

Open-vent volcanoes fuelled by depth-integrated magma degassing

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Abstract

Open-vent, persistently degassing volcanoes—such as Stromboli and Etna (Italy), Villarrica (Chile), Bagana and Manam (Papua New Guinea), Fuego and Pacaya (Guatemala) volcanoes—produce high gas fluxes and infrequent violent strombolian or ‘paroxysmal’ eruptions that erupt very little magma. Here we draw on examples of open-vent volcanic systems to highlight the principal characteristics of their degassing regimes and develop a generic model to explain open-vent degassing in both high and low viscosity magmas and across a range of tectonic settings. Importantly, gas fluxes from open-vent volcanoes are far higher than can be supplied by erupting magma and independent migration of exsolved volatiles is integral to the dynamics of such systems. The composition of volcanic gases emitted from open-vent volcanoes is consistent with its derivation from magma stored over a range of crustal depths that in general requires contributions from both magma decompression (magma ascent and /or convection) and iso- and polybaric second boiling processes. Prolonged crystallisation of water-rich basalts in crustal reservoirs produces a segregated exsolved hydrous volatile phase that may flux through overlying shallow magma reservoirs, modulating heat flux and generating overpressure in the shallow conduit. Small fraction water-rich melts generated in the lower and mid crust may play an important role in advecting volatiles to subvolcanic reservoirs. Excessive gas fluxes at the surface are linked to extensive intrusive magmatic activity and endogenous crustal growth, aided in many cases by extensional tectonics in the crust, which may control the longevity and activity of open-vent volcanoes. There is emerging abundant geophysical evidence for the existence of a segregated exsolved magmatic volatile phase in magma storage regions in the crust. Here we provide a conceptual picture of gas-dominated volcanoes driven by magmatic intrusion and degassing throughout the crust.

Introduction

Open-vent volcanoes are characterised by their persistent outgassing and mildly explosive activity between major eruptions (Andres and Kasgnoc, 1998; Francis et al., 1993; Rose et al., 2013; Vergnolle and Métrich, 2021). Many are well-studied because persistent low-level activity allows access and collection of extended timeseries of monitoring data. Open-vent volcanoes are found in all tectonic settings and are associated with a range of magma compositions and bulk viscosities (some examples—

38 not an exhaustive list—are shown in **Fig. 1** and **Table S1**). Open-vent volcanoes may be active over
39 millennia – for example, Masaya, Nicaragua (Stix, 2007), Stromboli, Italy (Allard et al., 1994), Etna, Italy
40 (Allard, 1997), Villarrica, Chile (Witter et al., 2004), Yasur Volcano, Vanuatu (Métrich et al., 2011) and
41 Erebus, Antarctica (Oppenheimer et al., 2011) – or years to decades, such as Soufrière Hills, Montserrat
42 (Christopher et al., 2010) and Fuego (Lyons et al., 2010) and Santiaguito, Guatemala (Holland et al.,
43 2011).

44
45 Volcanoes that transition from being ‘open-vent’ to ‘closed vent’ over years to decades timescales may be
46 classified as ‘persistently restless’. For example, Telica Volcano, Nicaragua, transitions between a ‘weak
47 seal’ and a ‘destabilised’ state, which may produce phreatomagmatic eruptions (Rodgers et al., 2015;
48 Roman et al., 2019). Long-dormant volcanoes may also convert to open-system behaviour when they
49 reactivate. Reactivation may initiate explosively, as at Santa María volcano in Guatemala (now the
50 location of the open-vent effusion of the Santiaguito flank volcano) (Lamb et al., 2019), or effusively, as
51 at Soufrière Hills volcano, Montserrat (Wadge et al., 2014). Following the initiation of activity in 1995,
52 Soufrière Hills has outgassed continuously for more than 25 years at the time of writing, despite not being
53 in a state of eruption for much of that time (Christopher et al., 2015).

54
55 Eruptions of open-vent volcanoes are typically gas-rich and may be highly hazardous. The nature of the
56 eruptive activity varies with magma composition. Mafic stratovolcanoes exhibiting open-vent behaviour—
57 such as the archetypal Stromboli volcano, Italy—exhibit frequent strombolian eruptions punctuated by
58 large violent explosions, or ‘paroxysms’ (Bertagnini et al., 2011; Métrich et al., 2005; Rosi et al., 2006).
59 Persistent degassing from mafic lava lakes may persist over decades or longer with accompanying
60 strombolian explosions and/or lake overflows or draining events (e.g., Ambrym, Vanuatu; Erta Ale,
61 Ethiopia; Masaya, Nicaragua) (Bouche et al., 2010; Lev et al., 2019; Moussallam et al., 2021). Open
62 system behaviour in more evolved systems is typically accompanied by episodic explosive activity
63 (typically vulcanian or violent strombolian in style depending on the melt composition; Cashman and
64 Sparks, 2013), effusion of viscous lava flows and domes, and/or gas venting episodes (Edmonds and Herd,
65 2007). The over-arching characteristics of open-vent activity in all settings, however, are that the
66 outgassing flux of volatiles far exceeds the volatiles that can be supplied from degassing of erupted
67 magma and that high levels of outgassing from a central vent continues between eruptions (Andres et al.,
68 1991). Open-vent volcanoes may therefore be thought of as predominantly gas emitters, with the magma
69 that is supplying the outgassing accumulating endogenously in the crust beneath (Allard, 1997; Anderson,
70 1975; Francis et al., 1993; Giggenbach, 1992; Giggenbach, 1996). Where in the crust the magma
71 accumulates is, however, an open question.

72
73 It has also been shown that open-vent volcanoes are the most prodigious volcanic outgassers of volatiles
74 into the atmosphere, worldwide (Andres and Kasgnoc, 1998; Carn et al., 2016). Additionally, extensive

75 records of the outgassing fluxes of open-vent volcanoes from many decades of in situ measurements
76 (Arellano et al., 2021; Carn et al., 2016; Carn et al., 2017; Fioletov et al., 2016) show that outgassing
77 *between* eruptions dominates the volcanic gas budget (Allard, 1997; Carn et al., 2016; Carn et al., 2017).
78 Indeed, satellite-based global observations of SO₂ flux confirm that persistent, or passive, degassing
79 accounts for ~90% of the global outgassing sulfur flux from volcanoes (Carn et al., 2016; Carn et al.,
80 2017) and that most of the top 20 volcanic outgassers, as quantified from UV sensors TOMS (Total Ozone
81 Mapping Spectrometer) and OMI (Ozone Mapping Instrument), may be classified as 'open-vent' (**Table**
82 **S1; Fig. 1**) (Carn et al., 2017).

83

84 **Background and aims of this review**

85

86 *Outstanding questions related to outgassing from open-vent volcanoes*

87

88 The most pressing questions surrounding the outgassing of open-vent volcanoes, and the consequent
89 implications for both monitoring and understanding how these volcanoes work, concern the sources and
90 mechanisms of volatile degassing. Volatiles (e.g., H₂O, CO₂) exsolve from magma upon reaching
91 saturation in the silicate melt or by partitioning into a pre-existing exsolved phase (e.g., sulfur, chlorine)
92 (Aiuppa et al., 2008; Candela, 1997; Cashman, 2004; Edmonds and Wallace, 2017; Edmonds and Woods,
93 2018; Métrich and Wallace, 2008; Wallace, 2005) (**Fig. 2**). Volatile degassing from melts occurs during
94 decompression (sometimes called 'first boiling'; **Fig. 2a**); this drives undercooling and crystallisation
95 (Cashman and Blundy, 2000) and, as a result of isobaric cooling and crystallisation, second boiling in
96 magma reservoirs in the crust (**Fig. 2b**). For open-vent volcanoes, where large fluxes of gases are
97 sustained with comparatively little magma erupted (**Table S1**), key questions include (1) the extent to
98 which exsolved volatiles derive from first and/or second boiling, (2) mechanisms of volatile transfer
99 upward through the magmatic system (i.e. as an exsolved magmatic volatile phase (hereafter MVP) or
100 retained within volatile-rich melts, and (3) the effect of the volatile transfer mechanism on resulting
101 volcanic activity (**Fig. 2**).

102

103 *Degassing in the volcanic conduit*

104 A popular model to explain the high observed outgassing fluxes of water, sulfur, CO₂, and halogen species
105 at mafic open-vent volcanoes is bimodal flow driven by convection, whereby buoyant, volatile-rich
106 magma rises up a conduit and degasses; then the denser, gas-free magma sinks over vertical length scales
107 of several kilometres (Kazahaya et al., 1994; Palma et al., 2011b; Shinohara, 2008; Stevenson and Blake,
108 1998) (**Fig. 3a**). Convective flow has been reproduced in both analogue and numerical experiments
109 (Beckett et al., 2011; Cardoso and Woods, 1999; Huppert and Hallworth, 2007; Molina et al., 2012). A
110 simple Poiseuille flow model of buoyancy-driven ascent of magma in a conduit is given by:

111

$$Q_{ascend} = \frac{\pi \Delta \rho_{d-a} g r_a^4}{8 \mu_a}, \quad (1)$$

112
113
114 (Kazahaya et al., 1994) where Q_{ascend} is the volume flux of ascending magma, g is the gravitational
115 constant, $\Delta \rho_{d-a}$ is the density difference between the bubble-rich magma at depth and the shallow
116 degassed magma, r_a is the effective conduit radius for ascending magma, and μ_a is the viscosity of
117 ascending magma. If no magma is erupted, then Q_{ascend} must be balanced by the volume flux of
118 descending magma minus the volume of the volatiles released to the surface (Kazahaya et al., 1994;
119 Stevenson and Blake, 1998). SO_2 fluxes of 10^2 - 10^3 tonnes per day (typical of many of the volcanoes
120 highlighted in **Fig. 1** and **table S1**), for example, require magma fluxes in the conduit of ~ 1 - $10 \text{ m}^3/\text{s}$.
121 Magma flux, in turn, is controlled by the conduit radius (assuming a cylindrical geometry) and the flow
122 velocity, which is a function of magma viscosity and density. If the gas is transported with the magma,
123 maintaining the same gas supply (assuming similar exsolved gas contents) requires magma with a
124 viscosity of 10^8 Pa s to occupy a conduit approximately ten times wider than magma with a viscosity of
125 10^4 Pa s .

126
127 Critically, however, H_2O -rich magmas do not maintain constant viscosities as they ascend because they
128 undergo extensive decompression-induced degassing and consequent crystallisation (Cashman and
129 Blundy, 2000; Lipman et al., 1985; Métrich and Rutherford, 1998). The addition of crystals may increase
130 the magma viscosity by orders of magnitude (Giordano et al., 2014 (Lejeune and Richet, 1995)).
131 Ultimately, it is likely that slowly-rising water-rich magmas will entirely crystallise, as observed in lava
132 domes, and therefore convection is unlikely. In lower viscosity magmas, changing the viscosity contrast
133 between the down- and upwelling liquids can affect the geometry of the exchange flow (Beckett et al.,
134 2014). It has been suggested that magma may overturn at various depths before reaching the surface (e.g.,
135 Masaya, Nicaragua; Aiuppa et al., 2018; Stix, 2007). Regardless of the exact location, however, magma
136 overturn within the shallow system requires that degassed magmas accumulate in subsurface storage
137 regions. At Etna, Italy, for example, there is evidence for endogenous accumulation of degassed magma at
138 a rate of $22.9 \pm 13.7 \times 10^6 \text{ m}^3 \text{ y}^{-1}$ in a storage region between 3 and 10 km beneath the surface (Coppola et
139 al., 2019); whether this magma crystallised *in situ* or degassed higher in the system is not known; a topic
140 debated by a range of authors in the past (Allard, 1997; Steffke et al., 2011).

