50 hPa. The imposed force is chosen to resemble the composite structure
 358 of the wave driving seen prior to the onset of the easterly forcing in Fig. 3a, and has the structure

$$G(\phi, z) = f_0 \exp \left\{ -\frac{\phi^2}{2\Delta\phi_f} \right\} \min \left\{ 1 - \left(\frac{z - z_f}{\Delta z_f} \right)^2, 0 \right\} \quad (16)$$

359 with $\Delta\phi_f = 10^\circ$, $z_f = 27.1$ km, and $\Delta z_f = 1.2$ km. It is switched on immediately at the onset of
 360 the run. Three ensembles are considered, *f8*, *f15*, and *f30*, with respective values for f_0 of 8×10^{-3}
 361 $\text{m s}^{-1} \text{d}^{-1}$, $15 \times 10^{-3} \text{m s}^{-1} \text{d}^{-1}$, and $30 \times 10^{-3} \text{m s}^{-1} \text{d}^{-1}$. The composite easterly force (3a) lies
 362 somewhere between the case *f15* and *f30*.

363 Before discussing the responses, it is worth estimating from (14) the magnitude of the threshold
 364 force F_c . The equatorial value of β and the radiative damping rate are specified externally in the
 365 dry dynamical core; β is $2.3 \times 10^{-11} \text{m}^{-1} \text{s}^{-1}$, and α is $2.9 \times 10^{-7} \text{s}^{-1}$. Assuming the length
 366 scale of the temperature response will be that of the forcing, L is about 7.8×10^5 m. The buoyancy
 367 frequency in this region is essentially determined by the imposed radiative equilibrium tempera-
 368 ture, and is very close to $2.1 \times 10^{-2} \text{s}^{-1}$. This gives a value of Γ of close to 0.01 m. The tropical

369 upwelling in the run with no topography is somewhat weaker than that in the base run; at the levels
370 of the imposed force a value of $3.0 \times 10^{-5} \text{ m s}^{-1}$ is a reasonable estimate. The critical forcing
371 F_c is then about $2.0 \times 10^{-3} \text{ m s}^{-1} \text{ d}^{-1}$. Each of the imposed forcings considered is therefore well
372 within the strong forcing regime, so in the absence of significant eddy responses, we expect the
373 vertical scale of the jet response to match that of the imposed forcing. For comparison with the
374 advective model, the advective timescale T is about 600 days, and the characteristic shear U/z_f is
375 about 0.0015 s^{-1} .

376 The anomalous zonal mean zonal wind, wave forcing, and residual vertical velocities, computed
377 with respect to the corresponding period in the unperturbed, no-topography run, are shown in
378 Fig. 6 for the three ensembles. The two weaker cases, *f8* and *f15*, produce a weak easterly anomaly
379 centered on the level of the imposed forcing with a comparable vertical length scale. The wind
380 response is roughly linear in the strength of the forcing, reaching by the end of the 1000 day
381 integration about 1 m s^{-1} in *f8*, and about 3 m s^{-1} in *f15*. The response of the eddy forcing in the
382 model at the level of the imposed force is weak in both cases. In contrast, by about 500 days into
383 *f30*, the eddies are reorganized to produce a strong easterly forcing, amplifying the effects of the
384 imposed force. As a result, by the end of the 1000 day integration the easterly anomaly reaches
385 about 20 m s^{-1} , comparable to the anomaly associated with the disruptions in the base run (cf.
386 Fig.3). Note that the contour interval for the zonal winds is different for *f30* than for *f15* and *f8*,
387 and that, on the color scale used for the eddy forcing, the imposed forcing of *f30* would only just
388 be visible. The fact that a substantial eddy response is seen in *f30* but not in *f8* or *f15* suggests the
389 existence of a threshold value for the wind response for this eddy response; this is consistent with
390 the easterly wave driving in Fig. 3a amplifying only once the wind anomaly has reached $6\text{-}8 \text{ m s}^{-1}$
391 (it is plausible that this feedback could ultimately arise in *f8* or *f15* if the integrations were carried
392 on for sufficiently long, this has not been explored).

393 The scaling discussed above suggests that in each case the secondary circulation should substan-
394 tially perturb the background upwelling. In the strongest forcing case $f30$ this can be seen (Fig. 6f);
395 the strength of the circulation relative to the shear is consistent with the estimated value of Γ . The
396 secondary circulation response in Figs. 6b,d is also apparent, but is subject to considerable noise.

397 The essential features of the response described by the one-dimensional advective model are
398 therefore confirmed in the dry dynamical core. Moreover, these experiments verify the presence
399 of a dynamical feedback in which the eddy forcing is reorganized to strongly amplify the applied
400 force if the latter is sufficiently strong. Unlike the initiation of the disruptions (which do not occur
401 in the $h_0 = 0$ case), this feedback is active even in the absence of surface topography.

402 *c. Nudged jet structures*

403 To further explore the nature of this dynamical feedback, additional integrations are performed
404 in which the equatorial zonal mean zonal winds are relaxed, or ‘nudged,’ towards a specified
405 profile, allowing the extratropics and the eddies to evolve freely. This approach has been used to
406 artificially produce QBO winds in comprehensive models (e.g. Giorgetta et al. 1999; Marsh et al.
407 2013).

408 Three configurations are considered; a reference case, and two cases with perturbed profiles. In
409 each case the model is integrated for 25000 days, and averages are computed from days 2000 to
410 25000. In the reference run wj , the winds are relaxed towards a tall equatorial westerly jet, with
411 fixed meridional curvature throughout the depth of the stratosphere. Surface topography remains
412 absent ($h_0 = 0$).

413 The nudging is imposed as a linear relaxation of the form $-\kappa(u - u_n)$, imposed on the zonal
 414 mean component of the zonal winds

$$\kappa(\phi, z) = \kappa_0 Z(z; z_n^b, z_n^t, \Delta z_n) \exp \left\{ 1 - \frac{1}{1 - (\phi/\Delta\phi_n)^6} \right\} \text{ if } |\phi| < \Delta\phi_n, 0 \text{ otherwise} \quad (17)$$

$$u_n(\phi, z) = U_0 Z(z; z_u^b, z_u^t, \Delta z_u) \exp \left\{ -\frac{\phi^2}{2\Delta\phi_u^2} \right\} \quad (18)$$

$$Z(z; z_b, z_t, \Delta z) = \frac{1}{2} \left(\tanh \left(\frac{z - z_b}{\Delta z} \right) - \tanh \left(\frac{z - z_t}{\Delta z} \right) \right). \quad (19)$$

415 The overall timescale of the nudging is $\kappa_0^{-1} = 1$ d, and parameters dictating the shape of the
 416 nudging region are $z_n^b = 18$ km, $z_n^t = 50$ km, $\Delta z_n = 2$ km, and $\Delta\phi_n = 20^\circ$. The reference jet has a
 417 maximum speed of $U_0 = 20$ m s⁻¹, and the shape of the jet is determined by $z_u^b = 20$ km, $z_u^t = 50$
 418 km, $\Delta z_u = 2$ km, $\Delta\phi_u = 10^\circ$.

419 The resulting winds are shown in Fig. 7a, along with the meridional component of the E-P flux.
 420 Despite the lack of surface topography, the westerly winds in the Northern Hemisphere allow
 421 waves to propagate upwards into the stratosphere, then equatorward. The equatorial fluxes are
 422 therefore southwards, and are relatively constant with height within the equatorial westerly jet.
 423 The westerly jet is associated with enhanced meridional PV gradients along the equator (Fig. 7d).

