

1 *40-character title: Lake Chala: 250-ky record of basaltic volcanism*

2 **History of scoria-cone eruptions on the eastern shoulder of the Kenya-Tanzania Rift**  
3 **revealed in the 250-ky sediment record of Lake Chala near Mt. Kilimanjaro**

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12 **Abstract**

13 Reconstructions of the timing and frequency of past eruptions are important to assess the  
14 propensity for future volcanic activity, yet in volcanic areas such as the East African Rift only  
15 piecemeal eruption histories exist. Understanding the volcanic history of scoria cone fields,  
16 where eruptions are often infrequent and deposits strongly weathered, is particularly  
17 challenging. Here we reconstruct a history of volcanism from scoria cones situated along the  
18 eastern shoulders of the Kenya-Tanzania Rift, using a sequence of tephra (volcanic ash)  
19 layers preserved in the ~250-ky sediment record of Lake Chala near Mt. Kilimanjaro. Seven  
20 visible and two non-visible (crypto-) tephra layers in the Lake Chala sequence are attributed  
21 to activity from the Mt. Kilimanjaro (northern Tanzania) and the Chyulu Hills (southern Kenya)  
22 volcanic fields, on the basis of their glass chemistry, textural characteristics and known  
23 eruption chronology. The Lake Chala record of eruptions from scoria cones in the Chyulu  
24 Hills volcanic field confirms geological and historical evidence of its recent activity, and  
25 provides first-order age estimates for seven previously unknown eruptions. Long and well-  
26 resolved sedimentary records such as that of Lake Chala have significant potential for  
27 resolving regional eruption chronologies spanning hundreds of thousands of years.

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30 Chyulu Hills volcanic field; tephra glass geochemistry

## 31 **1. Introduction**

### 32 **1.1. Volcanism in East Africa**

33 The East African Rift system (EARS) marks one of Earth's best-preserved continental rift  
34 systems - a fascinating natural laboratory in which to study compositionally diverse volcanism  
35 and wide-ranging volcanic hazards. Volcanism in Kenya and Tanzania tracks the eastern  
36 branch of the EARS, where rifting was initiated ~35 Ma in the Lokichar Basin of northern  
37 Kenya (Fig. 1; Macgregor, 2015). Twenty-one volcanoes in Kenya and 10 in Tanzania are  
38 believed to have been active during the Holocene, on the basis of their morphology and  
39 evidence of historical or recent geological activity (Global Volcanism Program, 2013). Whilst  
40 there is little data on past volcanism, an increasing number of volcanological, archaeological  
41 and palaeoenvironmental studies are beginning to compile and date eruptive deposits  
42 preserved both in terrestrial and lake-sediment sequences throughout East Africa (e.g. Poppe  
43 *et al.*, 2016; Campisano *et al.*, 2017; Fontijn *et al.*, 2018, McNamara *et al.*, 2018). These  
44 tephrostratigraphic studies indicate that many volcanoes of the Kenya-Tanzania Rift have  
45 erupted explosively during the Holocene, depositing ash over hundreds of kilometres (Fontijn  
46 *et al.*, 2010, 2018; Martin-Jones *et al.*, 2017; 2018; Lane *et al.*, 2018; McNamara *et al.*, 2018).

47 During the Late Pleistocene, volcanism within the EARS has been strongly bimodal. Silicic  
48 eruptions from central volcanoes occurred alongside less abundant basaltic eruptions from  
49 fissure vents – whilst intermediate compositions have been comparatively rare (Baker, 1987;  
50 Williams, 1984; MacDonald, 1995; McDoughall and Brown, 2009). The volcanoes of the  
51 Kenya-Tanzania Rift have been divided into northern and southern sectors by Williams *et al.*  
52 (1978) and Baker (1987). In the North, Emurangogolak, Silali, Korosi and Ol Kokwe (Fig. 1)  
53 have erupted large volumes of basalts characterised by sparse pyroclastic activity (White *et al.*  
54 *et al.*, 2012). In the southern Kenya Rift, basaltic eruptions have been typically minor but have  
55 occurred in the Ndabidi, Tandamara and Elementeita lava fields adjacent to the Olkaria,  
56 Suswa and Eburru calderas (Fig. 1; White *et al.*, 2012). Contemporary volcanism has also  
57 occurred to the East of the Rift, with Mt. Kenya, Mt. Meru and Mt. Kilimanjaro having erupted  
58 basalts alongside more differentiated nephelinite and phonolite compositions (Baker, 1987).

### 59 **1.2 Strombolian activity in the EARS**

60 Scoria (or cinder) cones are one of the most common volcanic landforms on Earth. They are  
61 characterised by Strombolian explosions that eject predominantly basaltic lapilli and ash tens  
62 to hundreds of meters into the air (Vespermann and Schminke, 2000; Fodor and Brož, 2015).  
63 Scoria cones are often described as parasitic, having formed due to eruptions along fissures  
64 on the flanks of a volcano rather than from the main vent. Hundreds of such cones form  
65 volcanic fields, which may be active over millions of years (Valentine and Connor, 2015) with  
66 average eruption recurrence intervals of hundreds to thousands of years (Ort *et al.*, 2008).  
67 Scoria cones are typically monogenetic, *i.e.* formed by a single episode of volcanism usually  
68 lasting days to weeks, although the more complex polygenetic character of some cones is  
69 now being recognised (*e.g.*, Poppe *et al.*, 2016). The eruption styles and landforms within  
70 volcanic fields are often variable: if magma interacts with water, explosive (phreatomagmatic)  
71 activity may form maar crater lakes and tuff cones, whereas if low-viscosity magma is erupted  
72 as a fire fountain, spatter cones of welded, ropey lava fragments may develop.

73 In equatorial East Africa, small basaltic cones occur along the flanks of both Mt. Meru and  
74 Mt. Kilimanjaro. Those on Mt. Meru are dated to 1.71 +/- 0.06 Ma (Wilkinson *et al.*, 1986),  
75 whereas those on Mt. Kilimanjaro are more recent and dated to 200-150 ka (Nonnotte *et al.*,  
76 2008, 2011). Roughly 100 km to the east of the Kenya Rift, hundreds of basaltic scoria cones  
77 form the Hurri, Nyambeni and Chyulu Hills volcanic fields. Chyulu, the youngest and most  
78 southerly of these, hosts a minimum of 250 cones (as estimated from satellite imagery).  
79 Activity here is thought to have commenced around 1.4 Ma, and the young Shaitani and  
80 Chainu cones (Fig. 1) erupted as recently as the mid-19<sup>th</sup> century CE (Saggerson, 1963;  
81 Goles, 1975; Haug and Strecker, 1995; Späth *et al.*, 2000; 2001). Detailed studies (Haug and  
82 Strecker, 1995; Späth *et al.*, 2000; 2001) have resolved long-term trends in the volcanic  
83 history of the Chyulu Hills, but focused on dating effusive rather than explosive eruptions.  
84 This study provides the first detailed reconstruction of scoria-cone eruptions from volcanic  
85 fields on the shoulders of the Kenya-Tanzania Rift, including the Chyulu Hills.

### 86 **1.3 Scoria cone eruptions recorded in lake sediments**

87 Reconstruction of scoria-cone eruption histories is challenged by their association with lava  
88 flows and spatially-restricted scoria deposits, meaning that the imprint of past scoria-cone  
89 eruptions may be unrecognisable. However, these eruptions are capable of generating high  
90 eruption columns and extensive tephra fallout (*e.g.* Pioli *et al.*, 2008; Kawabata *et al.*, 2015;  
91 Jordan *et al.*, 2016). Scoria is particularly unstable and weathers readily, especially in tropical

92 climates. Long inter-eruptive periods may therefore be associated with complex stratigraphic  
93 relationships, and deposits from older centres may be almost entirely denuded while new  
94 cones are developing (Németh *et al.*, 2012, 2014). The extent of scoria weathering and  
95 shape of the (remnant) cone may provide relative age control but, with datable minerals and  
96 charcoal often lacking, determining the absolute age of past eruptions can be difficult  
97 (Jaimes-Viera *et al.*, 2018; Nieto-Torres *et al.*, 2019).