141
142 Alternative models to explain the outgassing flux from open-vent volcanoes invoke the permeable flow of
143 an exsolved volatile phase through magma in the conduit. In viscous magmas, gas flow is governed by
144 bubble connectivity and the development of permeability. In the absence of crystals, permeability
145 development during decompression of a hydrous melt depends on decompression rate, magma
146 composition (viscosity) and shear (e.g., Giachetti et al., 2019; Hurwitz and Navon, 1994; Lindoo et al.,
147 2015; Okamura et al., 2006; 2008; 2013). Experimentally-determined vesicularity thresholds for

148 permeability development vary from ~30 to 80%, depending on the deformation regime (Okumura et al.,
149 2008). Experimental data suggest that efficient, channelised gas flow may occur at depths of a few
150 kilometres through rhyolite melt containing 5 wt% H₂O when subject to a shear strain >8 (Okumura et al.,
151 2008). Addition of crystals may substantially reduce the percolation threshold for system-scale
152 connectivity during vesiculation and may promote efficient gas loss from conduits even at low gas
153 fractions (Colombier et al., 2020; Collombet et al., 2021; deGraffenreid et al., 2019; Lindoo et al., 2017;
154 **Fig. 3b**). Degassing-induced rheological changes in shallow conduit magma may promote brittle
155 fracturing at the conduit walls, providing transient, highly permeable pathways for gas loss (Gaunt et al.,
156 2014; Gonnermann and Manga, 2003; Rust et al., 2004; Tuffen and Dingwell, 2005) (**Fig. 3c**) and
157 generating low frequency seismicity (Iverson, 2008; Neuberg et al., 2006).

158
159 In crystal-rich magmas, gas may be trapped in pore spaces between crystals, where it may accumulate
160 until the overpressure generated overcomes the local yield strength of the crystal framework (Belien et al.,
161 2010; Oppenheimer et al., 2015); this presents a mechanism by which gases may accumulate in crystal-
162 rich plugs and subsequently trigger strombolian eruptions (Oppenheimer et al., 2020; Suckale et al., 2016;
163 Woitischek et al., 2020) (**Fig. 3b**). Gas 'hold-up' (accumulation of gas within the magma) occurs when
164 gas supply from depth is balanced by gas loss from the system and may be implicated as a triggering
165 mechanism for paroxysmal eruptions more generally. For example, paroxysmal eruptions are often
166 preceded by increases in the height of the magma column which may be caused by gas retention; the
167 resulting lava effusion from either flank (Stromboli) or summit (Fuego) vents may then trigger
168 decompression of the shallow conduit (Calvari et al., 2011; Liu et al., 2020b; Ripepe et al., 2015).
169 Similarly, correlations between lava lake surface elevations and gas flux at Villarrica (Johnson et al.,
170 2016), Erta Ale (Bouche et al., 2010) and Masaya (Aiuppa et al., 2018; Williams-Jones et al., 2003) for
171 example suggest that temporal fluctuations in deep (> 1-2 km) gas supply may be important in modulating
172 surface activity at open-vent volcanoes and in advecting heat to maintain an open state.

173 174 *Degassing throughout the magmatic system*

175 The introduction of gas into shallow (top few km) reservoirs and conduits derived from deeper (>2-3 km)
176 in the system requires a mechanism of deep volatile exsolution. The principal source of that deep MVP is
177 crystallisation and second boiling, which can generate the equivalent of several weight percent for
178 andesite and dacite magmas (**Fig. 2b**). The MVP will initially be CO₂-rich, with increasing water for
179 higher degrees of crystallisation (**Fig. 2b**). Once formed, the MVP can migrate upward and out of the
180 crystal-rich magma reservoir and rise towards the surface (Degruyter et al., 2019; Huber et al., 2010;
181 Parmigiani et al., 2017). Given that the ratio of intrusive to extrusive magmatism is thought to be high in
182 all tectonic settings (from an average of 3:1 to 10:1 or higher in many arc regions); (Crisp, 1984; White et
183 al., 2006) and that plutonic rocks are generally volatile-poor (Bachmann et al., 2007), it follows that the
184 volatiles outgassed persistently by open-vent volcanoes likely have their source, at least in part, in second

185 boiling processes in crustal reservoirs. Evidence for this deep (>2-3 km and perhaps extending into the
186 mid or lower crust in some cases) MVP is provided by the “excess sulfur” emissions accompanying large
187 explosive eruptions of arc volcanoes, which have been explained by sulfur partitioning into substantial
188 accumulations of an exsolved MVP in magma reservoirs (Andres et al., 1991; Rose et al., 1982; Scaillet
189 and Pichavant, 2003; Wallace and Gerlach, 1994; Webster et al., 2011; Zajacz et al., 2012). A deep-
190 derived MVP is also implicated in gas fluxing observed in the volatile systematics of many melt inclusion
191 suites (Blundy et al., 2010; Caricchi et al., 2018; Métrich and Wallace, 2008), as well as in the diffuse
192 degassing of CO₂ along rifts (Foley and Fischer, 2017) and other volcanic centres (Werner et al., 2019).
193 Questions remain as to the MVP source depth and mechanism(s) by which a deep-derived MVP
194 segregates and migrates through crustal magma storage regions.

195
196 At low melt viscosities and low crystal fractions, bubbles may accumulate in foam layers at the roof zones
197 of eruptible melt lenses (Jaupart and Vergnolle, 1989; Vergnolle and Jaupart, 1986) and on foam
198 collapse, bubbly plumes may be generated (Degruyter et al., 2019; Parmigiani et al., 2016) (**Fig. 3d**). At
199 intermediate crystal fractions in more evolved systems, the MVP generated through deep (>2-3 km)
200 crystallisation and periodic influx from mafic recharge may rise buoyantly through crystal-rich mush via
201 viscous elongate fingering channels, which produce high permeability pathways for a deep MVP phase to
202 percolate (**Fig. 3e**) (Degruyter et al., 2019; Parmigiani et al., 2016). MVP accumulation in melt-rich caps
203 or lenses may aid eruption of crystal-poor rhyolites (Bachmann and Bergantz, 2004). At high crystal
204 fractions, the MVP may become trapped and accumulate in pore spaces between crystals; it may escape on
205 ductile or brittle deformation (capillary fracturing) when the crystal framework is disrupted (Belien et al.,
206 2010; Oppenheimer et al., 2015; Parmigiani et al., 2016) (**Fig. 3f**).

207
208 *Aims of this paper*

209 It is clear that a universal paradigm is required that applies to all open-vent volcanoes, of all magma types
210 (high and low viscosity), and which addresses important questions such as how and where the MVP forms
211 and its mode of its transport through the magmatic system. We review outgassing from open-vent
212 volcanoes and lay out the characteristic and generic features common to all settings and all magma
213 compositions. In particular, we examine how new insight into the dynamic nature of crustal magma
214 systems, including conceptual models of volcanic-plutonic systems linked by eruptible melt lenses and
215 unstable volatile-rich fluids (Cashman et al., 2017; Christopher et al., 2015), help us to understand
216 persistent outgassing and gas-rich eruptions from open-vent volcanoes. More specifically, we assess the
217 contribution of unerupted magmas and extensional tectonics to the outgassing fluxes observed at open-
218 vent volcanoes. In considering not only the outgassing characteristics, but also the available evidence for
219 the form and extent of the underlying magmatic system using petrology, geophysics and modeling, we
220 propose a new generic picture for understanding the volatile budget of these volcanoes.

221

222 **Key observations of open-vent volcanic outgassing**

223

224 *Outgassing fluxes from open-vent volcanoes are decoupled from eruptions*

225

226 Recent observations of volcanic outgassing from space have highlighted the number and diversity of
227 open-vent volcanoes that emit the overwhelming bulk of volcanic gases into the atmosphere every year
228 (**Fig. 1; Table S1**). Global satellite-based monitoring of volcanic gas emissions demonstrate
229 unequivocally that >90% of the global outgassing fluxes of sulfur dioxide are produced during ‘passive
230 degassing’ from an open-vent, where no eruption is taking place (Carn et al., 2017; Fioletov et al., 2016;
231 Werner et al., 2019). These open-vent volcanoes erupt magmas of a wide range of compositions and
232 rheological properties (**Table S1**), from low viscosity basalt to highly viscous crystal-rich andesite.
233 Moreover, as our understanding of volcanic outgassing increases, it is becoming ever clearer that ‘excess’
234 volcanic gas (over that which can be supplied by erupting magma) is the norm, rather than the exception
235 (Andres et al., 1991; Francis et al., 1993). Here we review the gas emission systematics from a number of
236 persistently degassing volcanoes with a wide range in magma compositions and rheological properties,
237 eruptive style and setting.

238

239 The flux of sulfur dioxide is commonly used as a proxy for the total volatile flux from a volcano (Aiuppa
240 et al., 2008). In most cases, SO₂ makes up 1-10 mol% of the gas phase from open-vent volcanoes, with the
241 bulk of the gas composed of water and CO₂ in variable proportions. These two major gas species (H₂O
242 and CO₂) are not easily measured, however, owing to their large and variable concentrations in the
243 background atmosphere. SO₂, in contrast, has a distinct and strong absorption in the UV region (Hoff and
244 Millan, 1981) and is not present in the background atmosphere, making it ideal for monitoring volcanoes.

245

246 Etna Volcano (Italy) is an archetypal ‘open-vent’ volcano. It has long been observed that the persistent gas
247 fluxes from Etna are too high to be supplied by the erupting magma (Allard, 1997). SO₂ fluxes between
248 1975 and 1995 varied from <1000 t/d during quiescent degassing to >10,000 t/d during fountaining
249 (Allard, 1997). Since then Etna has continued to outgas at prodigious rates (Andres and Kasgnoc, 1998;
250 Caltabiano et al., 1994; Salerno et al., 2009), with average SO₂ outgassing rates from 2005-2015
251 determined from space-based inventories showing an average rate of 2039 t/d (Carn et al., 2017).
252 Approximately 25-30% of that SO₂ flux can be accounted for by decompression-driven degassing of
253 erupted magma (**Fig. 4**). The high rate of volatile outgassing has been attributed to continuous, convective
254 bimodal flow, whereby alkali basalts ascend to shallow pressures, degas and then sink back down into the
255 edifice (Allard, 1997; Burton et al., 2003; Kazahaya et al., 1994), although there is little definitive
256 geophysical or geochemical evidence to support this.

257

258 Yasur Volcano (Vanuatu) is a persistent and continuous outgasser with small-scale strombolian activity
259 interspersed with larger paroxysms (Kremers et al., 2012; Métrich et al., 2011; Suckale et al., 2016;
260 Woitischek et al., 2020). Anecdotal and historical evidence suggest that continuous degassing has been
261 taking place for several centuries (Métrich et al., 2011). Frequent, strombolian eruptions eject small
262 volumes of crystal-rich trachybasalt generated in shallow reservoirs by ~50% crystallisation of more
263 primitive alkali basalts (Métrich et al., 2011). SO₂ fluxes ranged from 400-700 t/d during field campaigns
264 in 2006, 2010 and 2018, with much of the SO₂ emitted by passive degassing between explosions (Bani
265 and Lardy, 2007; Ilanko et al., 2020; Métrich et al., 2011; Woitischek et al., 2020). Again, the gas budget
266 cannot be accounted for by degassing of erupted magma (**Fig. 4**); instead, the SO₂ flux requires complete
267 degassing of 0.04-0.05 km³ per year of unerupted magma, which is ~50 times that erupted (Métrich et al.,
268 2011; Woitischek et al., 2020). If the recent measurements are extrapolated to the past, > 4 km³ degassed
269 magma has accumulated beneath Yasur over the past 100 years (Métrich et al., 2011).