424 The wind profile is then perturbed in two further cases *e1* and *e2* by introducing a shallow
 425 easterly anomaly, centred near 20 hPa (the height at which the jets emerge in the free running
 426 model). This is done by replacing the vertical profile Z of the reference jet by

$$\tilde{Z} = Z \left(1 - \delta \exp \left\{ -\frac{(z - z_e)^2}{2\Delta z_e^2} \right\} \right), \quad (20)$$

427 where $z_e = 30$ km and $\Delta z_e = 2$ km. The anomaly in *e1* is half the amplitude of the westerly jet (δ
 428 = 0.5) and so the winds remain westerly at all heights, while the anomaly in *e2* is strong enough
 429 to generate an easterly anomaly ($\delta = 1.5$).

430 Consistent with Fig. 6e, a wind anomaly of 8 m s^{-1} is sufficient to produce a substantial reorga-
431 nization of the wave fluxes. Figure 7b shows the winds in $e1$, as well as the anomalous meridional
432 E-P fluxes relative to the reference run. The cross equatorial flux is enhanced by about 40% in a
433 shallow layer centered at the level of the easterly anomaly. Most of this additional flux is absorbed
434 at the equator. This structure closely resembles that seen in Fig. 2. The wave fluxes in $e2$ are
435 modified through much of the extratropics. There is a strong local enhancement of flux focused
436 on the easterly jet from both hemispheres (Fig. 7c); these anomalies connect to the upper flanks of
437 the tropospheric sub-tropical jet indicating that the easterly anomaly has a substantially non-local
438 effect on the eddies. There are corresponding anomalies in the vertical fluxes as well (not shown).

439 Given the presence of the broad layer of cross-equatorial fluxes present in the base run, it is
440 plausible that the presence of an easterly anomaly within the westerly jet can act as a favourable
441 place for wave breaking and the absorption of easterly momentum. The weaker winds imply
442 slower group velocities and thus, for a given flux, larger wave activities which may be more subject
443 to breaking or damping. This mechanism has been invoked in a barotropic context to argue that
444 Rossby waves incident on an easterly anomaly larger than one-fifth of the value of the initial
445 westerly flow would ultimately lead to a wind reversal (Fyfe and Held 1990). The one-fifth value
446 is roughly consistent with the simulations shown in Fig. 6. However, Rossby waves with phase
447 speeds that would be expected to break on the equatorial anomaly should be unable to propagate
448 through the much weaker winds in the subtropics (O’Sullivan 1997). Moreover, these arguments
449 do not explain why the fluxes are enhanced throughout a broad region of the subtropics, remote
450 from the region of the imposed anomaly.

451 The meridional gradients of quasi-geostrophic potential vorticity, shown in Figs. 7d-f, remain
452 positive throughout the domain and therefore do not suggest that these additional fluxes are gen-
453 erated by barotropic or baroclinic instability. Instead, the region of enhanced PV gradients on the

454 subtropical flanks of the easterly anomaly may be acting as kind of ‘lightning rod’, promoting
455 wave propagation from the upper troposphere and focusing waves towards the developing easterly
456 anomaly which would, in the absence of such a structure, propagate more diffusely.

457 The results and discussion of the previous sections suggest that the easterly disruptions in the
458 dry dynamical core are produced in two stages. In the first stage, meridionally propagating eddies
459 produce an initial shallow easterly anomaly. The time required for this is subject to considerable
460 fluctuations, leading to the variable period between disruptions. The second stage begins once
461 this easterly anomaly becomes sufficiently strong, at which point the wave, mean flow feedback
462 just described sets in. E-P fluxes are enhanced and focused on the developing easterly jet, while
463 the secondary circulation associated with the westerly shear on the upper flank of the easterly jet
464 maintains the shallow, broad aspect ratio of the jet. This feedback process saturates at some point,
465 perhaps when the easterly anomalies become too strong to admit further wave driving.

466 **6. Relevance to the observed event**

467 We now consider to what extent the dynamics of the disruption observed in boreal winter of
468 2015-16 can be understood to follow the two-stage development just described. We make use of
469 the ERA Interim reanalysis for this purpose, but note that many of the relevant dynamical fields in
470 the deep tropics are only weakly constrained by observations (Abalos et al. 2015; Kawatani et al.
471 2016).

472 Figure 8a shows time series of the equatorial zonal mean zonal wind and the meridional diver-
473 gence of the meridional component of the E-P flux for the period of November 2015 through to
474 the end of March 2016. At the beginning of the period the westerly winds extend from near the
475 tropopause up to about 10 hPa with weak vertical shear in a layer from 70 to 40 hPa. From the end
476 of November through to mid-January, a sequence of large-amplitude easterly eddy forcing events

477 occur, centered roughly at the 80 hPa level but extending up to about 30 hPa. In late November, a
478 shallow easterly anomaly begins to form just about the 50 hPa level; by mid-January this shallow
479 region is centered somewhat below 40 hPa and is about 4 m s^{-1} weaker than the westerly winds at
480 30 hPa. From mid-January through to the end of February, there are a sequence of further large-
481 amplitude eddy-forcing events, now more clearly centered at the level of the easterly anomalies,
482 and by the beginning of March, net easterly winds have emerged.

483 The residual mean vertical velocity shown in Fig. 8b is perhaps weakly modulated by the shear
484 zones associated with the emerging easterly jet (e.g. lighter red contours near 30 hPa and darker
485 contours near 55 hPa from late February on), but in contrast to the dry dynamical core (cf. Fig. 3b),
486 the secondary circulation is much weaker relative to the background upwelling.

487 The vertical structure of the meridional component of the resolved wave driving shown in Fig. 8a
488 does not obviously match the vertical structure of the easterly jet which emerges. This is also true
489 of the net (vertical and meridional) resolved wave driving (not shown). However, the results of
490 the one-dimensional advective model suggest that the scale of the easterly jet is determined by
491 the scale of the forcing. As mentioned above, given the relatively weak observational constraints
492 in the tropics, particularly for such a derived quantity, one possible reason for this mismatch is
493 that the resolved wave driving is not correctly captured by the reanalysis; another is the presence
494 of unresolved wave driving. As a rough means of determining which aspects of the forcing are
495 likely to be relevant to the developing jet, a highly idealized ‘back-trajectory’ has been overlaid
496 on Fig. 8a, terminating at 40 hPa on 1 March 2016. The heavy black line and shaded envelope
497 corresponds to an upwelling velocity of $3 \pm 1 \times 10^{-4} \text{ m s}^{-1}$, suggesting that the wave forcing in
498 early winter at 50 or 60 hPa is most relevant. Wave driving much below that level is unlikely to be
499 so; since the winds do not change much in the westerly shear zone near 70 hPa, this forcing from

500 the meridional fluxes is likely compensated by forcing from both resolved and unresolved vertical
501 fluxes.

502 The scaling presented in Section 4 provides a useful framework for understanding the similarities
503 and differences between the observed event and the idealized dry dynamical core. Although the
504 disruption occurs somewhat higher in the dry dynamical core (closer to 20 hPa than to 40 hPa), the
505 vertical length scale of the easterly jets are quite comparable, with a half-width of about 2 km. The
506 background upwelling of about $3 \times 10^{-4} \text{ m s}^{-1}$, however, is nearly an order of magnitude stronger
507 in ERA Interim than in the dry dynamical core. This corresponds to an advective timescale T of
508 about 60 days.

509 The sensitivity parameter, Γ , is more difficult to estimate for the real atmosphere than for the
510 dry dynamical core, not least because the radiative timescale depends on the vertical scale of the
511 associated temperature anomaly (Fels 1982). Estimates of this timescale vary (Mlynczak et al.
512 1999; Randel et al. 2002; Hitchcock et al. 2010), but given the vertical length scales the relevant
513 timescale is likely of the order 10-30 days, corresponding to a value of α about twice that imposed
514 in the dry dynamical core. The meridional length scale on the other hand is somewhat smaller (as
515 shown below), so that assuming a value of Γ unchanged from that estimate for the dry dynamical
516 core is reasonable. Since the threshold forcing F_c depends quadratically on the background
517 upwelling (14), it is far larger for the real atmosphere than for the dry dynamical core. The as-
518 sumptions just outlined give a value of roughly $5 \times 10^{-1} \text{ m s}^{-1} \text{ d}^{-1}$, larger than all but the peak
519 values of wave forcing shown in Fig. 8a. This suggests that the observed disruption is in the weak
520 forcing regime, despite the considerably stronger wave forcing relative to the dry dynamical core
521 (Fig. 1b). This is consistent with the relatively weak anomalies to the upwelling in the reanalysis
522 data.