98 Sedimentation in many lakes is continuous and uniform over long periods, providing a better-  
99 resolved and potentially more complete archive of regional volcanic history. Tephra layers  
100 preserved in lake sequences can often be dated using age-depth models built from <sup>14</sup>C dating  
101 of organic components – giving an indication of past eruption timing. Where comparative  
102 geochemical data is available for nearby volcanoes, and they erupt distinct magma  
103 compositions, the geochemistry of the glass shards composing the tephra can be used to  
104 identify its source. Tephra sequences in lake sediments have been widely used to constrain  
105 volcanic histories, typically from voluminous volcanic eruptions (*e.g.*, Lane *et al.*, 2017).  
106 However, several studies have also used lake sediment tephrostratigraphy to unravel the  
107 complex history of basaltic cones (*e.g.*, Green *et al.*, 2014; Hopkins *et al.*, 2015; 2017;  
108 Németh *et al.*, 2008). Studies on East African lake sequences (Martin-Jones *et al.*, 2017a; b;  
109 McNamara *et al.*, 2018; Lane *et al.*, 2018) are beginning to shed light on EARS volcanism,  
110 revealing the frequency of past, sometimes undocumented, eruptions. When these records  
111 can be linked with their on-land equivalents, they can also give an indication of eruption size.

112 In this study we present an initial tephrostratigraphic record of scoria-cone eruptions from  
113 volcanic fields in southern Kenya and northern Tanzania, derived from a ~250-ky sediment  
114 sequence from Lake Chala, a deep crater lake bridging the Kenya-Tanzania border on the  
115 south-eastern flank of Mt. Kilimanjaro. Lake Chala is relatively unique in having permanently  
116 stratified and anoxic bottom waters (Buckles *et al.*, 2014), providing excellent preservation  
117 conditions for even very fine tephra layers from relatively modest and/or distant eruptions.

#### 118 **1.4 Lake Chala**

119 Lake Chala (3.3° S; 37.7° E) is situated in a steep-sided volcanic crater basin formed after a  
120 monogenetic parasitic eruption associated with Mt. Kilimanjaro's most recent phase of  
121 volcanic activity. Based on the basal age of 25 ka for the upper, <sup>14</sup>C-dated, 21.65 metres of  
122 profundal sediments (Blaauw *et al.*, 2011) and an extrapolated age model for the last 140

123 kyrs (Moernaut *et al.* 2010), Verschuren *et al.* (2013) estimated the age of the complete ~210-  
124 metre sediment infill in Lake Chala at >250,000 years. The lake is 94 m deep with a surface  
125 area of 4.2 km<sup>2</sup>, and receives water only from direct rainfall, runoff from its steep crater walls  
126 and subsurface inflow originating from the percolation of rainfall onto the forest zone on Mt.  
127 Kilimanjaro (Verschuren *et al.*, 2009). Its lower water column is permanently stratified, both  
128 today (Buckles *et al.* 2014; Wolff *et al.*, 2014) and over at least the last 25,000 years  
129 (Verschuren *et al.* 2009). Consequently, Lake Chala accumulates undisturbed and near-  
130 continuously laminated diatomaceous organic sediments, interrupted only by turbidites and  
131 tephtras (Verschuren *et al.*, 2013).

132 With the aim of reconstructing the palaeoclimate and landscape history of equatorial East  
133 Africa, two sediment sequences have been extracted from Lake Chala by the CHALLACEA  
134 and DeepCHALLA projects. The 21.65-m long composite CHALLACEA sequence was  
135 constructed from four overlapping piston cores collected in 2003 and 2005, and provided  
136 insights into climate dynamics spanning the last 25 kyr (*e.g.*, Verschuren *et al.*, 2009; Barker  
137 *et al.*, 2011; Sinninghe Damsté *et al.*, 2011). In November 2016, as part of the International  
138 Continental Scientific Drilling Program, the DeepCHALLA project (Verschuren *et al.*, 2013)  
139 retrieved a 215-m long sediment sequence, nearly reaching the crater floor. This study  
140 utilises both the CHALLACEA and DeepCHALLA sequences to explore its record of basaltic  
141 scoria-cone volcanism in the Mt. Kilimanjaro area.

142 Figure 1

## 143 **2. Methods**

### 144 **2.1 Locating and analysing tephtras**

145 We identified and recorded basaltic tephtra layers by visual inspection throughout the ~250-  
146 kyrs DeepCHALLA sequence, and documented their morphological features and crystal  
147 content under high-power microscopy. No visible tephtra layers were detected in the  
148 CHALLACEA record, therefore this sequence was analysed for the presence of crypto-tephtra  
149 using a method adapted from Blockley *et al.* (2005) and guided by peaks in XRF data (Kristen,  
150 2007) indicating high concentrations of glass shards. Contiguous 1-cm samples were  
151 collected from the 10-cm sediment interval surrounding an XRF peak, dried at 95 °C, wet-  
152 sieved to >25 µm and then density-separated using a sodium polytungstate solution of >1.95  
153 g/cm<sup>3</sup> to remove organic material. Extracted residues were counted under transmitted light

154 microscopy and counts plotted against core stratigraphy to determine tephra isochrons (*i.e.*,  
155 the depth of peak shard density), from which shards were then sampled for geochemical  
156 analysis. For this purpose, both macro- and microscopic tephtras were re-extracted from the  
157 sediment cores and sieved. Glass shards were mounted in epoxy resin stubs, which were  
158 then ground and polished to expose the glass for analysis.

159 Each tephra layer was assigned either a CHALLACEA (CH-) or DeepCHALLA (DCH-) code  
160 referring to the depth (in metres) of its basal contact according to the composite depth scale  
161 of both sequences, which corrects for core gaps and drilling-related artefacts but not for event  
162 deposits such as turbidites and tephra layers. Indicative ages (rounded to the nearest 0.1 ky  
163 for the CHALLACEA tephtras and 1 ky for the DeepCHALLA tephtras) were assigned to each  
164 tephra using the CHALLACEA <sup>14</sup>C-based chronology for the last 25 kyrs (CH-3.68 and CH-  
165 13.22; Blaauw *et al.*, 2011); links between Lake Chala seismic stratigraphy and known near-  
166 global climate events back to 140 ka (DCH-57.60, DCH-58.03, DCH-58.41, DCH-69.24,  
167 DCH-75.53, DCH-111.72 ; Moernaut *et al.*, 2010); and extrapolation of the average  
168 sedimentation rate over this 140-ky interval (0.80 m/ky) to the base of the DeepCHALLA  
169 sequence (DCH-228.60; this study). At this time, the associated 95% age uncertainties are  
170 estimated to be  $\pm 0.1$  ky (2.4%) for CH-3.68 and  $\pm 0.45$  ky (2.7%) for CH-13.22 (Blaauw *et al.*,  
171 2011);  $\pm 2$  ky (1.6-3%) for the six DeepCHALLA tephtras <140 ka (Moernaut *et al.*, 2010); and  
172  $\pm 10$  ky (4%) for DCH-228.60.

## 173 **2.2. Tephra geochemical analyses**

174 Individual glass shards from within all nine basaltic tephra layers, except CH-13.22, were  
175 analysed using a wavelength-dispersive Cameca SX-100 electron microprobe (WDS-EPMA)  
176 at the Department of Earth Sciences, University of Cambridge. CH-13.22 was analysed using  
177 a Jeol 8600 WDS-EPMA at the Research Laboratory for Archaeology and the History of Art,  
178 University of Oxford. Operating conditions were identical for both instruments, using a 10- $\mu$ m  
179 diameter defocussed beam, a 15-kV accelerating voltage and 6-nA beam current. Sodium  
180 was collected for 10 s, Cl and P for 60 s and all other elements for 30 s. The instruments  
181 were calibrated against a series of mineral standards, and results were quantified using the  
182 PAP absorption correction method. Assays of the MPI-DING standards KL2-G (basalt) and  
183 St-Hs6/80-G (andesite) (Jochum *et al.*, 2006) and an in-house Lipari obsidian (peralkaline  
184 rhyolite) were used to monitor accuracy across the two instruments. All major element data

185 presented in the text, tables and figures are normalised to 100% to account for variable  
186 secondary hydration of glasses. For raw data, see the Supplementary Material.

187 Trace element concentrations in individual glass shards were determined using a Thermo  
188 Scientific iCapQ coupled to a Teledyne G2 Eximer laser in the iCRAG laboratory at Trinity  
189 College Dublin. Depending on available sample areas, analyses used 20  $\mu\text{m}^2$ , 30  $\mu\text{m}^2$  or 40  
190  $\mu\text{m}^2$  laser spots. The laser was fired at a repetition rate of 5 Hz, with a count time of 40  
191 seconds on both sample and gas blank. Concentrations were calculated via calibration  
192 against NIST612 and using  $^{29}\text{Si}$  as the internal standard (concentrations determined via  
193 EPMA). Data reduction was performed in Lolite3.4, followed by a secondary correction using  
194 Ca as advocated in Tomlinson *et al.* (2010). To monitor instrument precision and accuracy,  
195 the MPI-DING reference materials GOR132-G (komatiite), St-Hs6/80-G (dacite) and ATHO-  
196 G (rhyolite) were analysed during each analytical run; see Supplementary Material for details.