270
271 Manam, a basaltic stratovolcano in the Western Bismarck arc, is one of the most active volcanoes in
272 Papua New Guinea. Continuous outgassing from two summit craters has been sustained at least over the
273 past few decades (Carn et al., 2017; Liu et al., 2020a). Sporadic strombolian eruptions produce low-level
274 ash plumes and are punctuated by occasional paroxysmal eruptions involving lava fountaining, lava flows,
275 pyroclastic density currents, and high ash plumes; five explosive events between August 2018 and June
276 2019 produced >10 km high eruption plumes. Manam is among the most prolific volcanic outgassers
277 globally, with an average SO₂ flux of 1480 t/d between 2005 and 2015 (Carn et al., 2017); and a 2019
278 campaign measured fluxes ≤ 7660 t/d over several days (Liu et al., 2020a). Assuming an undegassed
279 magmatic sulfur content of ~2000 ppm, this large SO₂ flux requires around 0.33 km³ of magma to degas
280 every year, which is likely to be an order of magnitude, and perhaps two, more than the erupted volume
281 (the erupted volumes have not yet been quantified) (**Fig. 4**).

282
283 Some of the most prolific and /or persistent global outgassers are lava lake volcanoes (Carn et al., 2017),
284 including Nyiragongo and Nyamuragira (Democratic Republic of Congo), Ambrym (Vanuatu), Masaya
285 (Nicaragua), Erebus (Antarctica), Erta Ale (Ethiopia) and Kilauea Volcano (Hawaii, USA). Degassing
286 from the surface of a lava lake takes the form of vigorous bubbling, low fountains, bubble bursting, gas
287 pistoning and overturn and resurfacing phenomena (Allard et al., 2016; Bani et al., 2012; Bouche et al.,
288 2010; Harris et al., 2005; Oppenheimer et al., 2009; Patrick et al., 2016; Swanson et al., 1979), with
289 upwelling and divergence zones providing evidence for rapid lateral magma motion across the lake's
290 surface (Harris, 2008; Harris et al., 2005; Lev et al., 2019; Pering et al., 2019). These observations, as well
291 as the high gas fluxes (a mean of 7356 t/d SO₂ from Ambrym between 2005 and 2015, with peaks
292 reaching >20,000 t/d SO₂) (Bani et al., 2009) and necessity for a continuous heat source to keep the lake
293 above its solidus temperature, have led to prevailing models of bimodal flow in the conduit to supply both
294 sufficient volatiles and heat (Kazahaya et al., 1994; Oppenheimer et al., 2009; Oppenheimer et al., 2004;

295 Palma et al., 2011a). Although analogue experiments can reproduce bimodal flow (Palma et al., 2011a;
296 Witham and Llewelin, 2006), we note that simple gas fluxing through the lava lake may supply sufficient
297 heat and volatiles to satisfy observational requirements. For example, degassing at the surface of the Erta
298 Ale (Ethiopia) lava lake occurs at fixed positions that are inferred to be directly above the conduit (Bouche
299 et al., 2010). Here visual, thermal and acoustic observations suggest that spherical cap bubbles rise to
300 burst at the surface; bubbly wakes that detach from the bubble bottom generate small fountains and hold
301 sufficient heat to ensure that the lava lake does not cool over time. In this scenario, a deep ($> \sim 1$ km)
302 source of gas is required, with no requirement for large-scale vertical bimodal flow. Moreover, the
303 dynamics of bubble behaviour within lava lakes may modulate degassing (Qin et al., 2018).

304
305 Intermediate composition magmas can also feed open-system vents, as illustrated by Soufrière Hills
306 Volcano, Montserrat and Santiaguito Volcano, Guatemala. Soufriere Hills erupts high viscosity (10^9 - 10^{12}
307 Pas) crystal-rich andesite (Melnik and Sparks, 2002). In contrast to the mafic systems, the typical eruptive
308 style is lava dome growth interspersed with episodes of vulcanian activity. SO_2 fluxes here have been
309 sustained since the onset of eruptive activity in 1995 (Christopher et al., 2015; Christopher et al., 2010;
310 Edmonds et al., 2010; Edmonds et al., 2014) and high gas emission rates have continued (to at least 2021)
311 since the cessation of eruptive activity in 2011. SO_2 fluxes have fluctuated between <100 and >2500 t/d
312 throughout the eruption (Christopher et al., 2015; Nicholson et al., 2013), with the highest SO_2 fluxes
313 observed immediately after large dome collapses that exposed the conduit (e.g., July 1998, July 2003)
314 (Herd et al., 2005). SO_2 fluxes were high and sustained during periods of both lava dome growth and
315 prolonged (months-years) periods during which the eruption paused (Christopher et al., 2015; Edmonds et
316 al., 2010). Petrological studies indicate that prior to eruption, sulfur solubility in the rhyolite melt was low
317 (<100 ppm) (Edmonds et al., 2001), consistent with partitioning of sulfur into an exsolved MVP in the
318 shallow storage region beneath the volcano (Clémente et al., 2004; Edmonds et al., 2001; Edmonds et al.,
319 2002). The high bulk viscosity of the magma precludes convective flow as a viable mechanism to supply
320 the outgassing fluxes; sustained degassing during eruptive pauses therefore requires both persistent
321 permeable pathways from depth to the surface and tapping of a substantial pre-segregated reservoir of
322 exsolved volatiles (Christopher et al., 2015).

323
324 Bagana Volcano (Papua New Guinea) has exhibited long-lived and continuous degassing, perhaps over
325 centuries (McCormick et al., 2012; McCormick Kilbride et al., 2019; Wadge et al., 2018). Bagana's
326 edifice is built of crystal-rich andesite lava flows and tephra (53-58 wt% SiO_2) and eruptive activity is
327 characterised by the emplacement of steep-sided lava flows, pyroclastic density currents and ash-rich
328 explosions (Bultitude and Cooke, 1981; Wadge et al., 2018). Observations (predominantly by satellite
329 radar) indicate that eruptive activity is strongly pulsatory, with eruptive periods separated by periods of
330 repose, throughout which strong degassing continues (Wadge et al., 2018). SO_2 fluxes at Bagana were first
331 measured in 1983 and reported at $> 3,000$ t/d (McGonigle et al., 2004). A recent global inventory of

332 volcanic SO₂ emissions measured via satellite-mounted UV sensor (Ozone Mapping Instrument, OMI)
333 reported Bagana's mean SO₂ flux as 1380 t/d for the period 2005-2015 (**Table 1**), placing it 3rd in the
334 global ranking of sustained SO₂ fluxes (Carn et al., 2017). The highest SO₂ fluxes occur during eruptive
335 periods (up to 10,000 t/d) but gas emissions remain high (≤ 2500 t/d) during eruptive pauses (McCormick
336 Kilbride et al., 2019). These high gas emissions cannot be supplied by the erupted magma, which has a
337 time-averaged eruption rate of 1 m³/s (Wadge et al., 2018) (**Fig. 4**). Instead, the observed rates of
338 degassing from 2005-2015 require around 5-6 times the observed magma flux when reasonable water and
339 sulfur melt concentrations for arc magmas are assumed (McCormick Kilbride et al., 2019).

340

341 To summarise, although a model of conduit convection may explain persistent degassing at some
342 volcanoes, it does not supply a universal explanation. In particular, the conduit convection model predicts
343 that high viscosity must be compensated by a larger conduit radius in order to supply gas at similar rate to
344 a lower viscosity system, yet there is no observational evidence for a systematic linear relationship
345 between magma viscosity and conduit radius. This problem is exacerbated by the lower solubility of sulfur
346 in rhyolitic melts (Cl  mente et al., 2004). Additionally, a convective model requires large accumulations
347 of degassed magma in the shallow crust, which poses a substantial space problem for long-lived open-vent
348 systems. Water-rich magmas may completely crystallise during slow ascent, severely inhibiting return
349 flow. A more parsimonious explanation for high gas flux across all volcano types is that gases are
350 supplied from a mixture of shallow (conduit) and deep (> 1-2 km and perhaps as deep as the mid-crust)
351 sources. Importantly, the flux of gases supplied from deeper magma storage regions to the shallow
352 systems has potential to both modulate and trigger eruptive activity and advect heat; this model also
353 allows for degassed magma accumulations to be distributed over a substantial depth range.

354

355 *Gas compositions at open-vent volcanoes are consistent with mixing between deep and shallow degassing*
356 *sources*

357

358 Additional information about how gases are delivered to the surface and from what depth they are sourced
359 comes from measurements of changes in volcanic gas compositions with eruptive activity. There has been
360 immense progress in quantifying the composition of volcanic gas emissions over the past two decades
361 (specifically the relative abundance of H, C, S and Cl species), principally driven by instrumentation
362 development (Aiuppa et al., 2010; Aiuppa et al., 2006; Liu et al., 2020a; Pering et al., 2020; Shinohara et
363 al., 2008). **Fig. 5** shows a compilation of gas composition data from a range of volcanoes, many of which
364 have open vents (not discriminated on the diagram). Volcanic gases are rich in H₂O and CO₂ and arc
365 volcanoes are typically richer in Cl than rift or intraplate volcanoes. Although we do not consider
366 hydrothermal systems here, we note that gases from volcanoes hosting a large hydrothermal system are
367 typically depleted in S and HCl and rich in H₂O and CO₂. Finally, **Fig. 5** shows the large variability in the
368 molar H₂O/CO₂, C/S and S/Cl ratios measured in volcanic gases at the surface.

369

370 Before scrutinising the natural data, it is useful to construct a framework for volcanic gas compositions to
371 understand how gas ratios may evolve during a) decompressional degassing (with some crystallisation)
372 and b) isobaric equilibrium crystallisation in a magma storage region in the crust (second boiling). We use
373 *MagmaSat* (Ghiorso and Gualda, 2015b) to model the solubility of H₂O and CO₂ under different pressure,
374 temperature and oxygen fugacity conditions (as shown in **Fig. 2**). Three examples are considered – Yasur,
375 Stromboli and Soufrière Hills – using appropriate basaltic and andesitic compositions (typical erupted
376 magma compositions for three examples are given in **table S2, supplementary material**). For example, to
377 initialise a decompressional degassing model for Yasur trachybasalt (**Fig. 6a**) we use a water content of 1
378 wt% and a CO₂ content of 0.2 wt%, consistent with petrological studies of melt inclusion compositions
379 (Métrich et al., 2011). Initial melt volatile contents are further modified by crystallisation during magma
380 ascent, which we model using RhyoliteMelts (Ghiorso and Gualda, 2015a). We model chlorine and sulfur
381 exsolution using both a closed and open system partitioning model (see **supplementary material** for
382 details).