523 The meridional structure of the zonal wind and meridional E-P flux anomalies (relative to cli-
524 matological values) are shown for three periods in Fig. 9a-c. During the period from Novem-
525 ber through mid-January, there are anomalously strong North-to-South cross-equatorial fluxes
526 throughout most of the depth of the westerly QBO jet; however, the fluxes are strongest at the
527 base of the jet, consistent with the episodes of convergence seen in Fig. 8a. The convergence of
528 these fluxes is strongest below the level where the easterly jet ultimately emerges. These elevated
529 fluxes can be traced to the top of the subtropical jet in the Northern Hemisphere. Figure 9d shows
530 strong meridional gradients of PV along the equator during this period consistent with the tall
531 westerly QBO jet, and with Fig. 7d.

532 From mid-January through to the end of February a shallow layer of meridional E-P flux is seen,
533 centred on the 40 hPa level where the jet is emerging (Fig. 8b). The developing easterly anomaly
534 also leads to a weakening of the meridional PV gradients at the equator and a strengthening of the
535 gradients on the Northern subtropical flank of the anomaly (Fig. 9e). Both the E-P flux anomalies
536 and PV gradients closely resemble those seen in the nudged simulation *e1* (Fig. 7b,e), though
537 the flux anomalies are somewhat stronger and the meridional length scale of the jet is somewhat
538 smaller.

539 By March the equatorial wave forcing at the level of the easterly jet shown in Fig. 8a has weak-
540 ened substantially. However there is still a shallow layer of elevated fluxes focused on the level
541 of the easterly jet apparent during this period in Fig. 9c which converges on the northward flank
542 of the emerging jet. There are also regions of elevated PV gradients centered on the subtropical
543 flanks of the easterly jet, similar in structure and magnitude to those seen in *e2* (Fig. 7f). However,
544 the increase in E-P fluxes in the Southern Hemisphere seen in *e2* (Fig. 7c) is not apparent in this
545 period. This may be explained by the deeper region of easterly winds separating the subtropical
546 jet in the troposphere from the developing easterly jet which was not present in the dry dynamical

547 core integrations. The pattern of fluxes more closely resemble those seen in the free-running dry
548 dynamical core integration (e.g. Fig. 2c).

549 On the basis of these comparisons with the dry dynamical core, we argue that the period from
550 November through to mid-January is analogous to the first stage of the disruptions in the ideal-
551 ized model (as discussed at the end of Section 5), before significant zonal wind anomalies have
552 formed. The eddy driving during this period can, from this perspective, be identified as the ‘trig-
553 ger’ for the event, and because of the vertical advection of the induced momentum anomalies, the
554 relevant forcing during this period lies somewhat below the level at which the easterly jet ulti-
555 mately emerges, likely near 50 or 60 hPa. The period from mid-January through March can then
556 be identified with the second stage of the development of the disruption. The similarity in the E-P
557 flux anomalies and meridional PV gradients between the reanalysis and the dry dynamical core
558 integrations suggests that the dynamical feedback demonstrated in the latter through controlled
559 experiments was also active at this point in the observed disruption.

560 This comparison suggests that to understand why the disruption occurred this year for the first
561 time in the observational record we must understand the nature and origin of the wave driving dur-
562 ing the onset period from November through mid-January. Figure 10a shows the profile of equa-
563 torial wave forcing arising from the meridional component of the E-P flux for each 1 November
564 to 15 January period from 1980-1981 through 2015-2016. Since it has been suggested (Newman
565 et al. 2016) that this event may be related to the large amplitude El Niño event of this year, the
566 winters of 1982-1983 and 1997-1998, other years with large-amplitude El Niño events are also
567 highlighted. The wave forcing from 80 hPa to 50 hPa in the winter of 2015-2016 was in fact the
568 strongest easterly forcing in the reanalysis record by a substantial margin.

569 Because the E-P flux anomalies during this initial period appear to be propagating out of the
570 tropospheric subtropical jet (Fig. 9a; this is confirmed by inspection of the vertical fluxes), it is

571 plausible that these initial wave driving events are associated with synoptic scale eddies propagat-
572 ing higher and deeper into the tropics than normal. This could be a result of more westerly winds
573 permitting more fluxes to propagate deep into the tropics. Figures 10b,c explore this possibility,
574 showing the zonal wind profile and meridional E-P flux along the 60 hPa isobar. The zonal mean
575 zonal wind between the equator and 20 hPa was amongst the most westerly in the record, while the
576 meridional E-P flux equatorwards of about 25 N, along with those during the winter of 1997-1998,
577 were substantially stronger than most other years. The fluxes in the deep tropics, however, remain
578 weak compared to the climatological fluxes at higher latitudes, suggesting that what was unusual
579 was not the overall level of wave activity at these levels, but the degree to which this wave activity
580 was able to propagate into the deep tropics, and the degree to which this activity was absorbed at
581 the equator, which is most obviously controlled by the zonal wind profile in the Northern tropics.
582 At the 60 hPa level these winds are most strongly controlled by the QBO itself, but the influence
583 of El Niño, which is associated with a strengthening of the upper flanks of the subtropical jets,
584 becomes more prominent lower in the stratosphere.

585 These results suggest the importance of several factors in leading to this disruption. First, the
586 QBO westerlies need to be sufficiently deep for the wave driving to produce an isolated easterly
587 jet (rather than simply encouraging or discouraging the descent of a shear zone). They also need
588 to have reached the tropopause, so that the associated westerlies are connected to the subtropical
589 jet, permitting extratropical Rossby wave propagation into the deep tropics. The seasonal cycle
590 of the subtropical jet, and the tendency for El Niño events to raise their upper flanks suggests that
591 the initial trigger is more likely to occur during El Niño events in Northern Hemisphere winter.
592 The dynamical feedback may also be stronger in Northern Hemisphere winter due to the presence
593 of stronger stationary waves, though the results of the dry dynamical core suggest that a topo-
594 graphic source is not essential. Finally, a sufficiently strong series of wave forcing events needs

595 to occur to initiate the easterly anomaly. Given that these extratropical waves carry only easterly
596 pseudomomentum, it is unlikely that an analogous westerly disruption could also occur.

597 **7. Conclusions**

598 A dry-dynamical core configuration is described in which a steady-state, tall, equatorial west-
599 erly stratospheric jet is quasi-periodically disrupted by shallow easterly jets. These disruptions
600 resemble in specific ways the disruption of the westerly QBO phase observed in early 2016.

601 Like the observed event, meridionally propagating eddies play a central role in producing the
602 disruption. The easterly jets appear to organize the forcing produced by the eddies, suggesting
603 the presence of a positive dynamical feedback. Further integrations demonstrate that reducing
604 the extratropical topographic source of stationary waves increases the average time between the
605 disruptions until they are eliminated altogether when the topography is removed.

606 Two possible mechanisms for such a feedback have been considered. The first involves the
607 secondary circulation, which in the dry dynamical core is strong enough to overwhelm the back-
608 ground tropical upwelling. The impact of this process on the emerging jet was considered in the
609 context of a one-dimensional advective model, subject to an imposed force. In this context, if an
610 applied easterly force is stronger than a threshold value, the secondary circulation acts to confine
611 the wind response in the vertical, and momentum is instead advected meridionally off the equator.
612 The threshold force, $F_c = w_0^2 N^2 / (4\alpha\beta L^2)$, depends on the background upwelling, static stability,
613 radiative damping rate, the meridional length scale of the forcing, and the meridional gradient
614 of the Coriolis parameter at the equator. For easterly forces weaker than this threshold the wind
615 response is advected upward. However, while this mechanism is likely important for establishing
616 the aspect ratio of the easterly jet and maintaining its shallow vertical scale, it cannot explain the
617 increasingly rapid strengthening of the easterly jet.

