

1 **Preservation of systematic Ni and Cr heterogeneity in otherwise**
2 **homogeneous mantle olivine: implications for timescales of post-**
3 **metasomatism re-equilibration.**

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8 **Highlights**

- 9 1. Peridotite xenolith from Bultfontein show evidence of proto-kimberlite metasomatism
10 2. Olivines record unique Ni and Cr disequilibrium, decoupled from Fo, Mn, Sc, Co, Zn
11 3. Incomplete subsolidus re-equilibration observed in Ni and Cr following metasomatism
12 4. Ni-Cr disequilibrium preserved in slow diffusion axes in olivine porphyroclasts
13 5. Period of metasomatism occurred <500,000 years prior to kimberlite eruption

14 **Abstract**

15 The flux of elements into Earth's sub-continental lithospheric mantle is facilitated by the
16 passage of small-fraction melts that either crystallise new phases or react with pre-existing
17 minerals. Metasomatised peridotite records the end product of this exchange but rarely
18 captures the process in the act due to subsolidus re-equilibration. We present the results of
19 a systematic investigation of a metasomatic melt channel preserved in a mantle peridotite
20 from the Late Cretaceous Bultfontein kimberlite (Kaalpvaal craton) that shows rare direct
21 evidence of the melt-rock reaction processes. We show that the metasomatic proto-
22 kimberlite melt underwent variable crystallisation of clinopyroxene, sulfides, phlogopite,
23 spinel and zircon together with interaction and diffusive exchange with the surrounding
24 olivine-rich mantle.

25 Element profiles across large olivine porphyroclasts (Fo₈₈) show significant core-to-
26 rim variations in Ni (NiO = 0.18-0.32 wt.%) and Cr (Cr = 35-60 ppm), while concentrations of
27 all other elements (e.g. Mg, Fe, Mn, Co, V) are remarkably homogeneous. Electron

28 backscatter diffraction analysis shows that the disequilibrium of Ni and Cr is greatest where
29 the crystal contains large components of the [100] and [010] axes. The disequilibrium is
30 preserved in certain orientations because diffusion of Ni and Cr in olivine is more anisotropic
31 than Fe-Mg and Mn, and slower in the [100] and [010] directions. We present the first
32 observations of Ni and Cr decoupling from other elements in mantle olivine and suggest that
33 this is a consequence of: (i) changing mineral-melt concentration gradients associated with
34 the reactive percolation of a precursory kimberlite melt; and (ii) late-stage sulfide and spinel
35 precipitation.

36 We use the diffusion limited re-equilibration of Ni in olivine to quantify the timing of
37 metasomatism prior to xenolith entrainment by the host kimberlite. Our modelling indicates
38 that reactive percolation occurred on the order of 10^3 - 10^5 years prior to entrainment; this
39 provides an additional line of support for the hypothesis that a period of metasomatism by
40 proto-kimberlite melts precedes the final kimberlite ascent to the surface. The broader
41 implication of our finding of variable rates of minor element diffusion in natural olivine is
42 that it highlights the importance of anisotropy and the impact of changing local
43 concentration gradients during subsolidus re-equilibration.

44

45 **Keywords**

46 Diffusion in olivine, metasomatism, crystallographic anisotropy, kimberlite, mantle xenolith

47

48 **1. Introduction**

49 The sub-continental lithospheric mantle represents one of Earth's largest and most long-
50 lived chemical reservoirs. Direct evidence from mantle xenoliths shows that beneath the
51 ancient cores of continents a once refractory melt residue has undergone billions of years of
52 refertilization (Gibson et al., 2008; Menzies and Hawkesworth, 1986; Pearson, 1995; Shu
53 and Brey, 2015; Simon et al., 2007; Woodhead et al., 2017). This long-term metasomatic
54 enrichment results from reactive percolation of low-viscosity melts, such as kimberlites and
55 carbonatites (e.g. Achterbergh et al., 2001; Dawson, 1981; J.B. Dawson and Smith, 1977;
56 Giuliani et al., 2014b; McKenzie, 1989; Rehfeldt et al., 2007). These small-fraction melts are
57 highly charged with volatiles, including CO₂ and H₂O, and cause fracturing of the sub-

58 cratonic mantle that may facilitate channelized flow of subsequent kimberlite melts (e.g.
59 Aulbach et al., 2017; Giuliani et al., 2016, 2014).