### 197 **2.3. Linking tephra to possible source regions**

198 This study presents new glass data for nine tephra layers in Lake Chala sediments alongside  
199 published compositions of Kenya-Tanzania Rift volcanoes in order to identify their potential  
200 source volcanoes. When collating the reference dataset, we focussed on compositions of  
201 eruptions within the <250-ky time frame of the DeepCHALLA record. However, published  
202 compositional and chronological data from many EARS volcanoes are limited. Further,  
203 available compositions are frequently whole-rock data, which contain variable amounts of  
204 phenocrysts and hence preclude direct comparison of element concentrations (e.g. Smith *et al.*,  
205 2005; Hopkins *et al.*, 2017; Pearce *et al.*, 2019). We therefore use a total alkali silica  
206 (TAS) plot (Le Bas *et al.*, 1986) to draw broad comparisons between eruptive compositions,  
207 and ratios of incompatible elements in glass or whole-rock samples to differentiate between  
208 eruption centres. Incompatible elements (e.g., Zr, Nb, Ba, La, and Pr) increase in abundance  
209 with magmatic differentiation, yet their ratios typically remain unchanged and can therefore  
210 be used as the chemical signature of an individual volcano. MacDonald (1987), Scott and  
211 Skilling (1999) and White *et al.* (2012) used Nb/Zr trends to discriminate among Kenyan Rift  
212 volcanoes, which are however more differentiated than the basaltic tephra in this study.

## 213 **Results**

### 214 **3.1 Lake Chala tephrostratigraphy**

215 This study focuses on nine basaltic tephra in the CHALLACEA and DeepCHALLA  
216 sequences (Table 1), which provide rare examples of far-travelled tephra from scoria-cone  
217 eruptions recorded in a lake-sediment archive. Seven of these tephra are macroscopically  
218 visible horizons in the DeepCHALLA record, containing light- to dark-brown scoriaceous,  
219 bubble-wall and ropey (>1000  $\mu\text{m}$  in length) shard morphologies (Fig. 2) indicative of  
220 fragmentation through Strombolian-style eruption. Olivine and clinopyroxene crystals  
221 frequently occur both as phenocrysts and free crystals (Fig. 2). The glass shards are  
222 predominantly of alkaline basaltic composition, however one tephra at the base of the  
223 sequence contains foiditic (low  $\text{SiO}_2$ ) glass (Fig. 3).

224 Cryptotephra investigation of the <25-ky sediments in the CHALLACEA sequence reveals  
225 regional volcanic activity into the Holocene. Background counts of glass shards are low,  
226 indicating that peaks in shard counts represent primary tephra deposition with little  
227 contamination from volcanic deposits forming the inner the crater walls. Two cryptotephra  
228 horizons (a modest peak of 100 shards/g dry sediment at 13.22 m depth and a more  
229 pronounced peak of 2230 shards/g dry sediment at 3.68 m depth) contain brown glass shards  
230 of alkaline basaltic composition (Fig. 3) similar to those in the DeepCHALLA sediments, and  
231 are therefore included in this study. For simplicity, these alkaline basaltic and foiditic tephra  
232 are hereafter referred to as mafic tephra.

233 Indicative ages for the nine mafic eruptions recorded in Lake Chala sediments are given in  
234 Table 1. The oldest tephra, found at the base of the DeepCHALLA sequence is dated to ~248  
235 ka. All other investigated tephra were deposited during the last ~134 kyrs, with DCH-57.60,  
236 DCH-58.03 and DCH-58.41 deposited over a relatively short period between ~66 and ~65  
237 kyrs. The youngest eruptions, recorded as cryptotephra in the CHALLACEA sequence, are  
238 dated to ~16.8 and ~4.2 ka.

239 Table 1

240 Figure 2

### 241 **3.3 Composition of the mafic tephra in Lake Chala**

242 Single-shard major and minor element concentrations are summarized in Table 2 and Figs.  
243 3-5. Glass compositions and incompatible-element ratios permit discrimination between the  
244 individual tephra, as described below.

245 3.3.1 *DeepCHALLA mafic tephtras*

246 The oldest of the mafic tephtra layers in Lake Chala sediments (DCH-228.60; ~248 ka) occurs  
247 only 0.55 m above the base of the DeepCHALLA drill core. Glass shards within DCH-228.60  
248 are foiditic and enriched in alkalis and the incompatible elements Zr, Nb, Ba, La, Pr and Nd;  
249 and depleted in MgO and TiO<sub>2</sub> relative to the other mafic tephtras (Fig. 3).

250 Tephtra DCH-228.60 as well as DCH-111.72 (~134 ka) and DCH-75.53 (~87 ka) show wide  
251 intra-eruptive variation in CaO content, distinguishing them from the more tightly constrained  
252 CaO concentrations in the six other tephtras. Trends of Ba/La in DCH-111.72 and DCH-75.53  
253 glass shards are distinct from one another and from those in the other mafic tephtras. Glass  
254 shards in DCH-111.72 can be readily distinguished from DCH-75.53, because the latter are  
255 enriched in TiO<sub>2</sub>, MgO, FeO\* and CaO; and depleted in SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, the alkalis and Ba (Table  
256 2, Figs. 4-5).

257 The four younger macroscopic mafic tephtras derive from eruptions dated to between ~79  
258 and ~65 kyrs, and occupy a similar area in the incompatible-element plots. Tephtra DCH-  
259 69.24 (~79 ka) contains higher MgO and lower Al<sub>2</sub>O<sub>3</sub> concentrations than DCH-58.41, DCH-  
260 58.03 and DCH-57.60 (~66-65 ka; Fig. 4). The latter three have similar major-element  
261 compositions, except that DCH-58.41 glass is enriched in MgO and depleted in Al<sub>2</sub>O<sub>3</sub> relative  
262 to DCH-58.03; and that DCH-57.60 glass contains lower concentrations of incompatible trace  
263 elements (i.e. Zr, Nb, Ba, La, Pr) than all other tephtras analysed.

264 3.3.2 *CHALLACEA mafic cryptotephtras*

265 Brown glass shards in the CHALLACEA cryptotephtras have major-element compositions that  
266 are most similar to the four youngest (~79–65 ky) visible tephtras in the DeepCHALLA  
267 sequence (Figs. 3-4). The two cryptotephtras cannot be distinguished from one another, as  
268 their composition is wide-ranging, each encompassing the entire range of MgO  
269 concentrations observed in the ~79–65 ky DeepCHALLA tephtras. The cryptotephtras do,  
270 however, contain slightly higher CaO and lower Al<sub>2</sub>O<sub>3</sub> than the two youngest DeepCHALLA  
271 mafic tephtras (Fig. 4).

272 Table 2

273 Figure 3

274 Figure 4

275 Figure 5

## 276 **4. Discussion**

### 277 **4.1 Multi-criteria correlations**

#### 278 *4.1.1. Major-, minor- and trace-element compositions*

279 The mafic tephtras in Lake Chala are more alkaline than basalts erupted from the Kenya Rift  
280 volcanoes (Fig. 3), and more closely resemble basalts erupted from the off-Rift volcanic  
281 centres of Mt. Kilimanjaro and the Chyulu Hills. The seven mafic DeepCHALLA tephtras for  
282 which we have trace-element data show distinct linear arrays in plots of Ba/La, Pr/La and  
283 Zr/Nb (Fig. 5). The four younger tephtras (DCH-69.24, DCH-58.41, DCH-58.03 and DCH-  
284 57.60) plot tightly along the same fractionation trends, whereas the three oldest tephtras  
285 (DCH-228.60, DCH-111.72 and DCH-75.53) display incompatible-element trends that are  
286 distinct both from each other and from those younger tephtras (Fig. 5).