618 The second mechanism considered involves a feedback between the mean flow and the wave
619 forcing. This feedback has been demonstrated in two sets of controlled experiments with the dry
620 dynamical core. In the first set, easterly forces of varying strengths were externally imposed,
621 modeled after the resolved wave forcing found prior to the disruptions. For magnitudes weaker
622 than the composited force, a weak easterly anomaly is produced whose structure agrees with that
623 predicted by the advective model. For magnitudes of the same order or stronger, the resolved
624 wave forcing acts to strengthen the easterly jet, producing more than a five-fold amplification of
625 the imposed force. In the second set, the zonal mean equatorial winds were nudged towards a
626 specified profile, allowing the extratropics and the eddies to evolve freely. Consistent with the
627 first set of experiments, imposing a shallow easterly anomaly is found to produce a narrow region
628 of enhanced wave fluxes arising from the top of the tropospheric subtropical jets, and focused
629 on the easterly anomalies. The enhanced fluxes are related to regions of enhanced meridional
630 PV gradients which may be acting as a kind of ‘lightning rod’ for drawing further wave activity
631 towards the easterly jet.

632 These results suggest that the disruptions evolve through two stages. First, an initial series of
633 weaker wave forcing events produces a weak, shallow easterly anomaly. Provided that the anomaly
634 becomes sufficiently strong, the second stage begins when a positive feedback arises. Extratropical
635 E-P fluxes amplify and focus on the developing easterly anomaly, producing the full easterly jet.

636 While one might be concerned about the sensitivity found in the dry dynamical core to, for
637 instance, the hemispherical sensitivity parameter ε (see Section 2), the similarity between the
638 large-scale flow in these integrations and analogous fields from the ERA Interim reanalysis sug-
639 gests the dynamical processes involved in the disruptions are robust. Indeed, evidence for a similar
640 two-stage evolution is found in the observed disruption. At the beginning of November 2015 the
641 westerly phase of the QBO stretched from the tropopause up to about 10 hPa. A series of wave

642 forcing events centered near 70 hPa but extending upwards to 40 hPa occurred from late November
643 through to about mid-January, producing a shallow easterly anomaly. This period can be identified
644 with the first stage. By mid-January the anomaly was nearly half the magnitude of the westerly
645 phase of the QBO, and from mid-January through February a series of further, stronger wave driv-
646 ing events occurred, centered on the 40 hPa level, leading to the full development of the easterly
647 jet. The pattern of E-P fluxes and PV gradients in the meridional plane during this mid-January-
648 February period are quite similar to those obtained in the dry dynamical core, suggesting that the
649 dynamical feedback identified in the dry dynamical core was active during the observed event.
650 This period can thus be identified with the second stage.

651 The eddy feedback cannot explain the wave driving required to initiate the easterly anomaly by
652 mid-January. The advective model suggests that the wave driving just below 40 hPa is relevant
653 for this anomaly, and indeed the wave driving at these levels during this period is found to be
654 stronger than any other year in the ERA Interim reanalysis record. While this is likely to be due
655 to a variety of factors as discussed in the previous section, the zonal winds connecting the QBO
656 westerlies to the upper flanks of the subtropical jets were also amongst the most westerly in the
657 record suggesting that the mean state was at least more conducive to this wave driving. Given
658 the tendency for El Niño to strengthen the winds in this region (e.g. Simpson et al. 2011), it is
659 likely that the strong El Niño event that occurred over the same period played a role in this event.
660 Increasing concentrations of greenhouse gases are also expected to lead to this kind of circulation
661 response (e.g. Shepherd and McLandress 2011), suggesting that such disruptions may become
662 more likely, though the expected strengthening of tropical upwelling may counter act this to some
663 extent.

664 While the wave driving during the onset of the disruption, i.e. during November through mid-
665 January, may well have been statistically unlikely and difficult to forecast, the relevance of a

666 feedback process suggested by the similarity of the observed event to the dry-dynamical core in-
667 tegrations suggests that seasonal forecast models should have some skill in predicting the second
668 stage. Failure to do so may imply a significant deficiency in seasonal forecast models of the sub-
669 tropical winds near the tropopause; indeed, the sensitivity of the dry-dynamical core integrations
670 described above and the fact that such a disruption has not previously been observed and is thus
671 a rare event is consistent with the idea that the occurrence of the disruption is highly sensitive to
672 background conditions. Regardless of whether such disruptions recur or not, the disruption may
673 thus prove to be a sensitive and valuable test of model performance in this critical region of the
674 atmosphere.

675 *Acknowledgments.* PHi acknowledges support from European Research Council ACCI grant
676 Project No. 267760. PHa acknowledges support from the IDEX Chaires d’Attractivité pro-
677 gramme of l’Université Fédérale de Toulouse, Midi-Pyrénées. TBi acknowledges helpful con-
678 versation with Mike Wallace regarding early work on the QBO. This work was partially supported
679 by NASA GNSS Remote Sensing Science Team Grant NNX16AK37G. The National Center for
680 Atmospheric Research (NCAR) is sponsored by the U.S. National Science Foundation (NSF).
681 The source code and data for the dynamical core integrations are available from the corresponding
682 author upon request.

683 APPENDIX A

684 **Momentum budget in 1D advective model**

685 The advective model considered in Section 4 can be justified by considering the zonally symmetric
686 zonal momentum equation in Cartesian coordinates. In flux form, conservation of total momentum

687 is clear

$$\partial_t u + \partial_z(uw) + \partial_y(uv) = \mathcal{F}. \quad (\text{A1})$$

688 Taking a meridional average $\langle \cdot \rangle$ over a narrow region about the equator, then if meridional varia-
689 tions from this average can be neglected in u and w , this leads to

$$\partial_t \langle u \rangle + \langle w \rangle \partial_z \langle u \rangle + \langle u \rangle \partial_z \langle w \rangle + \langle u \rangle \partial_y \langle v \rangle = \mathcal{F}, \quad (\text{A2})$$

690 from which (9) follows after use of the continuity equation. Variations in w with height then imply
691 corresponding meridional transport of mass and momentum out of the equatorial region.

692 APPENDIX B

693 **Threshold behaviour in the advective model**

694 Adopting the scaling discussed in the text, the vertical derivative of (9) is

$$\partial_t u_z + (1 - u_z) \partial_z u_z = \partial_z \mathcal{F}, \quad (\text{B1})$$

695 where all symbols are now their non-dimensional equivalents. (Note the forcing scale is $2F_c$.) This
696 is a linear first order partial differential equation for the vertical shear $u_z = \partial_z u$ that can be solved
697 along characteristics

$$\frac{dz}{ds} = 1 - u_z \quad (\text{B2a})$$

$$\frac{du_z}{ds} = \partial_z \mathcal{F}. \quad (\text{B2b})$$

698 The vertical velocity of these characteristics is not the upwelling velocity (which would be $1 - \frac{u_z}{2}$).

699 For the piecewise quadratic forcing (15), this leads to a second order ordinary differential equation

$$\frac{d^2 u_z}{ds^2} + F u_z = \begin{cases} F & \text{if } 0 < z < 2, \\ 0 & \text{otherwise.} \end{cases} \quad (\text{B3})$$

700 Solutions to (B3) for easterly forces ($F > 0$) are trigonometric, while for westerly forces ($F < 0$)
 701 they are exponential; we consider the former.