60 Mantle metasomatism via reactive percolation of enriched melts may result in: (i)
61 changes in modal mineralogy (modal/patent metasomatism; Harte, 1983); (ii) chemical
62 changes with no accompanying mineralogical change (cryptic metasomatism; Dawson,
63 1984); or (iii) refertilisation of mantle peridotite (stealth metasomatism; O'Reilly and Griffin,
64 2013). Numerous studies have attempted to date metasomatic events using radiometric
65 dating techniques on metasomatic phases such as zircon and titanite. The highest precision
66 are U-Pb dates on mantle zircons which indicate that they precipitate from metasomatic
67 melts within several million years of kimberlite emplacement at the surface. Some of these
68 U-Pb zircon ages are coeval with the host kimberlite (Kinny and Dawson, 1992; Konzett et
69 al., 2013, 2000, 1998) while others are much older (Giuliani et al., 2015, 2014b; Liati et al.,
70 2004; Woodhead et al., 2017). The episodic versus continuous nature, timescales and extent
71 over which metasomatic interactions occur in the sub-continental mantle are, however,
72 poorly constrained. The limitation of radiometric dating techniques is their resolution to an
73 order of millions of years which cannot tell us whether a metasomatic event was related
74 directly to the host kimberlite or not.

75 Higher resolution inferences about the timing of metasomatic events in the mantle are
76 reliant on mineral disequilibrium and diffusion timescales. In the deep lithosphere the
77 timescales of re-equilibration for minerals present in the refractory wall rocks of melt
78 channels (e.g. garnet, olivine and orthopyroxene) are fast ($< 10^5$ years, e.g. (Griffin et al.,
79 1996) so that each crystal only records the most recent chemical perturbation. At shallower
80 levels, the re-equilibration timescales are longer, so the preservation of elemental zoning is
81 more likely. Samples of mantle material entrained during this short re-equilibration time
82 therefore potentially preserve diffusion profiles in minerals that can be used to estimate the
83 timing of metasomatism prior to kimberlite emplacement.

84 Here we present high-resolution *in-situ* micro-analyses of major and trace elements in
85 rare, un-equilibrated, olivines found in a metasomatised peridotite from Bultfontein Mine,
86 South Africa. By combining olivine diffusion profiles for a range of major (Mg, Fe), minor and
87 trace elements (Ni, Mn, Cr, Co, Cu, Ti, V and Zn) with crystallographic controls we show how

88 Ni and Cr disequilibrium can be preserved in otherwise homogeneous mantle olivines. We
89 use diffusion modelling of Ni in olivine to quantify the timescales of equilibration of
90 refractory wall rock following reactive percolation of metasomatic melts through sub-
91 cratonic lithospheric mantle.

92

93 **2. Description of veined mantle peridotite BD3067, Bultfontein**

94 Our study focusses on a large fragment of veined mantle peridotite (BD3067) entrained by
95 the Bultfontein Group I kimberlite in Kimberley (S. Africa). The emplacement of this pipe at
96 $84 (\pm 0.9)$ Ma (Kramers et al., 1983) was coeval with widespread Group I kimberlite activity
97 across the southern Kaapvaal craton (Griffin et al., 2014). Numerous previous studies have
98 shown that the Bultfontein kimberlite sampled highly-heterogeneous and extensively-
99 metasomatised mantle (e.g. Erlank, 1987; Giuliani et al., 2013a, 2013b; Jones et al., 1982;
100 Kramers et al., 1983; Rehfeldt et al., 2007; Simon et al., 2007). Mineral separates from a
101 suite of Bultfontein xenoliths (including BD3067) were initially analysed for REEs together
102 with Sr-Rb, Sm-Nd and Pb isotopes by Kramers et al. (1983). They concluded that a
103 kimberlite melt, which preceded emplacement of the Bultfontein kimberlite, was
104 responsible for the metasomatism. Our study builds on this work by carrying out systematic
105 *in-situ* analyses of major, trace and minor elements in minerals from one of these samples,
106 BD3067. We combine these results with a petrographic and microstructural study in order
107 to quantify the timescales of mineral equilibration and length scales of metasomatism.