287 Incompatible-element ratios of all mafic Chala tephtras are broadly consistent with those of  
288 bulk alkaline lavas from the Chyulu Hills (Späth *et al.*, 2001) and Mt. Kilimanjaro's parasitic  
289 cones (Nonnotte *et al.*, 2011), which also plot within relatively well-defined values, particularly  
290 for Pr/La. Specifically, in plots of Pr/La and Zr/Nb, the four youngest mafic DeepCHALLA  
291 tephtras track the same linear path as the Chyulu lavas, indicating that this is their likely  
292 source. Also tephtra DCH-75.53 shows broadly similar incompatible-element ratios as the  
293 Chyulu lavas in terms of Pr/La and Zr/Nb, but is distinct in terms of Ba/La. However, since  
294 the latter ratio is less tightly constrained for Chyulu lavas, we tentatively suggest the Chyulu  
295 Hills is the most likely source for DCH-75.53. Without trace-element data we are unable to  
296 directly relate the <25-ky mafic cryptotephtras to Chyulu or Mt. Kilimanjaro volcanism,  
297 notwithstanding their similarity in major-element composition to the ~79-65 ka tephtras.

298 In contrast, the foiditic composition of tephtra DCH-228.60 is more similar to the high alkalinity  
299 of parasitic volcanic activity on Mt. Kilimanjaro (Fig. 3). Both DCH-228.60 and DCH-111.72  
300 follow a Pr/La trend similar to known Mt. Kilimanjaro tephtras (Fig. 5) and we thus tentatively  
301 relate these two older mafic tephtras to the later phase of Mt. Kilimanjaro volcanism as  
302 recognised by Nonnotte *et al.* (2011).

303 The compositional distinctions described above are subtle and it is therefore difficult, based  
304 on geochemistry alone, to confidently assign individual tephra layers to Mt. Kilimanjaro or the  
305 Chyulu Hills. Trends in published incompatible-element ratios for the Mt. Kilimanjaro and  
306 Chyulu lavas are not as tightly constrained as those in our glass data, and it is possible that  
307 only Pr and La, ratios of which are consistent between whole-rock samples and glass, behave  
308 as truly incompatible elements. The magmatic evolution of these two volcanic fields is likely  
309 complex, as both are fed by small magma batches derived from variable degrees of partial  
310 melting and crystal fractionation. The glass compositions of DCH-69.24, DCH-58.41, DCH-  
311 58.03 and DCH-57.60 follow crystallisation paths characteristic for olivine and clinopyroxene  
312 (Fig. 5). However, fluctuating concentrations of incompatible trace elements between them  
313 suggest a more complex magma evolution, possibly involving multiple influxes of fresh  
314 magma.

315 Haug and Strecker (1995) and Class *et al.* (1994) report both temporal and spatial variation  
316 in the geochemistry of lavas from the Chyulu Hills, with those erupted from cones in the south  
317 of the volcanic field enriched in SiO<sub>2</sub> compared to older northern lavas. Späth *et al.* (2000)  
318 attribute this change to the degree of partial melting increasing towards the South. Such  
319 differences in melting can create heterogeneous magma compositions across a single  
320 volcanic field, challenging comparisons using incompatible element ratios. Also, our  
321 understanding of compositional variation in magmas from the Chyulu volcanic fields is based  
322 only on studies of lava flows, which were erupted under different conditions, and likely at  
323 different times, to the tephtras generated by explosive eruptions we observe in Lake Chala -  
324 further complicating direct comparisons.

#### 325 *4.1.2. Constraints from known eruption histories*

326 Based on the compositional similarities discussed above, the two CHALLACEA cryptotephtras  
327 and five youngest DeepCHALLA tephtras (DCH-75.53, DCH-69.24, DCH-58.41, DCH-58.03  
328 and DCH-57.60) are attributed to previously undefined eruptions of the Chyulu Hills, whilst  
329 DCH-228.60 and DCH-111.72 are tentatively correlated to Mt. Kilimanjaro. Further support  
330 for these proposed attributions is found in the known eruption history of these volcanic fields.

331 The Chyulu Hills is unique in the EARS for its prevalent Holocene volcanism (Saggerson,  
332 1963; Haug and Strecker, 1995; Späth *et al.*, 2000; 2001). Lavas decrease in age from the  
333 northwest (~1.4 Ma) towards the southeast, where a lava flow East of Mzima Springs (Fig.

334 1b) overlies lake sediments radiocarbon-dated to ~26 ka (Späth *et al.*, 2000). Omenge and  
335 Okele (1992) provide oral accounts of <100-year-old flows from the Shaitani and Chainu  
336 cones in the South (Fig. 1). The Chyulu Hills is therefore younger than the most recent  
337 parasitic activity along the southeast flank of Mt. Kilimanjaro, which is dated to ~200-150 ka  
338 (Nonnotte *et al.*, 2008). Concordant with this older age, the Mt. Kilimanjaro scoria cones are  
339 typically more denuded and more densely vegetated than the Chyulu cones. The age of the  
340 two oldest (~248 ka and ~134 ka) mafic DeepCHALLA tephras is thus in line with the known  
341 timing of the final phases of volcanism on Mt. Kilimanjaro; and the prevalence of younger  
342 volcanic activity in the Chyulu Hills support our geochemical correlation of the two <25-ky  
343 CHALLACEA cryptotephras (4.2 and 16.8 ka) and five <100-ky DeepCHALLA tephras  
344 (between ~65 and ~87 ka) to that volcanic field. It should be noted, however, that the known  
345 eruption histories of Mt. Kilimanjaro and Chyulu are based on K/Ar dating of lavas collected  
346 from select areas of these volcanic fields, and hence may not capture all phases of their  
347 activity. Moreover, a complete cryptotephra scan of the long DeepCHALLA sediment record  
348 would likely document additional Chyulu Hills and Mt. Kilimanjaro eruptions.

#### 349 *4.1.3. Indications of source proximity*

350 Mt. Kilimanjaro and the Chyulu Hills, the volcanic centres closest to Lake Chala that erupt  
351 basalts, are located respectively <1-30 km to the west, and 60-120 km to the northeast (Fig.  
352 1). Tephra deposits typically become thinner and finer-grained with increasing distance from  
353 their source. Consistent with our source attributions, DCH-228.60 and DCH-111.72 contain  
354 broadly coarser (~300 µm) free crystals than the seven younger tephras (<100 µm; Table 1).  
355 Assuming modern wind patterns over the region, dispersal of tephra from the Chyulu Hills  
356 towards Lake Chala would vary seasonally, being promoted from November to April by  
357 northeasterly winds, and hampered from May to October by southeasterly winds (Wolff *et al.*,  
358 2014).

#### 359 **4.2 Scoria-cone eruption records from lakes: potential and challenges**

360 This study highlights the exceptional preservation conditions offered by lake systems such  
361 as Lake Chala, and their value in documenting eruptions otherwise unrecognised in the  
362 geological record. Tying the tephras incorporated in such lake records to their source is  
363 mainly challenged by the scarcity of published glass data characterising individual EARS  
364 volcanoes; this problem is amplified for intermittently active volcanic fields generating

365 localised and rapidly-weathered basaltic scoria deposits. This study uses incompatible-  
366 element ratios to compare glass from mafic Chala tephras with published whole-rock  
367 analyses of lavas allowing their correlation to the Chyulu Hills and Mt. Kilimanjaro volcanic  
368 fields. Our geochemical correlations remain tentative however, as they are substantiated only  
369 by examining the Lake Chala tephrostratigraphy against existing, but skeletal, stratigraphic  
370 and geochronological outcrop data.

371 Further, whereas geochemical and chronological constraints allowed us to trace the mafic  
372 Chala tephras to two volcanic fields, correlating them to individual scoria cones is more  
373 challenging. As historical (<100 years) eruptions have been documented from the southerly  
374 Shaitani and Chainu cones in the Chyulu Hills, these, or cones nearby, are a likely source of  
375 the most recent (>4.2 ky) mafic cryptotephra recorded in Lake Chala. Späth et al. (2001)  
376 report that many of the Chyulu Hills cones, particularly from the southern sector, cannot be  
377 distinguished from one another based on their petrographic and geochemical characteristics.  
378 These difficulties call for more detailed study of the Chyulu Hills and Mt. Kilimanjaro volcanic  
379 fields, focussing not only on the lava flows but also past pyroclastic activity. Such  
380 compositional and chronological analysis should target the full temporal and spatial extent of  
381 these volcanic fields, allowing a detailed understanding of their eruptive history and magma  
382 genesis. Long sediment sequences, such as that of Lake Chala, have great potential to feed  
383 into our understanding of complex monogenetic volcanic fields, yet this potential can only be  
384 fully realised in the presence of detailed comparative eruption data from on-land records.