702 The switch-on problem considered in the text assumed an initial vertical profile of shear $u_z(t =$
 703 $0, z) = U'(z)$, and no shear at the base of the domain for all time $u_z(t, z = 0) = 0$. The steady-state
 704 solution is determined by characteristics starting at the base of the domain; of interest here is the
 705 behavior of those which start within the forcing region when the force is switched on. For now the
 706 initial shear is taken to vanish ($U' = 0$).

707 The solution along characteristics is

$$u_z(s) = \sqrt{F} (z_0 - 1) \sin \sqrt{F}s + 1 - \cos \sqrt{F}s, \quad s < s_c \quad (\text{B4a})$$

$$z(s) = \frac{1}{\sqrt{F}} \sin \sqrt{F}s + (z_0 - 1) \cos \sqrt{F}s + 1, \quad s < s_c, \quad (\text{B4b})$$

708 where s parameterizes the characteristics, and the fact that $\frac{du_z}{ds}(s = 0) = f'(z_0)$ at the height $z_0 =$
 709 $z(0)$ where the characteristic is initialized has been used. The characteristics leave the forcing
 710 layer when $z(s_c) = 2$, after which the shear (and thus the vertical velocity) remains constant. The
 711 time s_c is given by

$$\sqrt{F}s_c = \arcsin R + \arcsin(1 - z_0)R, \quad R = (F^{-1} + (1 - z_0)^2)^{-\frac{1}{2}}. \quad (\text{B5})$$

712 Within the forcing layer, characteristics which enter the domain after the onset of the forcing will
 713 first be subject to the lower flank of the imposed forcing which strengthens with height, producing
 714 easterly shear. They then accelerate upwards until they pass $z = 1$, after which they are subject to
 715 the upper flank of the imposed forcing which weakens with height, reducing the shear and slowing
 716 their ascent. In steady state (e.g. for $z_0 = 0$), this recovers (11).

717 Consistent with the intuition that the easterly forcing should tend to accelerate the ascent of the
 718 parcels, those trajectories that start at $z_0 = 0$ always reach the top of the forcing layer, no later than
 719 $\pi z_f / 4\sqrt{\Gamma f_0}$ (in dimensional terms) after they enter. Trajectories which begin above the midpoint

720 of the forcing layer, for which $z_0 > 1$, are subject only to the upper flank of the forcing, and do not
 721 always reach the top of the forcing layer.

722 This can be seen, for instance, by considering (B4b) for the set of trajectories which start at
 723 $t = 0$. At the time $s_0 = \frac{\pi}{2\sqrt{F}}$, $z(s_0) = F^{-1/2} + 1$ becomes independent of the initial condition z_0 ,
 724 provided that the trajectories have not yet left the forcing layer. This will be the case for at least
 725 some trajectories if $F > 1$. In this case a cusp forms with easterly shear below and westerly shear
 726 above. The height $z(s_0)$ at which the cusp forms always lies within the forcing layer, above the
 727 midpoint.

728 More general forcing profiles can also be considered. Multiplying together (B2a) and (B2b)
 729 yields

$$\frac{d(u_z - \frac{1}{2}u_z^2)}{ds} = \frac{d\mathcal{F}}{ds} \quad (\text{B6})$$

730 which can be integrated to find

$$\frac{dz}{ds}(s) = \pm \sqrt{(U' - 1)^2 - 2(\mathcal{F}(s) - \mathcal{F}(0))}. \quad (\text{B7})$$

731 Characteristics turn over if they reach a height at which

$$\mathcal{F}(z(s)) - \mathcal{F}(z_0) = \frac{1}{2}(U' - 1)^2. \quad (\text{B8})$$

732 For $U' = 0$, this will necessarily occur for an arbitrary localized easterly force if the maximum
 733 amplitude of \mathcal{F} is greater than 1, or, equivalently, if its dimensional magnitude is greater than
 734 $F_c = w_0^2/4\Gamma$.