108 Mantle xenolith BD3067 contains a spectacular, bright green, trichotomous branching,
109 clinopyroxene-rich vein set in a matrix of olivine porphyroclasts and neoblasts (Figures 1a
110 and b). Orthopyroxene is absent, and the vein host rock is a dunite. At its widest point, the
111 clinopyroxene-rich vein measures 45 mm. Branch terminations and regions of more isolated
112 clinopyroxene crystallisation are associated with small amounts of phlogopite. The large
113 olivine porphyroclasts and clinopyroxenes (both 2-6 mm diameter) show significant internal
114 deformation, exhibited in olivine subgrains, recrystallisation and complex fine structures
115 (Figures 1c and d). Accessory phases include large anhedral zircons (up to 5mm diameter;
116 Figure 1e) and very small amounts of Cr-spinel that have crystallised within the main vein.

117 Away from this region we observe phlogopite, large (0.2-4 mm) Fe-Ni sulfides, interstitial
118 amongst the small (<50 μm) olivine neoblasts, and ilmenite.

119 Multiple thin sections were made from xenolith BD3067, including across the main
120 vein and regions further from it (Figure 1a).

121

122 **3. Mineral chemistry of veined mantle peridotite BD3067**

123 Olivine, clinopyroxene, phlogopite and sulfides were analysed for major and minor
124 elements using a Cameca SX100 electron probe microanalyser (EPMA). Minor and trace
125 element concentrations were determined on the same grains with an ESI UP193UC laser
126 inductively coupled to a Nexion 350D quadruple mass spectrometer (LA-ICP-MS). All of
127 these analyses were undertaken in the Department of Earth Sciences at the University of
128 Cambridge. Further details of these analytical setups are provided in Supplementary File A.
129 Representative clinopyroxene and olivine compositions are provided in Table 1 and 2,
130 respectively, sulfide compositions are given in Table 3 and all analyses in Supplementary File
131 B.

132

133 **3.1 Clinopyroxene**

134 The clinopyroxenes in BD3067 have a high Mg number ($\text{Mg\#} = 89\text{-}93$, average $\text{Mg\#} 91.1 \pm$
135 1.8 (2σ), $n=45$) and CaO content (22 ± 0.7 wt. %, $n=45$) and are diopsides (Table 1; Figure 2).
136 They have low TiO_2 (0.20 ± 0.03 wt.%, $n=45$) and Na_2O (1.23 ± 0.37 wt.%, $n=45$), and very
137 low Al_2O_3 contents (0.63 ± 0.15 wt.%, $n=45$). Diopsides in the main vein have lower Cr_2O_3
138 (0.74 ± 0.14 wt. %, $n=15$) and higher FeO (3.08 ± 0.25 wt.%, $\text{Mg\#} = 90.6 \pm 0.75$, $n=15$) than
139 those further away ($\text{Cr}_2\text{O}_3 = 0.98 \pm 0.86$ wt. %; $\text{FeO} = 2.89 \pm 0.58$ wt.%; $\text{Mg\#} = 91.1 \pm 1.78$,
140 $n=10$). This relationship is illustrated in Figure 3a. There is also a subtle difference in Ni
141 content with the distal diopsides having a higher Ni content and greater variability ($243 \pm$
142 110 ppm, $n=10$) than those found in the vein (185 ± 23 ppm, $n=15$).

143 The diopsides have remarkably high Zr (107 ± 38 ppm) and Hf (6.63 ± 2.85 ppm)
144 concentrations so that primitive mantle normalised values are 10 and 24, respectively
145 (Figure 3). They are enriched in light rare earth elements (REE) with concentrations 3 to 8

146 times primitive mantle. The heavy REE contents of the diopsides are by comparison low and
147 4-5 times lower than primitive mantle. As a result, the diopsides have high La/Yb (12-17) and
148 Zr/Hf (90-130), but very low Ti/Eu (1300-1600) compared to primitive mantle (McDonough
149 and Sun, 1995). While the concentration of incompatible trace elements in diopsides in the
150 main vein is more uniform, their composition lies within the wider range of pyroxenes away
151 from the main vein.