## 385 **5. Conclusions**

386 Mafic tephras preserved in Lake Chala sediments provide a new record of volcanic activity in  
387 the Chyulu Hills and Mt. Kilimanjaro volcanic fields spanning ~250,000 years. Excellent  
388 preservation conditions reveal tephra deposits from distant, and previously undocumented,  
389 scoria-cone eruptions. We tentatively attribute the two oldest mafic Chala tephras (dated to  
390 ~248 and ~134 ka) to Mt. Kilimanjaro, and seven younger tephras (dated to between ~87 and  
391 4.2 kyrs) to volcanism in the Chyulu Hills. We show that pyroclastic activity has occurred in  
392 the Chyulu Hills over at least the last ~87 kyrs, and the ash deposited within the late-Holocene  
393 section of the Chala sequence supports the documented evidence of more recent volcanism.  
394 Pyroclastic activity will likely continue to be a common feature of future volcanism in the  
395 Chyulu Hills, however detailed research into the volcanological hazard and associated socio-  
396 cultural implications is needed to evaluate the extent of any risks to the local population.

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408

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591 **Table 1.** Identification code, location in the sediment sequence (core section and depth, at  
592 base of deposit), physical properties and age of the mafic DeepCHALLA tephtras (DCH-) and  
593 CHALLACEA cryptotephtras (CH-) analysed in this study. Minerals are abbreviated as cpx  
594 (clinopyroxene), ol (olivine) and fsp (feldspar). The listed ages are based on <sup>14</sup>C dating for  
595 the two CHALLACEA cryptotephtras (Blaauw et al., 2011), and an extrapolated age model  
596 tied to seismic stratigraphy (Moernaut et al., 2010) for the seven DeepCHALLA tephtras. Age  
597 uncertainties are 2.4-2.7% for CH tephtras and 1.6-4% for DCH tephtras; see text.

<i>Tephra ID</i>	<i>Core section</i>	<i>Section depth</i>	<i>Composite depth (m)</i>	<i>Thickness (cm)</i>	<i>Glass size</i>	<i>Crystals</i>	<i>Crystal size (µm)</i>	<i>Age (ka)</i>
CH-3.68	2PII	45.4	3.68	-	<100	trace	<50	4.2
CH-13.22	3P-Va	104.6	13.22	-	<50	trace	<50	16.8
DCH-57.60	1B-18H-2	56.5	57.6	0.5	<1000	cpx, ol	<100	~65
DCH-58.03	1B-18H-2	99.5	58.03	0.5	<1000	cpx, ol	<100	~65
DCH-58.41	1B-18H-3	21.5	58.41	0.3	<1000	cpx, ol	<100	~66
DCH-69.24	1A-23H-2	23.3	69.24	0.3	<100	trace	<100	~79
DCH-75.53	1B-23H-3	51.8	75.53	0.3	<250	trace	<100	~87
DCH-111.72	1B-35H-1	34	111.72	0.5	<600	fsp, cpx	<300	~134
DCH-228.60	1E-45E-3	122.5	228.6	1	<1000	fsp, cpx	<250	~248

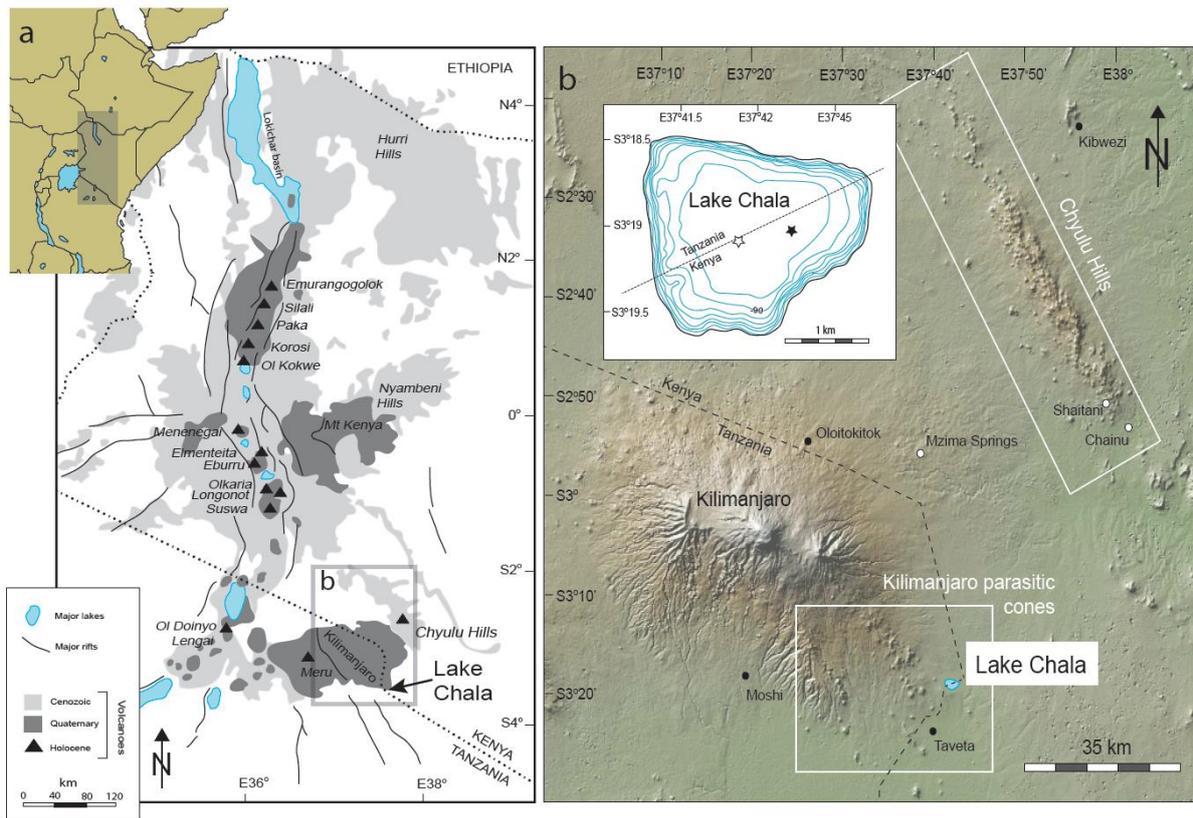
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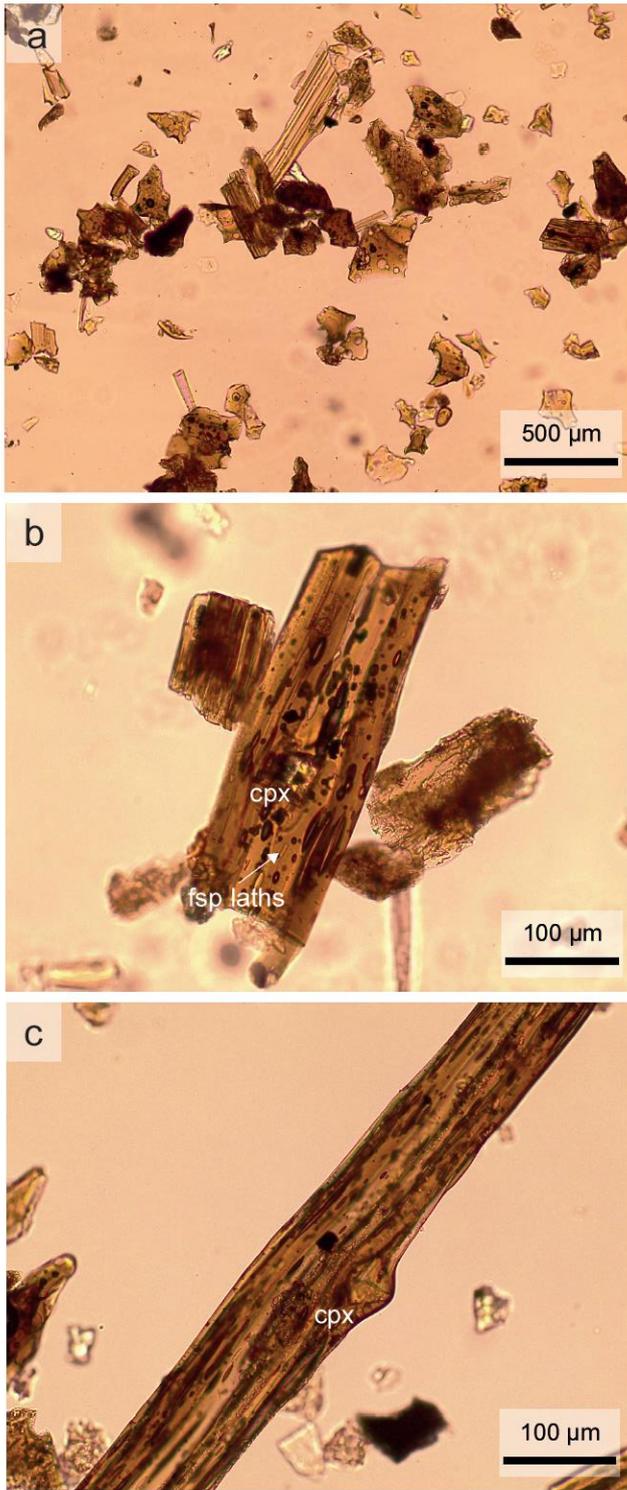
600 **Table 2.** Average major-, minor- (wt.%) and trace-element (ppm) concentrations of the mafic  
601 *DeepCHALLA and CHALLACEA* tephras, with standard deviation added in italics and the  
602 number of shards analysed at the top of each data column. Major and minor elements are  
603 normalised to 100% analytical totals, '<bd"' indicates concentrations below instrument  
604 detection limit. See Supplementary Information for raw data and analyses of reference  
605 materials.