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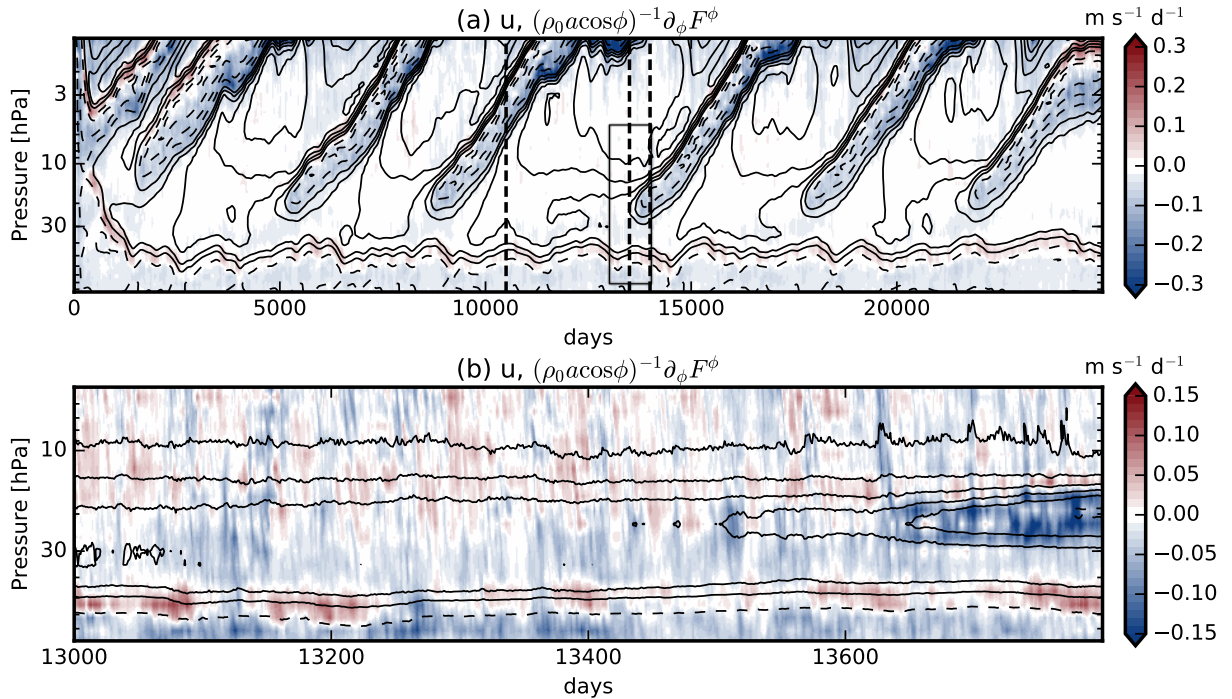
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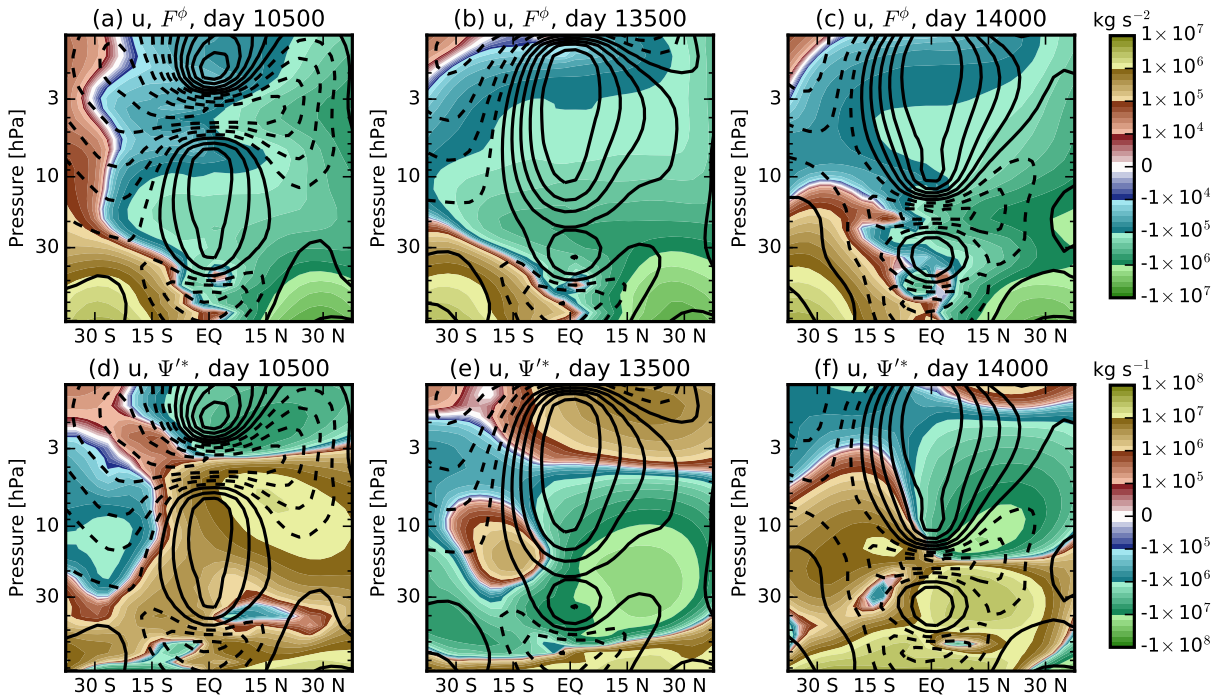
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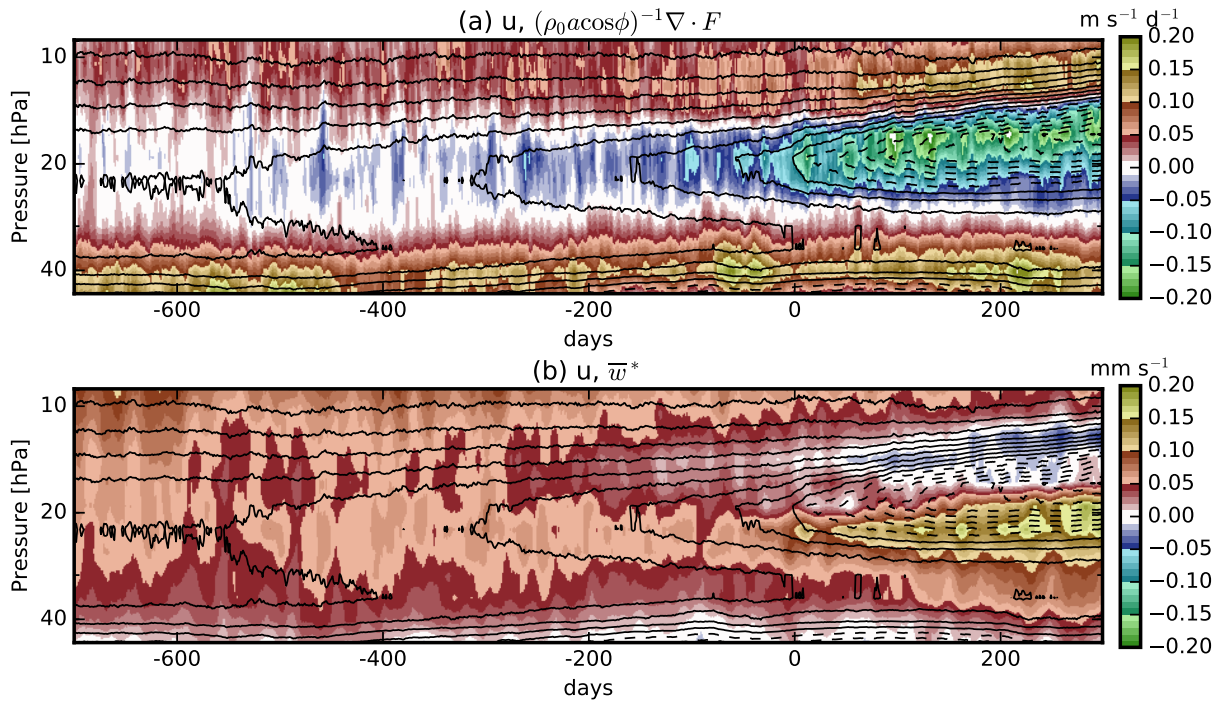
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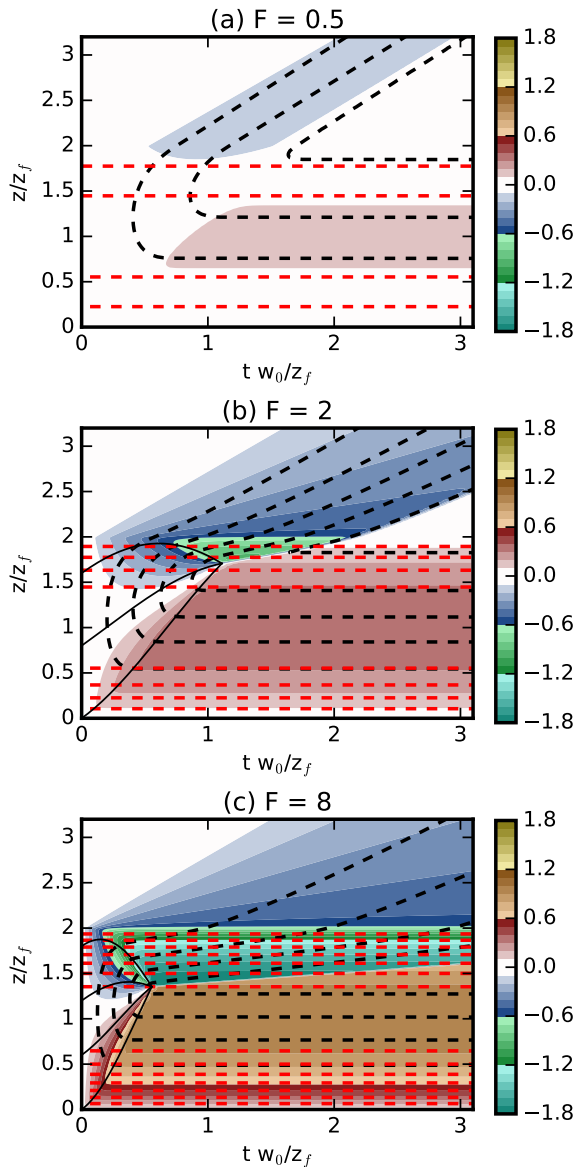