152 **3.2 Olivine**

153 The olivine neoblasts (nb) and porphyroclasts (pc) in BD3067 are characterised by uniform
154 but moderate forsterite contents (nb Mg# = 88.0 ± 0.36 , n=14; pc Mg# = 88.0 ± 0.22 , n=12,
155 Table 2; Figure 4). Both generations also have moderate MnO (nb = 0.15 ± 0.04 wt.%, n=14;
156 pc = 0.16 ± 0.03 wt.%, n=12) but low CaO ($<<0.1$ wt. %) concentrations. Similar to the
157 diopsides, all of the olivines have extremely low contents of Al (8 ± 5 ppm, n=12), and the
158 neoblasts have low NiO (0.21 ± 0.05 wt.%, n=14).

159 The olivines in BD3067 have higher P (99 ± 35 ppm, n=9), Ti (150 ± 23 ppm, n=9) and
160 Zn contents (105 ± 19 ppm, n=18) but similar concentrations of V (3.9 ± 0.7 ppm), Co ($133 \pm$
161 11.4 ppm) and Cu (1.3 ± 0.6 ppm) to those found in previous studies of mantle olivines (e.g.
162 (Aulbach et al., 2017; De Hoog et al., 2010). All of the olivine neoblasts are uniform in
163 composition but a number of the large porphyroclasts have Ni- and Cr-rich cores and Ni- and
164 Cr-poor rims. These rims have the same composition as the neoblasts and the un-zoned
165 porphyroclasts (NiO = 0.20 ± 0.02 wt.%, n=13; Cr = 37 ± 6 ppm, n=9). The cores of the zoned
166 olivines show a range in Ni (0.20 – 0.33 wt.% NiO), Cr (32 – 63 ppm) and Ti (117 – 204 ppm)
167 contents but are uniform for every other element and the same composition as the
168 neoblasts and un-zoned olivine. This is demonstrated by comparing the NiO and MnO
169 contents in Figures 4a and b.

170 We measured profiles of major and minor element concentrations across sixteen,
171 large (2 to 6.3 mm), olivine porphyroclasts in BD3067. Six of these profiles are across
172 olivines that border the main clinopyroxene vein (hereby referred to as proximal olivine).
173 The remaining ten profiles are across porphyroclasts distal to the main vein (hereby referred
174 to as distal olivine), where the proportion of clinopyroxene is lower and phlogopite and
175 sulfides higher. Examples of representative profiles of elements across olivine

176 porphyroclasts, determined by EPMA and LA-ICP-MS, in both the proximal and distal
177 sections are shown in Figure 5, and all profiles and analyses are presented in Supplementary
178 File B.

179 An important finding arising from our study is that olivine porphyroclasts which
180 border the main vein in BD3067 have homogeneous major and trace element compositions
181 (e.g. Figure 5d, Olivine M) whereas those away from this region (e.g. Olivine A, D) exhibit a
182 range in Ni, Cr and Ti but have constant Si, Mg, Fe and Mn, Co, V and Zn contents. Figures 4d
183 and 4f show a positive correlation between Ni and Cr, and Ni and Ti, respectively. This
184 reflects the zonation in these elements, however, the variation is not uniform across the
185 olivine porphyroclasts: some show much stronger core-to-rim variation than others; some
186 have parabolic zonation (e.g. Olivine A) and others more complex patterns. While the olivine
187 crystals that exhibit zonation in Ni and Cr tend to be the largest, there are others of the
188 same size that have homogeneous compositions. The gradient of the Ni and Cr profiles
189 varies within and between crystals. Figure 5 shows profiles for two endmembers: Olivine A
190 has the most strongly zoned core-to-rim profile and Olivine B has uniform concentrations of
191 NiO (0.2 wt.%) and Cr (34 ppm) that are the same as those in the neoblasts. The core
192 composition in the un-equilibrated profiles range from 0.28 - 0.32 wt.% NiO, and 50 – 63
193 ppm Cr. By contrast the rim composition of the zoned olivines are strikingly uniform with
194 0.18-0.2 wt.% NiO and 35-40 ppm Cr. The rims have the same chemical composition for all
195 elements as the olivine neoblasts.

196

197 **3.3 Sulfides**

198 We observed several large, irregular, metasomatic Fe-Ni sulfides in the thin sections of
199 BD3067 (Figure 6). The shape, size and composition of these grains are extremely
200 heterogeneous. Most of the sulfides have been heavily serpentinised to the low
201 temperature assemblage of magnetite, heazlewoodite and serpentine (Lorand and Grégoire,
202 2006), but some unaltered areas remain. The composition of the unaltered regions of the
203 sulfides are shown in Table 3. The heterogeneity in composition is especially prevalent in Cu
204 and Ni. Some regions have higher Cu content; however a low Cu concentration is common
205 in Kaapvaal peridotite mono-sulfide-solution and demonstrates that they have re-

206 equilibrated at a low temperature (Lorand and Grégoire, 2006). There also appear to be
207 exsolution lamellae in the sulfides that are often Ni-rich (up to 44 wt. % NiS).