	Challacea				DeepCHALLA													
	CH-3.68		CH-13.22		DCH-57.60		DCH-58.03		DCH-58.41		DCH-69.24		DCH-75.53		DCH-111.72		DCH-228.60	
	<i>n=5</i>		<i>n=35</i>		<i>n=30</i>		<i>n=24</i>		<i>n=29</i>		<i>n=30</i>		<i>n=26</i>		<i>n=16</i>		<i>n=30</i>	
SiO <sub>2</sub>	<b>45.24</b>	1.14	<b>46.48</b>	1.80	<b>45.06</b>	0.55	<b>45.25</b>	1.50	<b>45.26</b>	0.66	<b>44.81</b>	0.58	<b>44.56</b>	0.83	<b>46.07</b>	1.66	<b>44.89</b>	0.69
TiO <sub>2</sub>	<b>4.28</b>	0.16	<b>4.33</b>	0.29	<b>4.46</b>	0.22	<b>4.33</b>	0.62	<b>4.43</b>	0.14	<b>4.41</b>	0.12	<b>4.56</b>	0.22	<b>3.95</b>	0.26	<b>3.37</b>	0.22
Al <sub>2</sub> O <sub>3</sub>	<b>14.30</b>	0.58	<b>14.59</b>	1.13	<b>15.05</b>	0.67	<b>15.23</b>	1.19	<b>14.64</b>	0.35	<b>13.75</b>	0.34	<b>13.74</b>	1.39	<b>15.61</b>	0.71	<b>16.03</b>	0.56
MgO	<b>5.39</b>	0.65	<b>5.17</b>	1.75	<b>5.28</b>	0.76	<b>4.62</b>	0.55	<b>4.93</b>	0.44	<b>5.87</b>	0.49	<b>5.06</b>	1.18	<b>4.10</b>	0.81	<b>4.07</b>	0.48
FeO <sup>T</sup>	<b>13.44</b>	1.00	<b>12.52</b>	1.45	<b>13.35</b>	0.55	<b>12.75</b>	1.37	<b>12.78</b>	0.74	<b>13.54</b>	0.51	<b>13.65</b>	1.24	<b>12.21</b>	0.68	<b>11.19</b>	0.56
MnO	<b>0.21</b>	0.09	<b>0.18</b>	0.11	<b>0.18</b>	0.07	<b>0.20</b>	0.11	<b>0.22</b>	0.10	<b>0.21</b>	0.06	<b>0.23</b>	0.10	<b>0.21</b>	0.05	<b>0.26</b>	0.14
CaO	<b>11.05</b>	0.43	<b>10.74</b>	2.11	<b>10.70</b>	0.69	<b>10.53</b>	1.14	<b>10.96</b>	0.49	<b>11.24</b>	0.23	<b>11.72</b>	1.72	<b>10.10</b>	1.36	<b>9.79</b>	0.96
Na <sub>2</sub> O	<b>3.74</b>	0.43	<b>3.53</b>	1.23	<b>3.76</b>	0.46	<b>4.35</b>	0.64	<b>4.09</b>	0.57	<b>3.71</b>	0.31	<b>4.03</b>	0.85	<b>4.42</b>	0.64	<b>6.15</b>	0.60
K <sub>2</sub> O	<b>1.55</b>	0.14	<b>1.70</b>	0.65	<b>1.48</b>	0.13	<b>1.79</b>	0.27	<b>1.80</b>	0.17	<b>1.68</b>	0.13	<b>1.61</b>	0.26	<b>2.33</b>	0.28	<b>2.91</b>	0.29
P <sub>2</sub> O <sub>5</sub>	<b>0.80</b>	0.12	<b>0.77</b>	0.32	<b>0.67</b>	0.07	<b>0.94</b>	0.21	<b>0.90</b>	0.11	<b>0.78</b>	0.09	<b>0.85</b>	0.15	<b>0.97</b>	0.15	<b>1.33</b>	0.13
					<i>n=20</i>		<i>n=16</i>		<i>n=15</i>		<i>n=20</i>		<i>n=15</i>		<i>n=15</i>		<i>n=18</i>	
Rb					<b>33.1</b>	1.53	<b>42.4</b>	11.4	<b>43.9</b>	3.59	<b>40.4</b>	1.98	<b>43.7</b>	7.66	<b>69.2</b>	11.62	<b>73.1</b>	7.86
Sr					<b>802</b>	38.1	<b>940</b>	102	<b>963</b>	114	<b>878</b>	41.6	<b>1020</b>	138	<b>1212</b>	118	<b>1530</b>	149
Y					<b>30.2</b>	1.08	<b>37.1</b>	3.28	<b>35.2</b>	3.87	<b>32.3</b>	1.82	<b>38.0</b>	4.14	<b>36.4</b>	5.96	<b>36.2</b>	2.94
Zr					<b>331</b>	14.3	<b>439</b>	91.5	<b>452</b>	46.4	<b>366</b>	19.5	<b>468</b>	35.7	<b>430</b>	47.9	<b>480</b>	34.3
Nb					<b>79.6</b>	3.30	<b>106</b>	18.1	<b>108</b>	11.8	<b>93.2</b>	3.19	<b>126</b>	14.2	<b>151</b>	12.8	<b>187</b>	15.6
Cs					<b>0.32</b>	0.048	<b>0.45</b>	0.23	<b>0.44</b>	0.10	<b>0.40</b>	0.042	<b>0.44</b>	0.11	<lld	<lld	<b>0.74</b>	0.083
Ba					<b>462</b>	21.2	<b>614</b>	86.8	<b>647</b>	73.6	<b>542</b>	22.0	<b>598</b>	81.4	<b>1020</b>	95.7	<b>1192</b>	106
La					<b>58.3</b>	2.52	<b>82.1</b>	16.2	<b>84.8</b>	9.40	<b>71.6</b>	3.37	<b>102</b>	12.0	<b>115</b>	11.7	<b>162</b>	12.6
Ce					<b>123</b>	6.45	<b>174</b>	29.9	<b>179</b>	22.4	<b>148</b>	6.53	<b>215</b>	22.7	<b>221</b>	24.5	<b>307</b>	23.1
Pr					<b>14.0</b>	0.627	<b>20.1</b>	3.13	<b>20.6</b>	2.64	<b>16.7</b>	1.04	<b>24.1</b>	2.60	<b>23.3</b>	3.07	<b>32.0</b>	2.42
Nd					<b>60.1</b>	2.86	<b>83.4</b>	11.4	<b>85.1</b>	10.6	<b>69.0</b>	3.02	<b>92.5</b>	11.0	<b>89.6</b>	14.4	<b>117</b>	8.91
Sm					<b>11.9</b>	0.719	<b>15.2</b>	2.23	<b>15.8</b>	2.26	<b>13.1</b>	0.912	<b>16.6</b>	2.22	<b>15.3</b>	2.51	<b>18.5</b>	1.63