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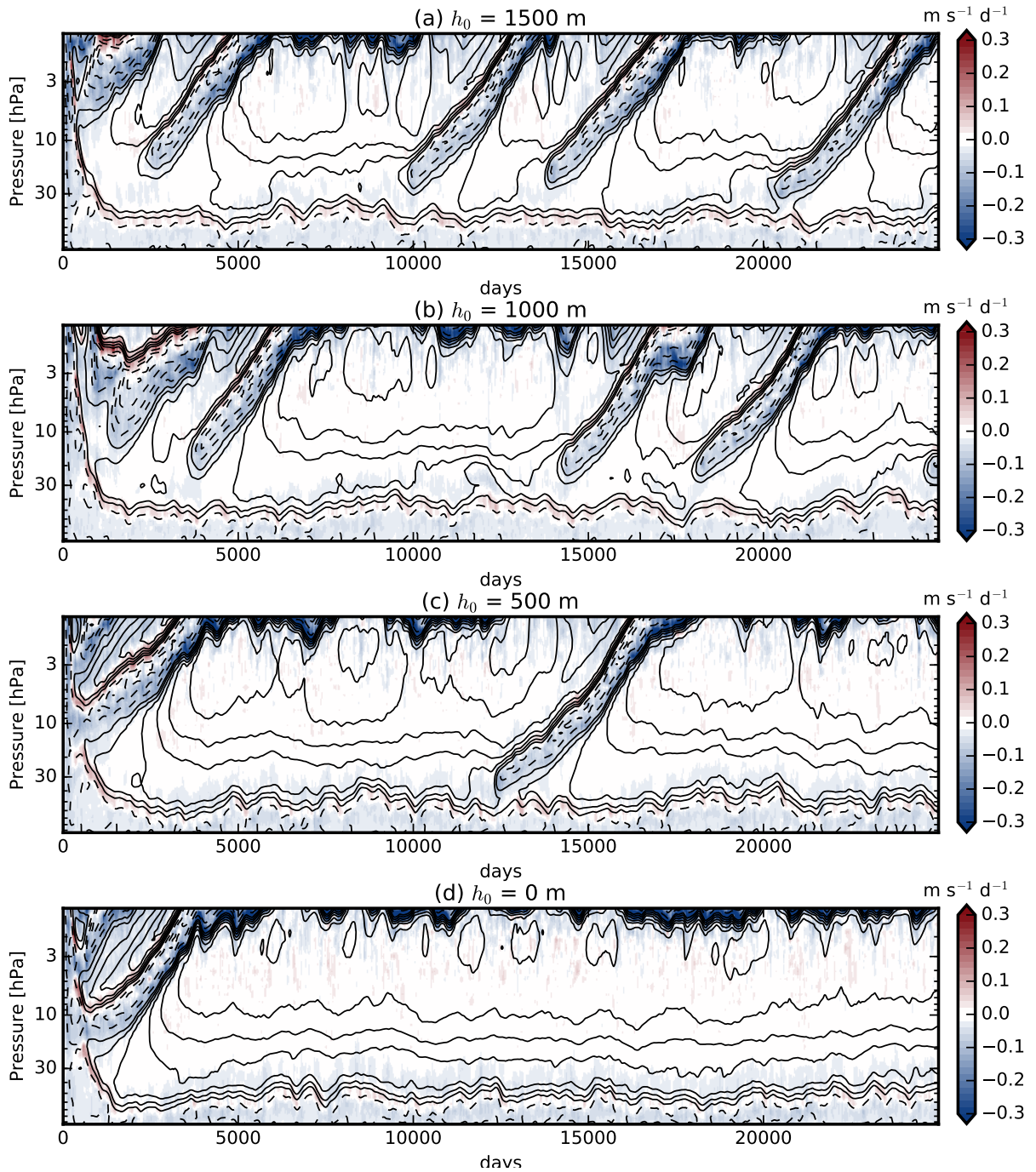
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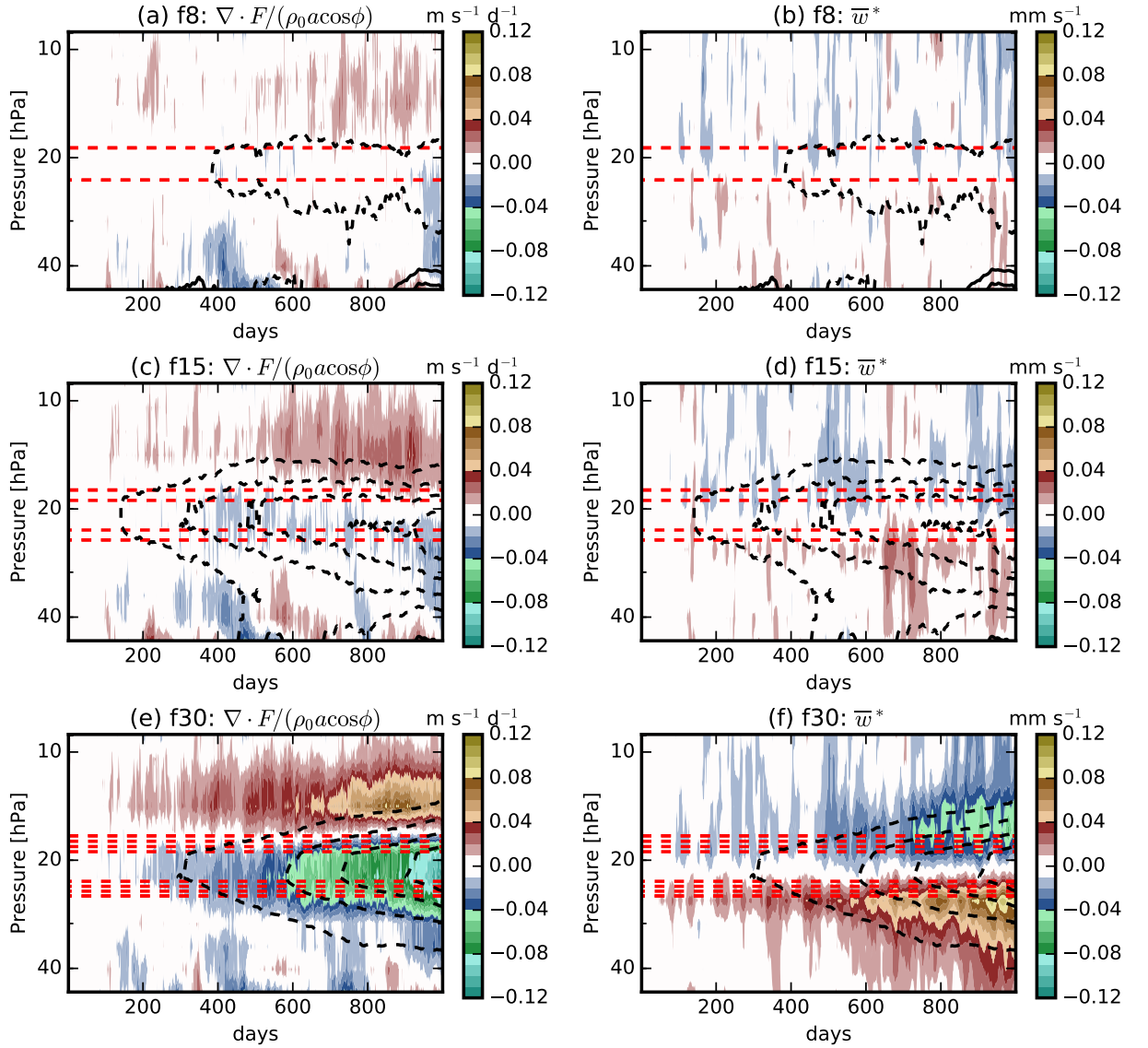
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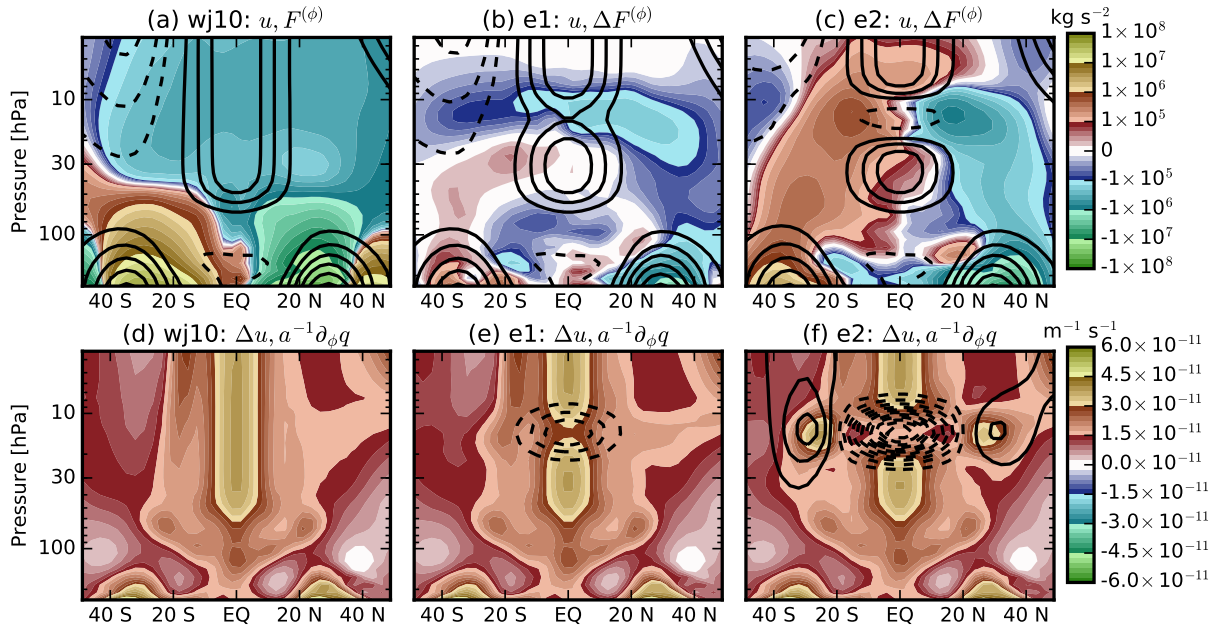