383

384 We use a suite of D_{Cl} (fluid-melt partition coefficient for chlorine) values collated from the literature
385 (Kilinc and Burnham, 1972; Lesne et al., 2011; Shinohara, 1994; Tattitch et al., 2021; Webster et al.,
386 1999; Webster et al., 2017). D_{Cl} is low (<10) for basaltic compositions and decreases as pressure decreases
387 (Tattitch et al., 2021), although the solubility behavior of Cl is complex and varies with melt composition
388 (Métrich and Rutherford, 1992; Métrich and Rutherford, 1998; Signorelli and Carroll, 2002), fluid
389 composition (Botcharnikov et al., 2004), temperature, oxygen fugacity and pressure (Botcharnikov et al.,
390 2004; Webster et al., 2009); a review is presented in the **supplementary material**. Some studies have
391 postulated an inverse relationship between D_{Cl} and pressure, i.e. that D_{Cl} decreases with increased pressure
392 (Alletti et al., 2009; Shinohara, 2009); this is explained by the large and negative pressure dependence of
393 NaCl partitioning into a melt and the HCl–NaCl exchange reaction between a silicate melt and an aqueous
394 fluid, which favors HCl in aqueous fluids at lower pressures (Shinohara, 2009). These pressure
395 dependencies cause chlorine to appear as HCl in low pressure (~0.1 MPa) volcanic gases and NaCl in high
396 pressure (~50 MPa) fluids. More work is required, however, to fully understand the implications of Cl
397 speciation on fluid-melt partitioning (Shinohara, 2009).

398

399 We use D_S (fluid-melt partition coefficient for sulfur) values derived from experiments at high pressure
400 and temperature using natural basalt samples from Masaya and Stromboli (Lesne et al., 2011), which
401 range from 1 to 5 at pressures >200 MPa and >100 for pressures <50 MPa (**Fig. 6a**). Lesne et al. (2011)
402 used synthetic samples based on natural Stromboli melts with an initial volatile inventory representing the
403 most volatile-rich melt inclusions from each volcano. For more evolved compositions, we use partition
404 coefficients derived from experiments (Botcharnikov et al., 2004; Botcharnikov et al., 2015; Webster et
405 al., 2011). These indicate that the fluid-melt partition coefficient for sulfur increases with melt

406 differentiation, reaching values of >500 for rhyolitic melts, and likely increases as the melt water content
407 decreases during decompression (Moune et al., 2009). Model results for Yasur, Stromboli and Soufriere
408 Hills are shown in **tables S3, S4, S5**.

409

410 A second set of models (a Yasur example is shown in **Fig. 6b**) simulates isobaric, closed system degassing
411 during crystallisation for four example pressures between 80 and 350 MPa, thus representing magma
412 stored in the crust that undergoes second boiling (details of the model are given in the **supplementary**
413 **material**; results are shown in **tables S6, S7 and S8**). The melt concentrations of H₂O, CO₂, Cl and S,
414 together with the molar fractions of each in the exsolved volatile phase, are shown in **Fig. 6(v) and 6(vi)**
415 as a function of melt fraction. The observed compositions of the volcanic gas at Yasur are shown in **Fig.**
416 **6(iii) and 6(vi)** for comparison (Métrich et al., 2011; Woitischek et al., 2020); and glass compositions are
417 shown in **Fig. 6(ii) and 6(v)** (Métrich et al., 2011). It is important to note that the models shown here
418 incorporate a significant amount of uncertainty; we use them to define the general trends expected for
419 magma degassing under a range of conditions.

420

421 The degassing behaviour of Yasur magmas (**Fig. 6**) shows an exsolved volatile phase that evolves from
422 carbon- and chlorine-rich compositions at high pressures, to sulfur- and water-rich compositions at low
423 pressures, consistent with our understanding of the effect of pressure on solubility and partitioning (**Fig.**
424 **6a, b; table S3**). Exsolved volatile phase C/S ratios attain a maximum (of > 300 for closed system
425 degassing and ~ 10 for open system degassing) at pressures of 100-230 MPa (**Fig. 6(iii)**). Sulfur is
426 preferentially exsolved over Cl at low pressures, leading to a sharp increase in exsolved volatile phase
427 S/Cl ratios and a sharp drop in the S/Cl ratios of melts at $P < \sim 100$ MPa (**Fig. 6(ii, iii)**). The ranges in
428 X_{melt}^S and X_{melt}^{Cl} thus derived match well with ranges of these elements preserved in melt inclusions and
429 matrix glasses from Yasur (**Fig. 6(ii)**) (Métrich et al., 2011). More generally, model predictions are
430 consistent with published measurements of volatile concentrations in melt inclusion and groundmass
431 glasses at Stromboli, Yasur and Etna (Métrich et al., 2011; Métrich et al., 2010; Métrich and Wallace,
432 2008; Spilliaert et al., 2006) and observed in the closed system experiments of Lesne et al. (2011).

433

434 The exsolved volatile phase is expected to have a molar C/S of 10-100 at pressures >100 MPa, decreasing
435 from ~10 at 80 MPa to ~1 at the surface. Volcanic gases at Yasur have a molar C/S of ~2-3 (Métrich et al.,
436 2011; Woitischek et al., 2020), consistent with gases being sourced from integrated, open system
437 degassing of the entire magma column to a pressure of 80 MPa (< 3 km depth). Open system degassing is
438 expected; the basalt has a low viscosity (<1000 Pa s; (Giordano et al., 2008)) and bursting of large bubbles
439 at the surface is the dominant style of activity (Kremers et al., 2012; Woitischek et al., 2020). Volcanic
440 gases have a molar S/Cl ratio of ~ 0.5 to 30 (Woitischek et al., 2020; Métrich et al., 2011; Oppenheimer et
441 al., 2006); this is again consistent with open system degassing of the entire magma column to a pressure of
442 80 MPa (**Fig. 6(iii)**).

443

444 Importantly, observations of volcanic gases, whilst consistent with models of decompressional degassing,
445 are also consistent with a fraction of the gases being derived from a deep (>2-3 km) exsolved volatile
446 phase generated during prolonged crystallization (**Fig. 6b**). In this scenario, as magma evolves from basalt
447 to trachybasalt (at 80 MPa, after about 50% crystallization) it generates ~ 0.6 wt% exsolved fluids. The
448 exsolved volatile phase is carbon and chlorine-rich at melt fractions $\geq \sim 0.7$ (**Fig. 6(vi)**), then 13rystalliz at
449 molar C/S ratios of ~1.5-2.5 and molar S/Cl ratios of 1-2. These values are consistent with volcanic gas
450 compositions observed at the surface (**Fig. 7a**), which raises the possibility that some, and perhaps a large
451 fraction, of the gases fluxing through the conduit and into the atmosphere may be derived from the fluids
452 produced during equilibrium 13rystallization of basalts at depths of ~3 km or deeper. Indeed, Métrich et
453 al. (2011) concluded from melt inclusion geochemistry that primitive basalts pond at ~ 3 km depth where
454 they fractionate during ~ 50-60% crystallization to form trachybasalts with 56-60 wt% SiO₂. 50%
455 equilibrium crystallization would produce ~1000 ppm exsolved S, ~1000 ppm exsolved Cl and 0.3 to 0.6
456 wt% H₂O (**Table S2; Supplementary Material**); this would require the intrusion of 0.04-0.09 km³
457 magma/year, similar to the volume required for the postulated vertical large-scale convection to shallow
458 depths necessary to supply outgassed SO₂ and HCl. Furthermore, the latent heat generated by extensive
459 shallow magma crystallization may be sufficient to thermally buffer the magma reservoir and to maintain
460 a hot conduit (Métrich et al., 2011). Ascent of a deep-derived exsolved volatile phase, possibly with
461 subsidiary melt, could advect heat to the conduit, allowing it to remain at a constant temperature over
462 decadal timescales.

463

464 Importantly, fluids generated by second boiling would be relatively water-poor owing to the relatively low
465 water content of Yasur basalts and the high solubility of water in silicate melts. The high water contents of
466 Yasur volcanic gases (Métrich et al., 2011; Woitischek et al., 2020) would thus seem to be good evidence
467 for some decompressional degassing and magma convection. The water content of the volcanic gases is,
468 however, an order of magnitude higher than expected from decompressional degassing alone, which may
469 suggest a contribution from meteoric waters. In summary, it is likely that the volcanic gases emitted to the
470 atmosphere record mixing between exsolved volatiles generated by deep (>2-3 km) isobaric second
471 boiling and by decompression-degassing accompanying convection, with the possible addition of shallow
472 meteoric water, although this is not well constrained (**Fig. 7a**).

473

474 Gas data for Stromboli volcano (**Fig. 7b**) illustrate the wide range of gas compositions observed during
475 eruptive activity (Aiuppa et al., 2010; Allard, 2010; Burton et al., 2007; Tamburello et al., 2012).
476 Stromboli's volcanic gases are dominated by H₂O (48 to 98 mol%, mean 80 mol%) and also contain CO₂
477 (2-50 mol%, mean 17 mol%) and SO₂ (0.2 to 14 mol%, mean 3 mol%). During paroxysms and
478 strombolian explosions, the carbon content of the emitted gases extends to 50 mol% CO₂ with a molar C/S
479 of >10 and up to 47, a low molar H₂O/CO₂ (typically 1-3) and high S/Cl ratios (mean 4.7 ± 0.08).

480 Between explosions, the gas molar C/S is <15, H₂O/CO₂ is 1.5 to 6.5 and S/Cl is 1-1.5. Stromboli is fed
481 by magmas with a much higher volatile content than at Yasur, as evidenced by studies of melt inclusions
482 (Métrich et al., 2010). As an approximation of the Stromboli system, we use a starting basalt composition
483 (**Supplementary Table S1**) with 3 wt% H₂O, 2 wt% CO₂, 0.2 wt% Cl and 0.25 wt% S (Métrich et al.,
484 2010) for the modeling (details given in **supplementary material**).

485
486 As for Yasur, and consistent with previous studies (Aiuppa et al., 2010; Allard, 2010; Métrich et al.,
487 2010), we find that decompressional degassing of the exsolved volatile phase for Stromboli compositions
488 causes the C/S ratio to decrease from >100 at pressures between 240 and 100 MPa to ~1-2 at the surface
489 (**Fig. 7b; table S4**). Also similar to Yasur, the volcanic gas molar S/Cl ratio increases with decreasing
490 pressure from <0.1 at depth to 1-10 at the surface, governed by the relative partitioning behaviour of Cl
491 and S with pressure (Lesne et al., 2011; Tattitch et al., 2021). The fluids generated during isobaric
492 crystallization (second boiling), in contrast, initially have low C/S and S/Cl but converge on C/S ~5-8 and
493 S/Cl ~ 1-2 after 50% crystallisation.

494
495 The high molar CO₂ content of the gases during strombolian explosions and paroxysms suggest triggering
496 by a deep-derived gas phase (Aiuppa et al., 2010; Allard, 2010; Burton et al., 2007; Métrich et al., 2010),
497 with the gases emitted during quiescent degassing fed by more shallowly-equilibrated gases. However, a
498 comparison of gas compositions to a decompressional degassing model (**Fig. 7b**) shows that the SO₂/HCl
499 systematics (Burton et al., 2007) are not obviously consistent with such an interpretation. Indeed,
500 decompressional degassing models predict 'deeper'-equilibrated gases to have a lower S/Cl than shallow-
501 equilibrated gases; this trend reflects the decrease in the fluid-melt partition coefficient for Cl with
502 decreasing pressure, in tandem with a dramatic increase in the fluid-melt partition coefficient for sulfur
503 (Lesne et al., 2011). As noted above, we have only a limited understanding of the chlorine systematics in
504 volcanic gases. However, a likely explanation is that degassing during paroxysms is more 'closed' than
505 during persistent degassing, consistent with the higher observed molar S/Cl as well as the high molar C/S.
506 The observed S/Cl ratio of ~ 2 of the quiescent plume at Stromboli (Burton et al., 2007), which accounts
507 for the bulk of the outgassing flux (Allard et al., 2008), is equally consistent with an exsolved volatile
508 phase being generated by decompression degassing or by second boiling processes at depth, or a mixture
509 of both sources (**Fig. 7b**).