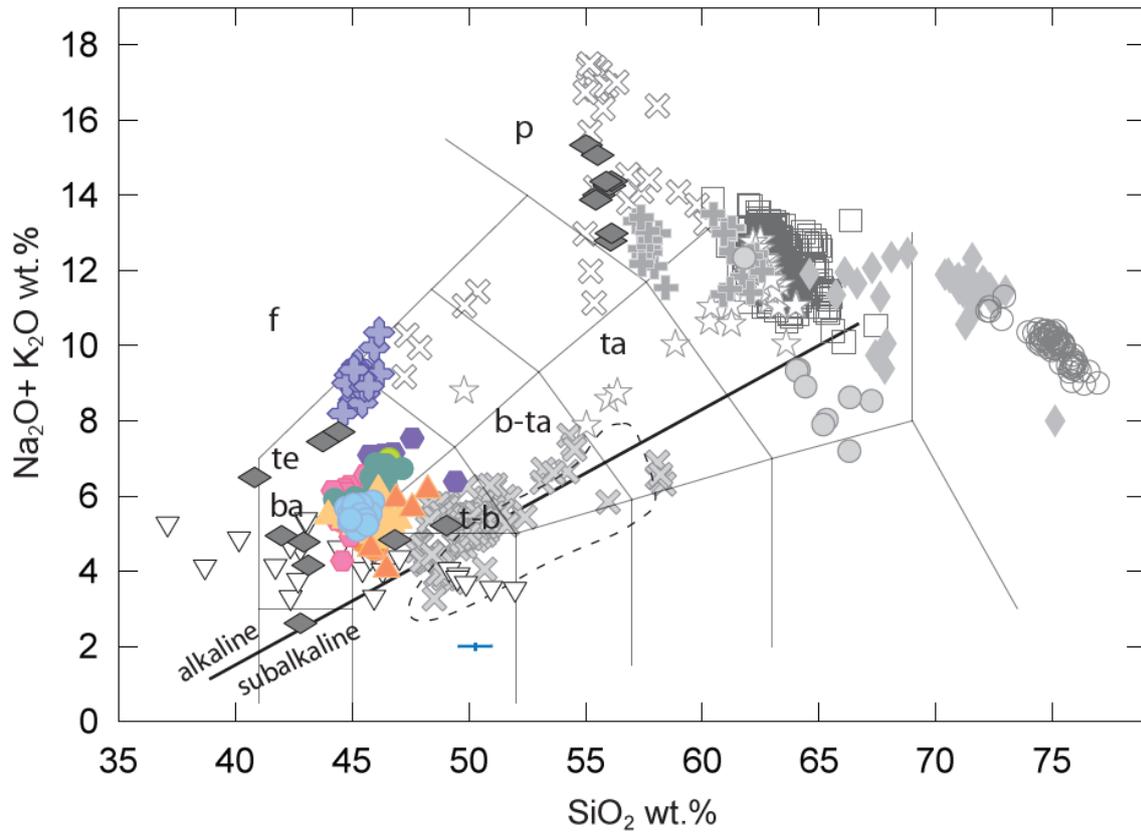
Eu	<b>3.53</b>	0.193	<b>4.59</b>	<i>0.495</i>	<b>4.48</b>	<i>0.561</i>	<b>3.77</b>	<i>0.207</i>	<b>4.73</b>	<i>0.563</i>	<bd	<bd	<b>5.10</b>	<i>0.408</i>
Dy	<b>7.11</b>	0.341	<b>9.52</b>	<i>1.03</i>	<b>8.61</b>	<i>1.42</i>	<b>7.48</b>	<i>0.520</i>	<b>8.94</b>	<i>1.32</i>	<bd	<bd	<b>8.24</b>	<i>0.805</i>
Er	<b>2.94</b>	0.165	<b>3.96</b>	<i>0.806</i>	<b>3.54</b>	<i>0.678</i>	<b>3.17</b>	<i>0.263</i>	<b>3.76</b>	<i>0.520</i>	<bd	<bd	<b>3.54</b>	<i>0.339</i>
Yb	<b>2.26</b>	0.180	<b>2.96</b>	<i>0.451</i>	<b>2.74</b>	<i>0.611</i>	<b>2.35</b>	<i>0.232</i>	<b>2.80</b>	<i>0.577</i>	<bd	<bd	<b>2.82</b>	<i>0.444</i>
Hf	<b>7.55</b>	0.421	<b>10.7</b>	<i>1.60</i>	<b>10.1</b>	<i>1.61</i>	<b>8.32</b>	<i>0.766</i>	<b>10.7</b>	<i>1.50</i>	<b>8.84</b>	<i>1.58</i>	<b>8.95</b>	<i>0.945</i>
Ta	<b>4.72</b>	0.214	<b>6.47</b>	<i>0.764</i>	<b>6.59</b>	<i>0.847</i>	<b>5.54</b>	<i>0.306</i>	<b>7.63</b>	<i>1.02</i>	<b>8.51</b>	<i>0.916</i>	<b>10.2</b>	<i>0.816</i>
Pb	<b>4.17</b>	0.583	<b>6.14</b>	<i>3.41</i>	<b>6.74</b>	<i>3.08</i>	<b>5.03</b>	<i>0.643</i>	<b>6.25</b>	<i>1.54</i>	<b>10.0</b>	<i>3.14</i>	<b>11.7</b>	<i>2.70</i>
Th	<b>6.52</b>	0.346	<b>9.68</b>	<i>2.54</i>	<b>9.62</b>	<i>1.64</i>	<b>8.00</b>	<i>0.642</i>	<b>11.2</b>	<i>1.52</i>	<b>14.7</b>	<i>2.42</i>	<b>19.2</b>	<i>1.70</i>
U	<b>1.65</b>	0.162	<b>2.71</b>	<i>1.22</i>	<b>2.62</b>	<i>0.491</i>	<b>1.98</b>	<i>0.169</i>	<b>2.93</b>	<i>0.424</i>	<b>3.46</b>	<i>0.441</i>	<b>4.83</b>	<i>0.551</i>



**Fig. 1** Study area and geographical context. (a) Volcanic and tectonic features of the East African Rift System in Kenya and northern Tanzania (grey box in the inset map of East Africa shows location), with spatial extent of Cenozoic and Quaternary volcanism (light and dark grey shading) redrawn after Späth et al. (2001), and locations of Holocene volcanoes (triangles) extracted from the Global Volcanism Program (2013) online database. (b) Digital elevation model of the study area in southeastern Kenya and northeastern Tanzania (box in panel a) with location of Lake Chala relative to the Mt. Kilimanjaro and Chyulu Hills volcanic fields (white boxes). Major settlements and locations referred to in the text are indicated with black and white circles, respectively. The inset map shows the bathymetry of Lake Chala (in metres) with indication of the CHALLACEA (white star) and DeepCHALLA (black star) coring sites.



**Fig. 2** Distinctive mafic glass-shard morphologies in Lake Chala sediments. (a) typical mafic shard assemblage. (b-c) individual shards with clinopyroxene phenocrysts and minor feldspar laths.



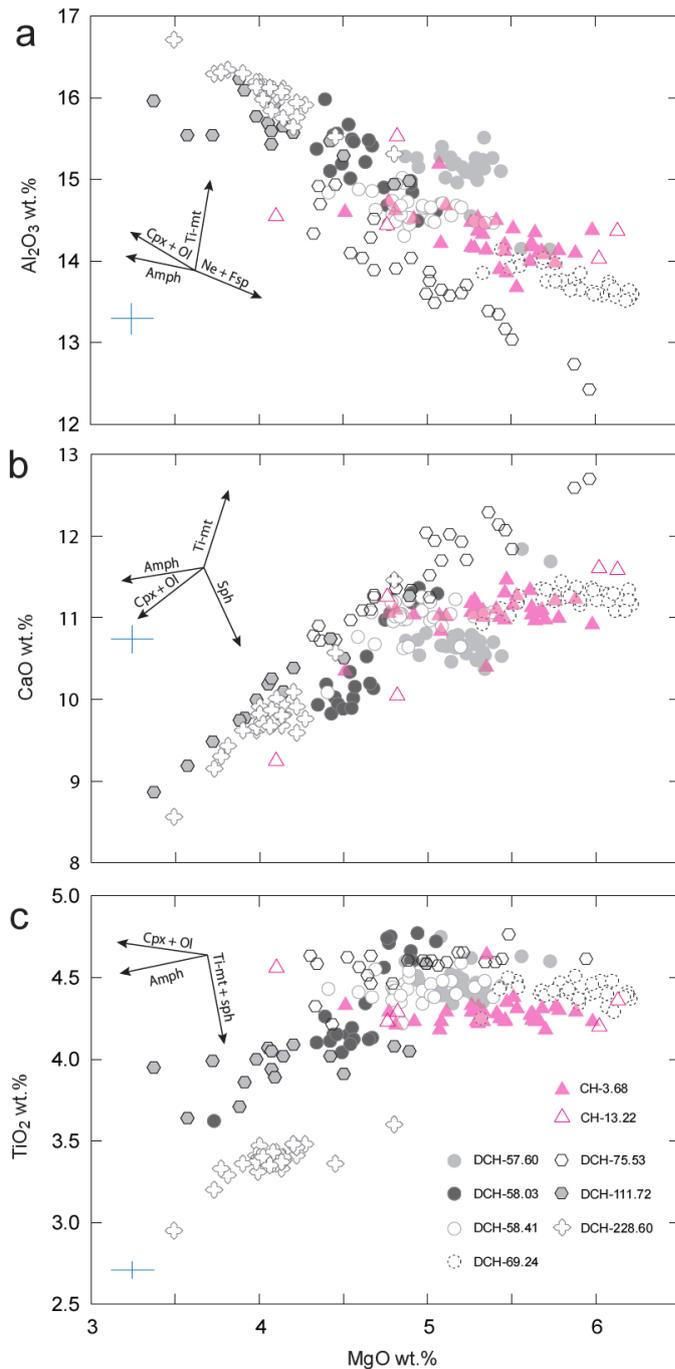
*mafic Chala tephras*

<i>Challacea</i>		<i>DeepCHALLA</i>				
▲ CH-3.68	▲ CH-13.22	● DCH-57.60	● DCH-58.41	● DCH-75.53	◆ DCH-228.60	
		● DCH-58.03	● DCH-69.24	● DCH-111.72		