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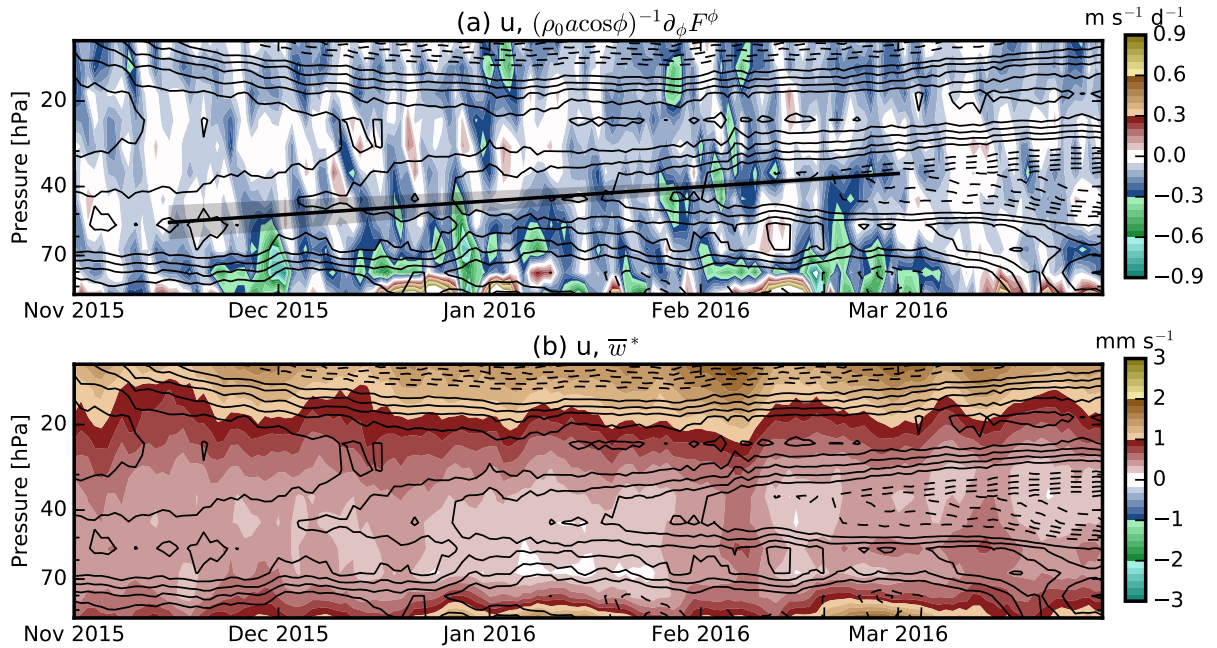
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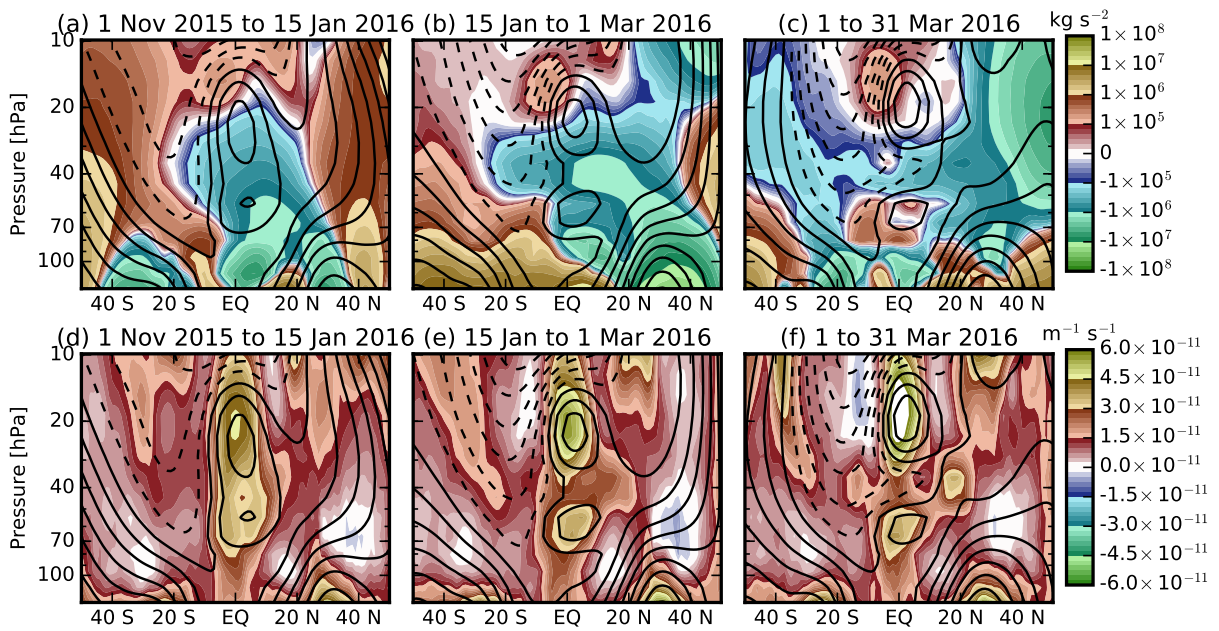
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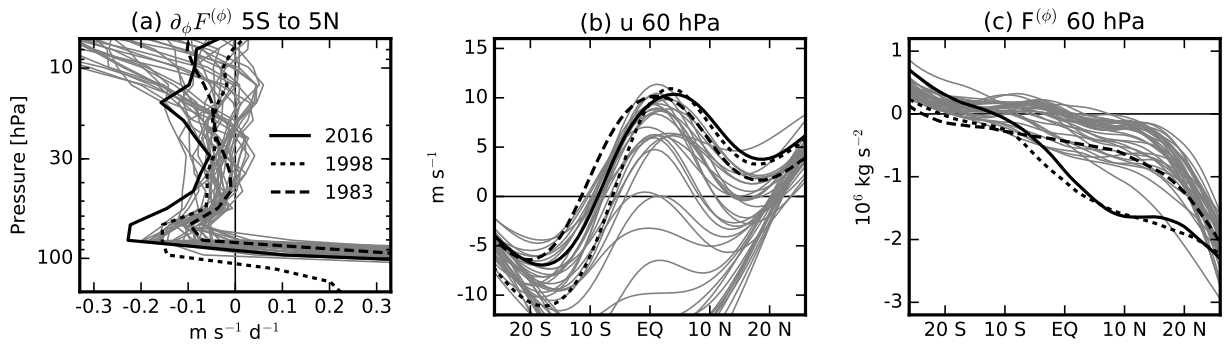
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