510
511 Petrological studies provide additional constraints on the Stromboli magmatic system. Stromboli is fed by
512 primitive, volatile-rich high K₂O (HK) basalts with 49-51 wt% SiO₂ and CaO/Al₂O₃ > 0.6 (Métrich et al.,
513 2010) stored at depths of 7-10 km beneath the summit (Bertagnini et al., 2003; Métrich et al., 2010). Large
514 paroxysms erupt this low-density CO₂-rich HK basalt as 'golden pumice' (Pichavant et al., 2009; Rosi et
515 al., 2000), with little evidence for mixing with shallow-stored magma, consistent with rapid and primarily
516 closed system decompression (Métrich et al., 2021; Pichavant et al., 2009). Eponymous strombolian

517 activity, in contrast, ejects crystal-rich, degassed shoshonitic basalt (51-54 wt% SiO₂) stored at 2-4 km
518 beneath the summit and produced by 20-30% fractional crystallisation of HK basalts at depth (Landi et al.,
519 2004; Métrich et al., 2010; Métrich et al., 2001; Métrich et al., 2005; Vergnolle and Métrich, 2021).
520 Deep and shallow magmas mix only during smaller paroxysms (LaFelice and Landi, 2011a). CO₂-rich
521 fluids derived from ponding and crystallising basalts at depth, in contrast, flux through the shallow
522 system, dehydrating the overlying magma and promoting extensive crystallisation within the shallow
523 conduit (Landi et al., 2004; Métrich et al., 2001). Resulting crystal networks may trap rising fluids to form
524 gas pockets; the release of these accumulated gases when they overcome the forces holding the crystals
525 together (the effective yield strength) may explain the 'normal' strombolian activity (Barth et al., 2019;
526 Belien et al., 2010; Oppenheimer et al., 2015; Suckale et al., 2016; Woitischek et al., 2020) that produces
527 highly degassed, crystalline and high viscosity bombs, remnants of the degassed 'plug' (Caracciolo et al.,
528 2021; Gurioli et al., 2014; Lautze and Houghton, 2007). Triggers for paroxysmal activity, in contrast, are
529 debated. One suggestion is that they may be triggered by rapid (days) ascent of HK magma (La Felice and
530 Landi, 2011; Métrich et al., 2010; Métrich et al., 2021) caused by increases in overpressure in the deep
531 storage area or by the greater buoyancy of gas-rich basaltic magma (Allard, 2010; Métrich et al., 2005;
532 Métrich et al., 2021). A 'top-down' trigger has been suggested for paroxysms preceded by high gas hold-
533 up and lava effusion, which promote decompression of the shallow conduit (Calvari et al., 2011; Ripepe et
534 al., 2015). These contrasting scenarios raise interesting questions about the role of deep volatiles in
535 modulating eruptive activity.

536
537 Volcanic gas compositions measured at Mount Etna (Italy) reveal that paroxysmal phases of ash emission
538 and lava fountaining during 2001 (Aiuppa et al., 2002) and mid- and late November 2002 (Aiuppa et al.,
539 2004) were accompanied by volcanic gases with low molar S/Cl ratios (< 1) and high SO₂ fluxes (15,000
540 t/d). Conversely, a trend of increasing S/Cl ratios and decreasing SO₂ flux accompanied the transition of
541 volcanic activity toward mild strombolian activity and finally passive degassing with minor effusive
542 activity. A sulfur and halogen degassing model developed to explain these trends (Aiuppa, 2009) suggest
543 that the S/Cl ratio in the gas phase increases by decompression degassing as magma nears the surface
544 because of the increasing preference of Cl for the melt and of S for the gas (see **Fig. 7b**), as observed in
545 geochemical studies (Spilliaert et al., 2006). The Cl-rich gas emitted during the paroxysms may therefore
546 represent a deeper exsolved volatile phase, perhaps generated through second boiling processes at depth.
547 Such a fluid phase would fuel the development of deep, volatile-rich melts co-existing with the Cl-rich
548 exsolved volatile phase implicated in driving paroxysms at Etna. Evidence of high S/Cl ratios in volcanic
549 gases during fountaining (Allard et al., 2005), in contrast, may record a large shallow influx of undegassed
550 magma accompanied by relatively shallow degassing at low pressures.

551
552 Now that gas geochemical monitoring is commonplace, and often automated, trends prior to explosive
553 eruptions at open-vent volcanoes are increasingly well characterised. Pulses of CO₂ are often observed

554 prior to paroxysms and other forms of explosive eruption, manifest as increases in the C/S ratio days to
555 weeks prior to eruption (Aiuppa et al., 2017; Aiuppa et al., 2007; de Moor et al., 2017). At Villarrica
556 volcano, Chile, for example, an increase in volcanic gas C/S after January 2015 preceded the 3 March
557 2015 paroxysm (Aiuppa et al., 2017). The same pre-eruptive period saw an increase of >50 m in the
558 height of the persistent lava lake (from 27 February; Johnson et al., 2018), suggesting increased gas hold-
559 up. Similar signals preceded explosive activity at Turrialba Volcano, Costa Rica, in 2014 and 2015, where
560 pulses of deeply derived CO₂-rich gas ($C/S_{total} > 4.5$) have been observed up to 2 weeks before eruptions
561 (de Moor et al., 2016). These signals of 'deep-derived' exsolved volatiles, arriving at the surface in the
562 absence of (or preceding) magma, provide further evidence of a significant, exsolved and segregated
563 exsolved volatile phase at depth that is capable of fluxing up through the shallow plumbing system prior to
564 and during explosive eruptions, including paroxysms, supported by studies of gas fluxes and scoria
565 textures, which illustrate degassing-driven mingling between deeper hotter melt and degassed, more
566 crystalline magma derived from the upper parts of the conduit (Gurioli et al., 2008).

567
568 At intermediate open-vent volcanic systems, evolved melts with high fractions of exsolved volatiles may
569 dominate the magma reservoir and the contribution of second boiling to the exsolved volatile phase may
570 be far more significant. Although few long time series of volcanic gas compositions exist for these
571 systems, one exception is Asama Volcano in central Japan, a persistently degassing volcano that erupts
572 every few years (Shinohara et al., 2015). Here, periods of high gas flux coincide with periods of eruptions
573 and elevated seismic activity. Low SO₂ emission rates characterise periods of low eruptive activity.
574 SO₂/HCl ratios in the gas are high during eruptive periods and lower during eruptive pauses, a pattern
575 consistent with eruptive periods dominated by decompressional degassing (Shinohara et al., 2015). There
576 is no clear variation in C/S between active and inactive periods (Shinohara et al., 2015). In contrast,
577 Soufrière Hills Volcano, which erupted crystal-rich andesite episodically between 1995 and 2011, showed
578 a clear pattern of molar S/Cl > 1 during eruptive pauses and S/Cl < 1 during eruptive episodes dominated
579 by lava dome building that remained consistent over many years of observation (Christopher et al., 2010;
580 Edmonds et al., 2001). These gas characteristics can be explained by cessation of gaseous HCl flux during
581 eruptive pauses whilst a near-constant (or slowly declining) SO₂ flux is sustained (Christopher et al.,
582 2010). The data are insufficient to assess whether systematic temporal variations in molar C/S exist.

583
584 Volcanic gas compositions arising from models of decompressional degassing versus isobaric second
585 boiling are compared with observations in **Fig. 8**. The Soufrière Hills andesite is crystal-rich with a
586 rhyolitic carrier liquid; fluid-melt partition coefficients for chlorine and sulfur are estimated to be ~20-30
587 and ~200-500, respectively, at the pressures of storage prior to eruption; with decreasing pressure, D_{Cl}
588 decreases to ~1 and D_s increases to >1000 (Tattitch et al., 2021; Webster et al., 2011) (**Tables S5, S8**).
589 Bulk chlorine and sulfur contents are poorly constrained; we use 0.15 wt% for Cl (informed by melt
590 inclusion studies; (Edmonds et al., 2001) and 0.3 wt% S for initializing the isobaric degassing models at

591 F=1. In general, the deep MVP generated by second boiling has an initially high molar C/S ratio, which
592 then decreases and converges on a composition of ~ 1 -2 after $\sim 60\%$ crystallisation at a range of pressures
593 (**Table S8**), and an initially low molar S/Cl ratio that increases and converges on values between 2 and 3
594 (**Fig. 8a**). These values yield Cl and S melt concentrations of 700-1000 ppm Cl and 50-100 ppm S after
595 90% crystallisation, consistent with melt inclusion studies of SHV rhyolitic melt inclusions (Edmonds et
596 al., 2001) (**Table S3, supplementary material**). After the 90% crystallisation required to generate
597 rhyolite melt, there is 4 - 7 wt% exsolved water-rich MVP (**supplementary material table S8**) with a
598 molar C/S of ~ 1 . Rhyolitic melts beneath Soufrière Hills Volcano are therefore likely to have significant
599 fractions of MVP that must be migrating to the surface, even during eruptive pauses, to supply the
600 outgassing flux (Christopher et al., 2015).

601

602 Rhyolitic melt starting with 8 wt% H₂O, 1 wt% CO₂, 0.1 wt% Cl and 0.01 wt% S (consistent with the melt
603 concentrations measured in melt inclusions, **table S3**) subjected to slow degassing-induced crystallisation
604 yields a Cl-rich volcanic gas (**Fig. 8a; table S5**), consistent with the S-poor melt. Variable mixing
605 between the deep (> 2 -3 km) MVP generated during second boiling and a decompression-derived MVP
606 during eruptions could thus yield a volcanic gas with a low S/Cl ratio during eruptive periods
607 (contributions from both deep- and decompression-derived MVP) and a high SO₂ flux with a high S/Cl
608 ratio during eruptive pauses (contributions dominated by the deep MVP generated through second
609 boiling), which is precisely what we observe (Christopher et al., 2010; **Fig. 8b**). This example clearly
610 demonstrates the importance of a segregated deep MVP in sustaining outgassing at more evolved open-
611 vent volcanoes; this mechanism may be generic to other, similar volcanic systems globally (e.g.,
612 Tungurahua and Reventador, Ecuador; Bagana volcano, Papua New Guinea).