208

209 **3.4 Zircon**

210 Energy Dispersive X-ray Spectroscopy (EDS) major element mapping of the largest
211 zircon, which is present in a bottlenecked region of the main vein and adjacent to olivine
212 porphyroclasts, shows that it is unzoned but regions of baddeleyite (ZrO_2) are present
213 adjacent to cracks in the crystal.

214

215 **4. Crystal orientation of olivine**

216 Previous studies of element diffusivity in olivine have highlighted the importance of
217 crystal orientation (e.g. Costa and Morgan, 2010; Dohmen and Chakraborty, 2007; Spandler
218 and O'Neill, 2010). We used a FEI Quanta 650FEG SEM equipped with a Bruker e- Flash HR
219 Electron Back-Scatter Diffraction (EBSD) detector in the Department of Earth Sciences at the
220 University of Cambridge to make EBSD maps of variably equilibrated olivine porphyroclasts
221 in veined xenolith BD3067, in regions parallel to the electron microprobe and LA-ICP-MS
222 profiles. We used the EBSD data to calculate the orientation of crystals in the thin sections,
223 i.e. angles of the plane to the $\langle 100 \rangle$ axes. These reveal that the angle between the crystal
224 face and the position of the [001] axis correlates with the level of heterogeneity in Ni and Cr
225 in olivines distal to the vein (Figure 7). This is significant because diffusion of major and
226 minor elements in olivine is anisotropic and faster along the [001] axis than the [100] and
227 [010] axes (Dohmen & Chakraborty, 2007). The proximal olivines are all homogeneous,
228 regardless of the crystallographic orientation.

229 Figure 7 shows the correlation between the range in NiO across the crystal and the
230 angle between the profile and the [001] fast diffusion axis. The profiles taken in crystals
231 normal to the [001] axis show the greatest degree of heterogeneity, for example Olivines A
232 and D (Figure 5a and c), whereas the most equilibrated profile in the distal sections (across
233 Olivine B) is parallel to the [001] axis. There are a number of intermediate levels of
234 equilibration that also reflect the angle of the cut plane.

235

236 5. Calculated and inferred melt chemistry

237 Olivines found in lithospheric mantle peridotites typically have NiO concentrations of
238 0.34 – 0.43 wt.% that remain roughly constant over a range of Fo contents (Foley et al.,
239 2013 and references therein; Bussweiler et al., 2015). The peak NiO concentration in the
240 parabolic profiles (0.36 wt. %) across olivine porphyroclasts in the veined Bultfontein
241 xenolith BD3067 are similar to non-metasomatised mantle olivines (De Hoog et al., 2010;
242 Foley et al., 2013). Figure 8 compares the composition of the olivine neoblasts and
243 porphyroclasts proximal and distal to the clinopyroxene vein with the composition of olivine
244 cores and rims in kimberlitic olivine from the Kimberley region, as compiled by (Giuliani,
245 2018). The rims of the kimberlitic olivines represent magmatic olivine that has crystallised
246 from the magma, and the cores represent xenocrystic mantle olivine (Giuliani, 2018). The
247 BD3067 proximal olivine and the distal olivine neoblasts and porphyroclast rims have a NiO
248 and Fo content that resembles the magmatic kimberlite rather than un-metasomatised
249 mantle olivine. The olivine porphyroclast cores have higher NiO, similar to mantle olivine.
250 The MnO content of all of the olivine crystals is very similar to the kimberlite magmatic
251 olivine.