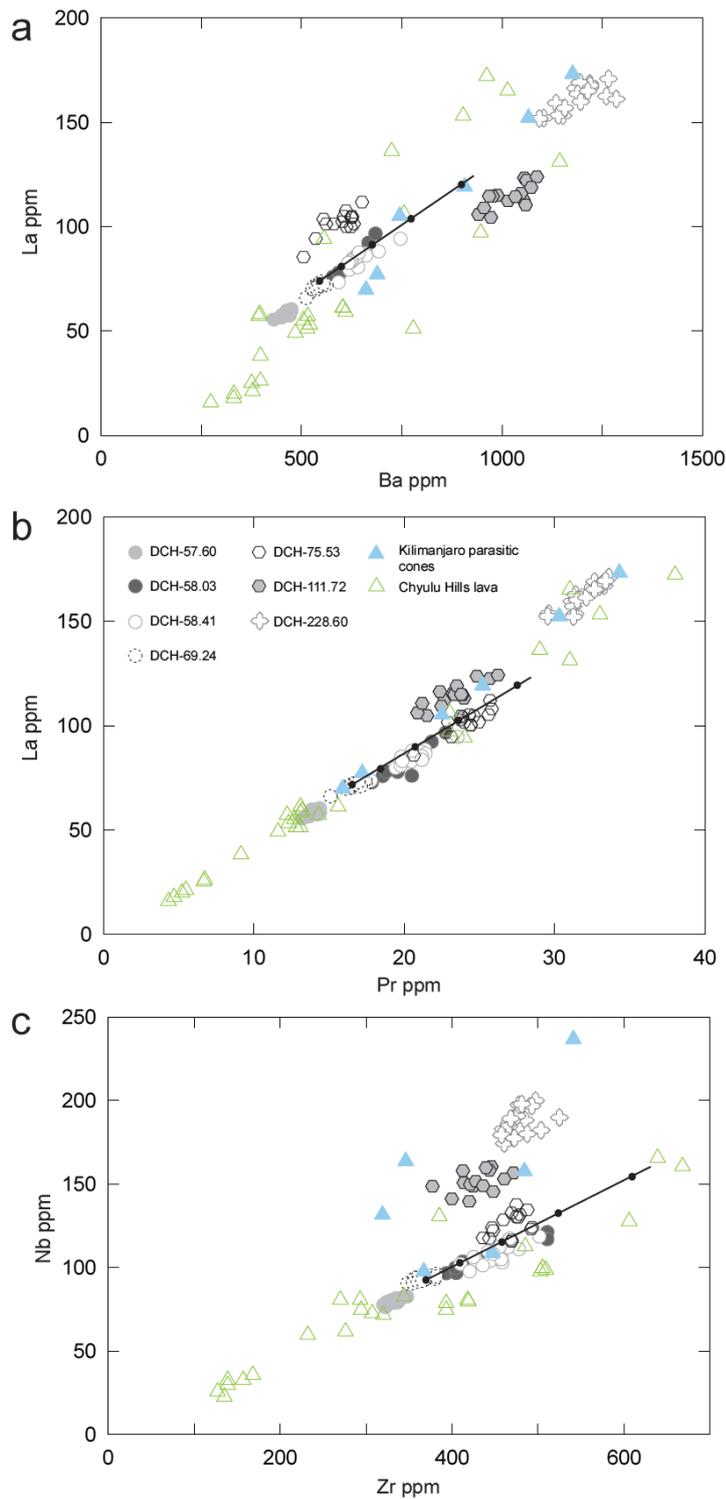
*Central volcanoes and scoria cones*

<i>Kenyan Rift</i>				<i>Off-Rift</i>	
● Silali	□ Menengai	○ Olkaria	⊕ Suswa	▽ Chyulu	⊗ Meru
⊗ Korosi	◆ Eburru	☆ Longonot	⊖ Elementeita, Ndabibi, Akira & Emurangogolak	◆ Kilimanjaro	

**Fig. 3** Total alkali-silica (TAS) plot showing glass composition of the nine mafic Chala tephra analysed in this study. Error bars (blue cross, 2 sigma) are based on repeat analyses of the MPI-DING basalt standard KL2-G, averaged across all analytical runs. Compositional fields referred to in the text are foidite (f), tephrite (te), basanite (ba), trachybasalt (t-b), basaltic trachyandesite (b-ta), trachyandesite (ta) and phonolite (p) (Le Bas et al., 1986). Alkaline and subalkaline fields are according to Irvine and Barager (1971). Published compositions of tephra and lavas from Kenyan and Tanzanian volcanoes (and dated to the Late Pleistocene or Holocene; Fig. 1), are plotted for comparison. Volcanoes with glass data are underlined in the legend, other data are based on whole-rock analyses. Source publications as follows: **Elementeita**, **Ndabibi**, **Akira** and **Emurungoglak** Rift basalts: compositional envelope from White et al. (2012); **Silali**: the ~130-120 ka Kapedo Tuffs, from Tryon et al. (2008); **Korosi**: the Bedded Tuff Member from Tryon and McBrearty (2002; 2006) and Blegen et al. (2018); **Menengai**: the ~36 ka tuff from Blegen et al. (2016); **Eburru**: analyses of lava, pyroclastics (Ren et al., 2006) and obsidians (Brown et al., 2013); **Olkaria**: obsidians from Marshall et al. (2009) and Brown et al. (2013); **Longonot**: pumice (Scott and Skilling, 1999) and lavas (Rogers et al., 2004); **Suswa**: lavas and tuffs (White et al., 2012); **Meru**: pumice (Fontijn and Lane, unpublished data), lavas and tuff (Roberts, 2003); **Kilimanjaro**: Nonnotte et al. (2011); **Chyulu**: Späth et al. (2001).



**Fig. 4** Single-shard major-element concentrations in the DeepCHALLA (grey symbols) and CHALLACEA (triangles) mafic tephras. Error bars (blue crosses, 2-sigma) are based on repeat analyses of the MPI-DING standard KL2-G, averaged across all analytical sessions. The trajectories of Al<sub>2</sub>O<sub>3</sub> (a), CaO (b) and TiO<sub>2</sub> (c) concentration against MgO (all as wt. %) indicate magma evolution dominated by clinopyroxene and olivine fractionation; the vectors in each panel indicate the approximate crystallisation paths.



**Fig. 5** Relationships between concentrations of the incompatible elements La and Ba (a), La and Pr (b) and Nb and Zr (c) in single shards of the seven DeepCHALLA mafic tephras analysed (grey symbols), compared with those in lavas from Mt. Kilimanjaro 200-150 ka parasitic cones (blue closed triangle; whole-rock analyses by Nonnotte et al., 2011) and in lavas from the Chyulu Hills volcanic field (green open triangles; whole-rock analyses by Späth et al., 2001). Average 2-sigma errors, from repeat analyses of

*the MPI-DING standard StHs6-80G over two analytical sessions (10.5 ppm for Zr, 6.94 ppm for Nb, 36.6 ppm for Ba, 1.16 ppm for La and 0.357 ppm for Pr) are smaller than symbol sizes in this plot. Vectors (with 10% steps) show modelled simple fractional crystallisation of olivine and clinopyroxene using IgPet software (Carr and Gazel, 2017), with DCH-69.24 (the most mafic tephra) as the starting composition and using mineral-melt partition coefficients compiled by Späth et al. (2001).*