613

614 *Geophysical evidence for decoupled flow of an exsolved magmatic volatile phase in the crust*

615

616 Seismicity related to shallow degassing and eruption

617 Low frequency (LF, or long period, LP) earthquakes are a common feature of active volcanoes (McNutt
618 and Roman, 2015). When LP earthquakes are closely spaced in time, the signals may merge to form a
619 continuous tremor signal. LP earthquakes (and tremor) are thought to be caused by fluid pressurisation,
620 including the resonant response of fluids in conduits or dykes (Chouet, 1996; Neuberg et al., 2000). Very-
621 long-period (VLP) and ultra-long-period (ULP) events detected using broadband seismometers originate
622 at shallow depths (≤ 1.5 km) and are associated with eruptions or vigorous fumarolic activity (McNutt and
623 Roman, 2015; Sanderson et al., 2010). Although the specific interpretations of VLP and ULP events vary,
624 there is general agreement that they provide evidence of short-term deformation accompanying eruptive
625 activity (Chouet et al., 1999; James et al., 2006; Oppenheimer et al., 2020; Ripepe et al., 2015; Suckale et
626 al., 2016). Recent studies highlight links between seismic signals and degassing flux at many open system
627 volcanoes (Zuccarello et al., 2013). A direct link between VLP signals and strombolian activity was first

628 identified at Stromboli volcano: very long period (VLP) signals sourced from a few hundred metres depth
629 in the conduit were thought to originate from the rise and bursting of large slugs of gas within the conduit
630 (Chouet et al., 1999). More recent data clearly show a VLP signal preceding each event together with
631 synchronous thermal and SO₂ flux signals accompanying each explosion (Gurioli et al., 2014; Tamburello
632 et al., 2012). Although the form of gas transport up the conduit linked to these seismic (and infrasound)
633 signals has long been interpreted as a rising gas slug (Blackburn et al., 1976; Jaupart and Vergnolle,
634 1988; Ripepe et al., 2001), an alternative model calls for gas accumulation in, and release from, a crystal-
635 rich, shallow plug (Del Bello et al., 2015; Gurioli et al., 2014; Oppenheimer et al., 2015; Suckale et al.,
636 2016). Correlations between VLP events and degassing have also been observed at Etna (Zuccarello et al.,
637 2013), Merapi (Hidayat et al., 2002), Asama (Kazahaya et al., 2011) and Erebus (Aster et al., 2008).
638 Similarly, episodic explosive activity modulated by accumulation and release of a gas phase beneath a
639 rigid or semi-rigid plugs may explain shallow (~300 m) VLP signals at Fuego (Waite et al., 2013) and
640 inflation-deflation cycles and periodic explosions at Santiaguito (Bluth and Rose, 2004; Johnson et al.,
641 2014).

642

643 Seismicity and strain signals related to migration of volatiles at depth

644 Deeper geophysical signals related to the movement or pressurisation by an MVP are limited. MVP-
645 related seismic signals in the upper crust have been observed at Popocatepetl volcano, where VLP signals
646 accompany volcanic degassing bursts at a depth of ~1.5 km. One interpretation is that these signals record
647 the opening of an escape pathway for an exsolved volatile phase that accumulated because of second
648 boiling in a shallow sill (Chouet et al., 2005). Sharper pressure transients associated with expanding gas
649 pockets may generate VLP signals to depths of ≤ 3 km (Arciniega-Ceballos et al., 2008). Another example
650 of upper crustal movement of MVP comes from Soufrière Hills, where strain signals observed during
651 vulcanian explosions and gas emission events record inflation of a shallow conduit and near-simultaneous
652 contraction of deeper magma reservoirs (> 5 km depth) (Hautmann et al., 2014). This strain pattern has
653 been interpreted as rapid upward migration of a buoyant MVP, initiated by a sudden destabilisation of
654 large pockets of already segregated fluid in the magma reservoirs (Christopher et al., 2015; Hautmann et
655 al., 2014; Linde et al., 2010).

656

657 Deep long period earthquakes (DLPs) associated with volcanoes have been observed in the mid-lower
658 crust or mantle (Aso and Tsai, 2014; Melnik et al., 2020; Wech et al., 2020). Although their origin is
659 enigmatic, some studies have linked DLPs to an exsolved MVP. A striking example is Mauna Kea,
660 Hawai'i, where more than a million DLPs have been recorded in the past 19 years (Wech et al., 2020).
661 These events are not linked to eruptions but have been ascribed to the second boiling of deep (near-Moho)
662 magma intrusions. Other interpretations of DLPs include thermal stresses set up by cooling magmas (Aso
663 and Tsai, 2014) and rapid changes of magmatic pressure in the lower crust caused by rapid nucleation and
664 growth of gas bubbles in response to the slow upwelling of volatile-saturated magma (Melnik et al., 2020).

665 The latter explanation for the Klyuchevskoy volcanic group relates to primary melts that may contain ≤ 4
666 wt% H₂O and 0.6 wt% CO₂, which would cause volatile saturation at 800 MPa (~30 km). Alternatively,
667 these DLPs may record pressurisation of a deep reservoir and consequent transfer of the magma towards
668 the surface. The relatively fast upward migration of long-period activity at Klyuchevskoy (months)
669 suggests that a hydraulic connection is maintained between deep and shallow magmatic reservoirs
670 (Shapiro et al., 2017) and the upward transport includes a large fluid component (Koulakov et al., 2012).

671

672 Seismic tomography within the crust beneath volcanoes yields a picture of how melt versus MVP-rich
673 areas may be distributed (Kuznetsov et al., 2017; Londoño et al., 2018; Vargas et al., 2017). A porous
674 medium saturated with gas has a low compression modulus that yields low velocity P-waves but no
675 decrease in S-wave velocity (a low Vp/Vs ratio). High P-wave velocities and low S-wave velocity (high
676 Vp/Vs ratios) may, in contrast, indicate the presence of melts, i.e. an active magma reservoir (Kuznetsov
677 et al., 2017). In this way, repeat tomographic studies can monitor temporal changes in the structure of
678 magmatic systems. At Nevado del Ruiz, Colombia, for example, the distribution of low and high Vp/Vs
679 regions changes on yearly timescales (Londoño et al., 2018; Vargas et al., 2017). Nevado del Ruiz is an
680 open-vent volcano with considerable fluxes of SO₂ emitted continuously (Lages et al., 2018). Here a high
681 Vp/Vs anomaly 2-4 km prior to 2010 is interpreted as a volatile-rich melt reservoir, the lower boundary of
682 which moved upward in 2011-2012 and was replaced by a region of low Vp/Vs, interpreted as a gas-rich
683 region undergoing second boiling; this was associated with intense, persistent outgassing at the surface
684 (Vargas et al., 2017). Tomographic studies of Mt. Spurr, an intermittently open-vent volcano, show finger-
685 shaped seismic anomalies with a high Vp/Vs ratio beneath the location of intensive fumarolic activity in
686 2004-2005 that are interpreted to represent separate conduits of magma and volatiles (Koulakov et al.,
687 2018). A shallow (0-2 km) region of low Vp/Vs directly above is interpreted as a large-scale degassing
688 event, whereby gases were segregated and migrated up to the surface (Koulakov et al., 2018). Although
689 limited, these studies illustrate the potential for future monitoring of volatile and melt distributions
690 beneath open volcanic systems.

691

692 *Open-vent persistent volcanic outgassing is promoted in complex, extensional tectonic regions*

693

694 Open-vent volcanoes that generate high outgassing fluxes (**Fig. 1**) are often located in regions of complex
695 tectonics and local extension. The correspondence between the locations of open-vent volcanoes and
696 major crustal extensional structures highlights the role of tectonics in promoting magma intrusion, MVP
697 segregation, and MVP migration to the surface. Although the processes that modulate MVP behaviour are
698 not well known, the association of persistent degassers with extensional regions suggests that a) extension
699 leads to high intrusive/extrusive magma ratios and therefore provides large upward fluxes of exsolved
700 volatiles through second boiling; b) extension may promote the gravitational segregation of low density

701 MVP phases in shallow reservoirs, allowing their migration and outgassing; and/or c) faults and shear
702 zones in extensional regions may become permeable pathways for deep fluids (**Fig. 9**).

703

704 Many persistently degassing open-vent volcanoes —Popocatepetl (Mexico), Fuego and Pacaya
705 (Guatemala), Turrialba and Poas (Costa Rica), Telica and Masaya (Nicaragua)— are located within
706 grabens along the Central American Volcanic Arc (CAVA) and Trans-Mexican Volcanic Belt (TMVB).
707 For example, the open-vent volcano Masaya lies within a large arc-parallel basin – the Nicaraguan graben
708 – that contains Lake Nicaragua and Lake Managua (Morgan et al., 2008) (**Fig. 1; 9a**). Masaya exhibits
709 cycles of intense outgassing that coincide with lava lake activity (Delmelle et al., 1999; Stoiber et al.,
710 1986) (Delmelle et al., 1999; Stoiber et al., 1986) but few eruptions—there has been no major effusive
711 activity since 1965 (Harris, 2009)—although Masaya has a history of large basaltic plinian eruptions (at 6
712 ka, 2.1 ka, 1.9 ka; (Pérez et al., 2020; Walker et al., 1993; Williams, 1983). Presently, there is a lava lake
713 at Masaya and evidence for a shallow subvolcanic reservoir (Aiuppa et al., 2018); petrological studies
714 have reconstructed the equilibration pressure of the super-hydrous melts responsible for explosive activity
715 to below the seismic Moho (Pérez et al., 2020). The extension rate in Nicaragua has been estimated from
716 initiation of arc splitting and dating of volcanic products (Plank et al., 2002). The observed crustal
717 thickness of 30-35 km greatly exceeds the ~10 km expected for the estimated 100 km extension over 15
718 Ma, suggesting that intrusive magmatism has infilled the space created by extension at a rate of 90-180
719 km³/km/Ma (Morgan et al., 2008). Moreover, the estimated intrusive flux for Nicaragua is ~100 times the
720 estimated volcanic output rate (Carr et al., 2003). This intrusive/extrusive ratio is much larger than the
721 global average, which is ~ 5:1 (with a range of 1:1 to >35:1) (Crisp, 1984; White et al., 2006). Over the
722 entire arc, regions of greatest extension also have the highest magma productivity and the strongest
723 geochemical slab signature (as demonstrated by geochemical indices Ba/La and Yb/La; (Burkart and Self,
724 1985; Carr et al., 2003). Nicaragua, specifically, has the largest magma productivity (intrusive and
725 extrusive together), the highest rates of extension and slab flux and the strongest slab melting and source
726 melting signals (Carr et al., 2003). Although it is unclear whether the large magma fluxes are a cause or
727 an effect of upper plate extension, the large fluxes of intrusive magmas beneath the Managua graben allow
728 ample opportunity for extensive second boiling and decompression-degassing and production of a deep
729 exsolved MVP. Venting of these deep-derived fluids advects heat to the shallow system and maintains a
730 hot conduit.

731

732 Ambrym, a top-ranking volcanic open-vent outgasser (**Fig. 1, 4**) located in the New Hebrides arc (**Fig.**
733 **9b**), is situated in the transition zone between a compressional regime in the central arc (Calmant et al.,
734 2003) and an extensional regime in the south (Beier et al., 2018). The relative motion between the central
735 and neighbouring northern and southern arc segments, respectively, is accommodated by dextral strike-
736 slip zones (Pelletier et al., 1998). Ambrym, with its 12 km wide caldera and the resurgent domes of
737 Marum and Benbow, is located exactly at the transition from regional subsidence to strike-slip faulting

738 (Picard et al., 1994). Changes in the stress field from compression to extension (plus rotation) has created
739 a complex polybaric magmatic system (Beier et al., 2018), including accumulation of large intrusive
740 volumes, crystallisation in shallow reservoirs and resulting large fluxes of exsolved volatiles that
741 contribute to the persistent outgassing observed at Ambrym (Allard et al., 2016).