252 We have used experimentally-determined partition coefficients to calculate the
253 incompatible trace element concentrations of melts in equilibrium with the diopsides in
254 BD3067. Figure 9 compares these to the composition of the Bultfontein kimberlite (Roex et
255 al., 2003) and a global carbonatite average (Bizimis et al., 2003). The four different sets of
256 partition coefficients used reflect four different melt compositions at mantle pressure and
257 temperature: (i) a silicate melt (Suzuki et al., 2012); (ii) carbonatite melt (Dasgupta et al.,
258 2009); (iii) kimberlite melt (Keshav et al., 2005); and (iv) a silico-carbonate melt (Girnis et al.,
259 2013). The main differences between the sets of partition coefficients are revealed in the
260 behaviour of the HFSE, especially Nb, Ta, Zr and Hf. The partition coefficients using
261 kimberlite and silico-carbonate melts calculate the melt to be much more enriched in Zr and
262 Hf than the silicate and carbonatite melts. The implication that the melt was enriched in Zr is
263 supported by the abundance of large metasomatic zircons in the vein (e.g. Figure 1e). Of the
264 cpx equilibrium melts calculated there is a good fit between the carbonatite cpx-melt and

265 the Bultfontein kimberlite which matches the Zr/Hf ratio but differs in the Nb/Ta ratio. The
266 silico-carbonate cpx-melt predicts the same Nb/Ta fractionation but no fractionation in
267 Zr/Hf. Apart from the enrichment in Zr and Hf the silico-carbonate melt coefficients produce
268 a good fit with the Bultfontein kimberlite. It is possible the composition of the Bultfontein
269 kimberlite shown is depleted in Zr and Hf due to zircon fractionation.

270 The diopsides have very high Ca/Al, low Ti/Eu and high Zr/Hf ratios. These are often
271 identified as features of carbonatite metasomatism (Rudnick et al., 1993), however the lack
272 of magmatic carbonate and the abundance of a metasomatic silicate phase (diopside)
273 implies that the metasomatic agent was not carbonatitic in nature. Instead we suggest that
274 the melt was a silico-carbonate melt, consistent with the results of the recent study by
275 Soltys et al. (2018) who reconstructed the parental composition of the Bultfontein
276 kimberlite to be a transitional silicate-carbonate melt, and those of (Simon et al., 2003)
277 whose suggest clinopyroxene in mantle peridotite from the Kaapvaal is the product of
278 metasomatism by kimberlite-like magmatism.

279

280 **6. Mineral disequilibrium**

281 Chemical disequilibrium is a common observation in mantle peridotites that have
282 experienced metasomatism, but the decoupling of Ni and Cr disequilibrium from any other
283 element in mantle olivine that we have observed is a unique and puzzling finding. The
284 olivine profiles reported here do, however, show a broad similarity to profiles observed in
285 olivines from the Springwater pallasite (Leitch et al., 1979; Zhou and Steele, 1993). In these
286 meteoritic olivines, long wavelength diffusion profiles (mm scale) are observed in several
287 elements, including Ni, Cr, Ca and Al, but flat profiles are observed in Fo content, Mn and V.
288 In this case the outward diffusion of Ni and Cr is attributed to the decrease in the
289 equilibrium concentration of the olivine with the adjacent metal phase, during cooling. The
290 flat profiles are recognised as resulting from the lack of any other sink for these elements
291 (e.g. V and Mn) outside of the olivine.

292 To our knowledge there has been no other study presenting decoupling of Ni and Cr
293 from other elements in olivine. In BD3067 Ni and Cr are the only elements to show
294 systematic core-to-rim variation although a few of the other minor elements show a degree

295 of scatter. In some cases, for example Ti, this scatter weakly correlates with the Ni and Cr
296 zoning (Figure 4f). The core to rim multi-element decoupling, on millimetre length scales,
297 indicates that these variations are not caused by overgrowth of olivine formed by
298 fractionation of a crystallising melt, as is observed in olivine xenocrysts in kimberlite
299 magmas (e.g. Arndt et al., 2010; Bussweiler et al., 2015; Giuliani et al., 2017; Pilbeam et al.,
300 2013). It is unlikely that this is a growth effect due to the shape of the profile and the
301 distance of the zoned olivines from the main vein, where the infiltrating melt would have
302 facilitated growth. Instead, this is an effect of subsolidus re-equilibration.