742

743 Etna has developed on the margin of the Hyblean plateau, the foreland to the Late Tertiary Maghrebian-
744 Calabrian thrust belt, a compressional regime that started extending at ~ 0.5 Ma (Hirn et al., 1997; Laigle
745 and Hirn, 1999) (**Fig. 9c**). Crustal-scale normal faults imaged by reflection seismology extend over 20 km;
746 their size, depth, location and evidence of activity suggests that these faults are the source of large
747 earthquakes, which are associated with enhanced volcanism in time and space (Hirn et al., 1997). The
748 specific location of Etna might be related to extension within a narrow zone of active normal faulting that
749 stretches from the Hyblean Plateau in eastern Sicily to northern Calabria (Monaco et al., 1997). A high
750 seismic velocity zone with a lateral dimension of ~6 km has been imaged beneath the summit at 9-18 km
751 depth (Hirn et al., 1997). This body probably comprises cumulates produced from intrusive magmas,
752 fragments of which are occasionally erupted as cognate xenoliths (Corsaro et al., 2014). This cumulate
753 body likely contains significant volumes of volatile-rich melts generated through second boiling as well as
754 regions dominated by an exsolved volatile phase. These fluids may mix with intruding basalts and ascend
755 to shallow levels in the plumbing system shortly before eruptions, contributing to the large and persistent
756 outgassing fluxes of Etna.

757

758 Persistently outgassing volcanoes in extensional (continental) regions include Erta Ale, Oldoinyo Lengai,
759 Nyiragongo and Nyamuragira in the East African Rift and Erebus in the West Antarctic Rift system. A
760 global link between outgassing and tectonics was suggested by Tamburello et al. (2018) to explain the
761 distribution of high CO₂-emitting volcanic areas, which are focused in the extensional regions of arcs and
762 in continental rifts. However, oceanic regions of extension are conspicuous for their lack of persistent
763 volcanic outgassers. Iceland, for example, sits astride the Mid-Atlantic Ridge and has frequent eruptions
764 but has no lava lakes or persistently outgassing conduits. Instead, frequent eruptions follow short periods
765 of unrest (including increased outgassing) and return rapidly to closed-system behaviour once the eruption
766 is over, although diffuse CO₂ degassing between eruptions may be linked to magma intrusions at depth
767 (Ilyinskaya et al., 2018). In this respect activity is more similar to other ocean islands such as Reunion,
768 Galapagos or the Canary islands, where sulfur-rich degassing occurs during, but not between, eruptions
769 (Di Muro et al., 2016).

770

771 **Depth-integrated magma degassing drives persistent outgassing and eruptive activity at open-vent**
772 **volcanoes**

773

774 A conclusion that can be drawn from the data and models presented above is that open-vent volcanoes
775 may be thought of as predominantly gas, rather than lava, emitters. A corollary is that degassed magmas
776 accumulate in the crust beneath open-vent volcanoes, thereby growing the crust endogenously. Open-vent
777 volcanoes often occur in regions of crustal extension, which yield the accommodation space for large
778 volumes of intruded magmas that ultimately form dry plutons once they crystallise, exsolve and lose their
779 volatiles. Open-vent volcanoes are active for decades to millennia; their longevity may be controlled by
780 the tectonics of the crust, which may cause different arc segments to ‘switch on and off’ over time (de
781 Moor et al., 2017). Eruptions of open-vent volcanoes may be triggered by the ascent of segregated
782 exsolved volatiles that flux through the shallow system, or by volatile-rich melts that migrate rapidly from
783 deeper levels in the crust, exsolving large volumes of volatiles as they ascend. Therefore, although
784 traditionally it has been assumed that magma is the ‘carrier’ for advecting volatiles—requiring mass
785 balance in the upper crust to account for open-vent outgassing fluxes (i.e., the convection model)—we
786 have shown instead that large volumes of intruded magma at depth, stored at multiple levels throughout
787 the crust, provide a potential source of segregated exsolved volatiles, which inevitably must contribute to
788 the large outgassing fluxes at open-vent volcanoes. Our model both removes the necessity for the volatiles
789 to be supplied by continuous, large-vertical-scale bimodal flow and alleviates the space problem caused by
790 the need to store large volumes of degassed magma within the shallowest parts of the crust.

791

792 *For basalt-dominated open-vent volcanoes (Fig. 10a)* with basalt or alkali basalt lava lakes or open-vents
793 (e.g. Stromboli, Yasur, Villarrica, Masaya, Fuego), volatiles may be delivered to the atmosphere through a
794 combination of deep and shallow mechanisms, both consistent with the volcanic gas compositions
795 observed at these volcanoes (Figs 5-7). Primitive basalts (which may be saturated in an exsolved volatile
796 phase even at mantle depths in some cases) typically undergo $\geq 50\%$ crystallisation in the crust to produce
797 basaltic andesites or trachybasalts that dominate the shallow storage regions beneath these volcanoes.
798 Exsolved volatiles generated through second boiling may migrate via capillary flow in crystal-rich mush
799 in mid- and upper crustal magma storage regions, accumulating and segregating, perhaps giving rise to
800 pockets of exsolved volatiles that may ascend rapidly to the surface and trigger paroxysms (Fig. 3).
801 Primitive melts may be drawn into the conduit in the wake of these pockets of exsolved fluids. Although
802 conduit convection may allow magmas to ascend to near atmospheric pressure, outgas and then sink,
803 convection likely acts in tandem with the fluxing of deep-derived exsolved volatiles through the shallow
804 conduit system. Together these processes may explain much of the outgassing volatile flux, as exemplified
805 by Stromboli (Fig. 7b). The balance between convective degassing and deep MVP fluxing likely differs
806 between volcanoes depending on both the depth of magma storage and crystallisation and the total volatile
807 content of the magma. Volatile-rich magmas stored at relatively shallow depths are likely to produce a
808 large volume of exsolved volatiles during even modest amounts of crystallisation. We note that this
809 concept of exsolved volatiles being integrated over large depth ranges in the crust to supply open vent
810 outgassing is consistent with geochemical evidence from volcanic rocks that suggest that melt as well as

811 crystals in magmas mingle over similarly large depth ranges (Cashman and Edmonds, 2019; Ruth et al.,
812 2018). A high magma intrusion rate will buffer the melt composition in the subvolcanic reservoir to
813 produce monotonous erupted compositions and long-lived outgassing. The latent heating generated by
814 extensive subvolcanic crystallisation combined with the rise of deep-derived exsolved volatiles (which
815 efficiently advect heat) may produce sufficient heat to maintain hot conduits and lava lakes.

816
817 *For intermediate composition open-vent volcanoes (dominated by andesites and dacites) (Fig. 10b)* (e.g.
818 Bagana, Soufrière Hills, Santiaguito, Anatahan), magma crystallisation over long timescales generates
819 extensive regions of mush. Crystallisation of basalts at lower crustal depths may generate low viscosity
820 hydrous or even 'superhydrous' basaltic andesite or andesite melts, as inferred for Kamchatka (Goltz et
821 al., 2020). These melts may be further enriched in incompatible elements (including volatiles) upon
822 mixing with highly evolved water-rich melt lenses in deep crustal mush. Petrological and experimental
823 studies suggest mid-crustal water contents of 5-11 wt% in basaltic andesites from the Lesser Antilles
824 (Edmonds et al., 2016; Melekhova et al., 2017). These volatile-rich melts may rise up to the mid and
825 upper crust through percolation along grain boundaries or by channelised reactive flow. Intrusion of
826 volatile-rich basaltic andesite into shallower, more evolved mush-dominated reservoirs can induce partial
827 melting, gas sparging (Bachmann and Bergantz, 2006) and/or trigger gravitational destabilisation or
828 eruption (Christopher et al., 2015). The volatile-rich melts may generate substantial fractions of exsolved
829 volatiles in mid and upper crustal mush-dominated reservoirs, which may accumulate and segregate from
830 their rhyolitic melt lenses over millennia. These volatile-rich lenses may be later tapped by eruptions and
831 drive persistent and long-lived volcanic outgassing. Importantly, bimodal flow and convective degassing
832 are precluded in volcanoes dominated by crystal-rich, hydrous intermediate-composition magma because
833 of extensive decompression-induced crystallisation and resulting high bulk viscosity of the magma. In
834 these systems, persistent degassing requires volatile migration that is independent of magma migration.

835
836 *A generic model for the degassing regime at open-vent volcanoes* brings together our understanding of
837 magmatic crystallisation, mixing and storage processes with our observations of volcanic gas flux and
838 composition at open-vent volcanoes. Intrusive, unerupted magmas crystallising at a range of crustal depths
839 generate a substantial exsolved volatile phase, which is fluxed into the overlying system and up through
840 conduits. Volatile fluxing advects heat and brings with it small volumes of primitive melts that replenish
841 the melt resident in the shallow magma storage and conduit systems. Although basalt-dominated
842 reservoirs may also experience subsidiary convection, convection is unlikely in more evolved, mush-
843 dominated magmatic systems, where the outgassing flux will instead be dominated by the fluxing of a
844 deep-derived MVP generated through second boiling. In this model, large bodies of crystal-rich mush
845 generated through extensive crystallisation remain *in situ* at a range of depths, with no requirement for
846 magmas to convect to atmospheric pressure and back down again. Volcanic gases emitted from these

847 volcanoes are the integrated products of the degassing of melts at a range of crustal depths that have
848 undergone various degrees of crystallisation and mixing.

849

850 **Conclusions**

- 851 1. Open-vent volcanoes produce *large outgassing fluxes*, much greater than can be supplied by
852 erupting magmas. Open-vent volcanoes may be thought of as gas vents connecting the mantle
853 and/or crust to the atmosphere.
- 854 2. Open-vent volcanoes produce *explosive and gas-rich eruptions*, e.g. violent strombolian,
855 vulcanian, paroxysms, that are triggered by the rise of volatile-rich melts and/or fluxing of
856 segregated exsolved volatiles from deeper mush-dominated magma storage regions.
- 857 3. *Volcanic gas compositions at open-vent volcanoes* are likely derived from a mixture of exsolved
858 volatile produced from decompressional degassing, whereby magmas degas during their ascent to
859 atmospheric pressure; and isobaric (or polybaric) second boiling in the crust, which generates a
860 substantial volume of exsolved volatiles during crystallization.
- 861 4. High fluxes of deep exsolved volatiles are sourced from the *second boiling of intrusive magmas* in
862 the mid to lower crust. These deep-derived exsolved volatiles flux through shallow volcanic
863 systems, advecting heat, sustaining persistent degassing and triggering eruptions. These processes
864 are particularly important for more evolved, water-rich volcanic systems.
- 865 5. *Bimodal flow and magma convection* may operate in low viscosity basaltic systems, which brings
866 magma up to near atmospheric pressure to outgas and then sink back down, but this mechanism
867 acts in tandem with fluxing by a deeper-derived volatile phase. Convection is not likely to be
868 important in more water-rich, more evolved volcanic systems, due to the extensive degassing-
869 induced crystallisation in the conduit, which will stall magma return flow by viscous inhibition.
- 870 6. Intrusion and degassing of magma into the crust beneath open-vent volcanoes is accommodated
871 by *extensional tectonics* and the extension plays a role in allowing exsolved fluids to migrate up to
872 the shallow volcanic systems. The location and longevity of open-vent volcanic outgassing and
873 activity is likely controlled by tectonics.
- 874 7. Open-vent volcanic outgassing is an integrated product of the *degassing of a vertically-protracted*
875 *magmatic storage and transport system*, not merely a shallow magma reservoir. A great challenge
876 for volcano monitoring in the future will be to detect and understand both geochemical and
877 geophysical signals from the mid and lower crust to enhance eruption forecasting.
- 878 8. Accurate measurements of outgassing volatile and magma fluxes from individual volcanoes and
879 from volcanic regions may greatly improve existing estimates of *intrusive/extrusive magma fluxes*
880 *and their link to tectonics*.

881

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885

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890

891 **Conflicts of interest/Competing interests**

892 There are no conflicts of interest to report.

893

894 **Availability of data and material**

895 All data generated in this paper through modeling are available in the Supplementary Material.

896

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1497

1498 **Figure captions**

1499

1500 **Fig. 1** Global distribution of open-vent volcanoes, listed in **Supplementary table S1**, encompassing a
1501 broad range of magma compositions and tectonic settings. ¹Average SO₂ flux (2005-2015) from Carn et al
1502 (2017)

1503

1504 **Fig. 2** Schematic illustration of a volcanic system to illustrate potential sources of outgassing volatiles.
1505 Magmas degas volatiles in response to (A) decompression during transit to the surface. The plot shows the
1506 concentrations of water (red) and carbon dioxide (blue) in a basaltic melt during decompression to the
1507 surface from a pressure of 200 MPa (modelled using MagmaSat; (Ghiorso and Gualda, 2015b)). The green
1508 shaded area shows the amount of exsolved volatile phase produced during degassing. In this case the
1509 basalt has a bulk concentration of 2 wt% H₂O and a range in CO₂ concentrations from 0.1 to 1 wt%. The
1510 exsolved volatile phase thus produced may outgas to the atmosphere during eruptions. Magmas also
1511 degas in response to (B) crystallisation in magma reservoirs in the crust. Crystallisation drives up the
1512 concentrations of volatiles in the residual melt and causes the formation of a substantial exsolved volatile
1513 phase after differentiation of the magma to highly evolved compositions. In this closed system degassing
1514 model involving crystallisation occurring at pressures of 80 to 350 MPa, a primitive basalt begins (at F=1)
1515 with 1 wt% H₂O and 0.1 wt% CO₂. After 50% crystallisation, the magma has reached basaltic andesite
1516 composition, and after 80%, approximately dacite composition

1517

1518 **Fig. 3 Conceptual frameworks to understand magmatic volatile phase (MVP) segregation from**
1519 **magmas in conduits (a, b, c) and in reservoirs (d, e, f) that may be relevant to open-vent volcanic**
1520 **systems.** In conduits, (a) convection is driven by density differences, with volatile-rich melts ascending,
1521 vesiculating, outgassing, then sinking (Kazahaya et al., 2004; Palma et al., 2011b; Shinohara et al., 1995;
1522 Stevenson and Blake, 1998). Crystals are generated by degassing-induced crystallisation in degassed,
1523 sinking melts (Beckett et al., 2014). (b) For open-vent volcanoes exhibiting strombolian activity, volcanic
1524 gases may accumulate in shallow crystal-rich plug made up of degassed and crystallised magma (Barth et
1525 al., 2019; Belien et al., 2010; Gurioli et al., 2014; Oppenheimer et al., 2015; Suckale et al., 2016;
1526 Woitischek et al., 2020); explosions may be caused by overpressure in the gas pockets overcoming the
1527 local yield strength in the crystal pack. (c) At low confining pressures and high magma viscosities, there
1528 may be sufficient strain at the conduit walls to induce brittle failure, with gas loss along permeable
1529 channels (e.g. Santiaguito, Mount St Helens 2004-2006) (Dingwell, 1996; Edmonds and Herd, 2007;
1530 Gonnermann and Manga, 2003; Tuffen and Dingwell, 2005). In crustal magma reservoirs, it has been
1531 proposed that the MVP may segregate under different regimes depending on magma crystal content. (d) in
1532 crystal-poor melt lenses the dominant regime may be buoyant bubble rise, the timescale for which is
1533 governed by the density difference between melt and MVP, the melt viscosity and the bubble size
1534 (Parmigiani et al., 2016). (e) In more crystal-rich mobile mushes, the MVP may rise buoyantly by viscous
1535 fingering, forming interconnected channels which may allow potentially much faster MVP segregation
1536 (Parmigiani et al., 2016). (f) In crystal-rich, melt-poor mush the MVP may become trapped in pore spaces,
1537 becoming mobilised once a critical overpressure is reached inside the pores, which may induce capillary
1538 fracturing (Degruyter et al., 2019; Parmigiani et al., 2016)

1539

1540 **Fig. 4** A plot of mean outgassing SO₂ flux (from Carn et al., 2017) against magma flux for a range of
1541 volcanoes. The magma flux required to supply the SO₂ flux is shown as a black square and the time-
1542 averaged eruption rate is shown as an open square. Estimates of degassing and erupted magma flux are
1543 sourced from: Ambrym (Allard et al., 2016); Manam (Liu et al., 2020a); Bagana (McCormick Kilbride et
1544 al., 2019; Wadge et al., 2018); Etna (Allard, 1997; Allard et al., 2006); Yasur (Métrich et al., 2011);
1545 Masaya (Zurek et al., 2019); and Stromboli (Allard et al., 2008). The individual studies use a combination
1546 of melt inclusion evidence and observed gas fluxes to infer the flux of degassing magma; and geological
1547 evidence to infer the magma eruption rate; please see papers for detail

1548

1549 **Fig. 5** A review of volcanic gas compositions. (a) H₂O–CO₂–S_t (i.e. SO₂ + H₂S) and (b) (b) H₂O–HCl–S_t
1550 (i.e. SO₂ + H₂S) gas compositions for a range of volcanoes, made up of both direct sampling (Fischer,
1551 2008; Hammouya et al., 1998; Symonds et al., 1994) and Multigas data (Aiuppa et al., 2008; Aiuppa et al.,
1552 2015; Burton et al., 2007; Sawyer et al., 2008; Shinohara and Witter, 2005). Red shaded regions = arc
1553 volcano emissions; blue = hydrothermal emissions; and yellow = intraplate/rift emissions

1554
1555 **Fig. 6** Models reconstructing the degassing of volatiles during decompression degassing and during
1556 isobaric crystallisation of alkali basalt melts at Yasur Volcano, Vanuatu. (a) Decompression of a
1557 trachybasalt from Yasur is accompanied by the exsolution of water and CO₂ (using Magmasat; (Ghiorso
1558 and Gualda, 2015a) from initial values of 1 wt% and 0.2 wt% respectively, based on melt inclusion and
1559 volcanic gas studies (Métrich et al., 2011; Woitischek et al., 2020). Fluid melt partition coefficients for Cl
1560 and S are shown in (i), melt volatile concentrations in (ii) and exsolved volatile phase composition in (iii).
1561 Open and closed system degassing models are considered, where open system degassing incorporates
1562 integration of the gas phase of the magma column at each pressure step. (b) Isobaric crystallisation leads
1563 to second boiling through enrichment of the melt in volatiles. Shown here are fluid-melt partition
1564 coefficients for Cl and S (i), melt volatile concentrations (ii) and the composition of the exsolved volatile
1565 phase (iii) for crystallization models (from a melt fraction, *F*, of 1 to a melt fraction of 0.1) at pressures of
1566 80, 160, 240 and 350 MPa. The pressures of crystallization are marked to show how the range in
1567 compositions links to pressure. Details of the models are given in **supplementary material**. Observed
1568 glass compositions (Métrich et al., 2011) and volcanic gas compositions (Oppenheimer et al., 2006;
1569 Woitischek et al., 2020) are marked on (a) and (b)

1570
1571 **Fig. 7** Composition of the exsolved volatile phase with pressure and melt fraction for (a) Yasur, Vanuatu
1572 and (b) Stromboli, Italy. Shown in solid black lines in both (a) and (b) are the decompressional degassing
1573 models for open and closed system degassing, marked with some of the pressure steps, in MPa. In dashed
1574 and red lines (see legend) are the isobaric second boiling models to describe the exsolved volatile phase
1575 produced during equilibrium crystallisation and degassing at various pressures, marked in red with the
1576 melt fraction remaining (1 to 0, where 1 is the case where there is no crystallisation, 0 denotes fully
1577 crystallised). A yellow box marks the compositions of volcanic gases observed at the surface (Aiuppa et
1578 al., 2010; Allard, 2010; Oppenheimer et al., 2006; Woitischek et al., 2020)

1579
1580 **Fig. 8** Mixing between a deep-derived MVP, generated through extensive second boiling, and an MVP
1581 derived from decompressional degassing may explain the gas systematics at Soufrière Hills Volcano,
1582 where high SO₂ fluxes and low S/Cl are observed during dome building and high SO₂ fluxes and high S/Cl
1583 during eruptive pauses. (a) the molar S/Cl of the MVP varies with S outgassed (in wt% of the
1584 melt+exsolved volatile phase) for isobaric degassing during second boiling and for decompressional
1585 degassing of S-poor rhyolite. Melt fraction remaining, *F*, is marked onto the trajectories for isobaric
1586 second boiling. Note the composition of the 'deep' MVP in equilibrium with rhyolite will differ if
1587 different bulk magma compositions of sulfur and chlorine are used, but the relative trends shown in the
1588 Fig. will remain the same. (b) The volcanic gas compositions at the surface may be explained well by a
1589 mixing model whereby a deep MVP generated through second boiling mixes with an MVP generated
1590 during decompressional degassing and crystallisation of sulfur-poor crystal-rich andesite (with a rhyolitic

1591 melt phase). Depending on the relative sizes of the two MVP reservoirs, the effect of mixing on the
1592 volcanic gas composition changes. For equal-sized reservoirs in terms of mass of the MVP phase per unit
1593 of magma, a scenario might be envisaged whereby during dome-building the shallow MVP dominates,
1594 generating Cl-rich gases; and during eruptive pauses (open-vent degassing) the deep MVP dominates,
1595 generating high S/Cl gases and a high SO₂ flux

1596

1597 **Fig. 9** Open-vent volcanoes are often in complex, extensional tectonic settings. (a) Masaya, Telica, Fuego
1598 and Pacaya, in the Central American Volcanic Arc, are closely located in regions of local crustal
1599 extension, associated with the Nicaraguan Depression (Masaya, Telica) and the rotational block tectonics
1600 of Guatemala (Fuego, Pacaya). (b) Ambrym is located at the boundary between a compressional and
1601 extensional regime in the Hebrides Arc. (c) Mount Etna, Italy, is located in an extension region stretching
1602 from Eastern Sicily to the south of Italy (see text for detail). Maps generated using GeoMapApp

1603

1604 **Fig. 10** Schematic diagram to illustrate the principal mechanisms of magma degassing at persistently
1605 active open-vent volcanoes. a) At basalt-dominated volcanoes, magmas rise to shallow storage regions in
1606 the crust to form shallow basic plutons. Some magma may rise and convect in the conduit. The exsolved
1607 volatile phase that outgasses quasi-continuously from the volcano is sourced from a mix of deep (second
1608 boiling) and shallow (convective degassing) sources. Volcanic activity at these volcanoes is dominated by
1609 gas-driven strombolian activity and paroxysms, and there may be a semi-stable lava lake. b) At andesite
1610 and dacite-dominated volcanoes, magmas undergo multi-level fractionation in the crust to form evolved
1611 melts which rise to shallow storage regions, exsolving a substantial exsolved volatile phase through
1612 second boiling. The persistent outgassing observed at these volcanoes is sourced principally from the
1613 second boiling process, which takes place during the solidification of hybrid and felsic plutons at depth. In
1614 both cases, magma intrusion and open-vent outgassing is promoted by crustal extension, which provides
1615 accommodation space for magma intrusion at depth and for the gravitational segregation of lower density
1616 exsolved volatile phases to the upper parts of the storage region. Deep generation of superhydrous melts
1617 may advect volatiles up to subvolcanic reservoirs