303 No single factor can satisfactorily account for the observed decoupling, rather
304 several aspects have produced, and influenced the preservation of, the elemental patterns
305 we see. These include: (i) mineral-melt and mineral-mineral concentration gradients during
306 subsolidus re-equilibration; (ii) the presence of Ni/Cr stoichiometric phases; and (iii)
307 anisotropy of diffusion in olivine. We suggest that the observed profiles are caused by an
308 evolving melt composition and locally changing concentration gradients as new phases
309 crystallise and change the local equilibrium conditions. The profiles represent the re-
310 distribution of elements during subsolidus re-equilibration (Cherniak and Liang, 2014,
311 2012), but the fact that we only see the profiles in Ni and Cr, and to a lesser extent Ti, is
312 influenced by the local mineral assemblage and the diffusion of these elements in olivine; Ni
313 and Cr because they exhibit strong anisotropic diffusion in olivine, and Ti because it's high
314 charge makes diffusion very slow in olivine.

315 **6.1 Concentration gradients: Subsolidus re-equilibration**

316 Re-equilibration during and after metasomatism is influenced by compositional
317 differences and concentration gradients between: (i) the infiltrating melt; (ii) the pre-
318 existing mantle phases; and (iii) the new metasomatic phases. We have established that the
319 infiltrating melt was a silico-carbonate melt and suggest that this resorbed orthopyroxene
320 and crystallised clinopyroxene (Lim et al., 2018; Simon et al., 2003). This early pulse of
321 volatile-rich proto-kimberlite melt introduced a suite of elements into the system, altering
322 the equilibrium conditions of the local mantle peridotite. Following the melt infiltration
323 event, the new mineral assemblage required elemental re-distribution to re-equilibrate.

324 In the main vein of BD3067 the olivine and clinopyroxene are homogeneous.
 325 Equilibrated coexisting mantle olivine and clinopyroxene have the same Mg# ($K_D = 1$)
 326 (Pearson et al., 2003; Seckendorff and O'Neill, 1993). None of the olivine and clinopyroxene
 327 pairs in BD3067 are in equilibrium. In the main vein of BD3067, the homogeneous
 328 clinopyroxene and bordering olivine have Mg# of 90.6 ± 0.75 ($n=15$) and 88.0 ± 0.21 ($n=12$),
 329 respectively. The distal olivine porphyroclasts and the clinopyroxene away from the main
 330 vein are not equilibrated (cpx Mg# = 91.1 ± 1.78 , $n=30$; olivine Mg# = 88.0 ± 0.36 , $n=14$),
 331 and the clinopyroxenes show an even greater range in Mg#. As for Ni, based on a
 332 comparison with D_{Ni}^{ol-cpx} of close to 0.13 for equilibrated mantle xenoliths (Gibson et al.,
 333 2013; Stosch, 1981), we observe that the lower Ni, proximal clinopyroxenes (Ni = 185 ± 23
 334 ppm) are close to equilibrium with: (i) unzoned olivines; (ii) neoblasts; and (iii) porphyroclast
 335 rims (Ni = 1548 ± 271 ppm) .

336 The clear differences between the levels of equilibration in mineral grains in the main
 337 vein and the regions distal to it in BD3067 highlight controls of: (i) local temperature; (ii)
 338 concentration gradients; and (iii) mineral assemblage. The olivine porphyroclasts bordering
 339 the main vein all have homogeneous compositions (e.g. Olivine M in Figure 4e). It is likely
 340 that the diffusivity of Ni in olivine and subsequent re-equilibration was increased near the
 341 vein due to the higher concentration gradients and temperatures associated with the larger
 342 volume of melt undergoing channelized flow (Griffin et al., 1996).

343 **6.2 Late-stage crystallisation of accessory phases**

344 Our data suggest that the olivine content became buffered at Fo₈₈, perhaps due to
 345 the resorption of orthopyroxene (Bussweiler et al., 2015; Giuliani et al., 2017; Pilbeam et al.,
 346 2013), but the Ni and Cr contents were still changing in response to the precipitation of late
 347 stage minerals such as sulfides and chrome-spinel and a lag in diffusion. We suggest that the
 348 infiltrating silico-carbonate melt was depleted in Ni (Roex et al., 2003; Soltys et al., 2018)
 349 and the Cr concentration will depend on the abundance of garnet in the melt source. The
 350 clinopyroxenes are not particularly enriched in Cr and Cr-spinel is only a minor phase in the
 351 vein assemblage, which implies that the melt was not significantly enriched in Cr.

352 A number of authors have suggested that an immiscible sulfide melt is produced at a
 353 late stage of kimberlite melt evolution (Giuliani et al., 2013a; Lorand and Grégoire, 2006)

354 and (Aulbach et al., 2017) attribute low Ni in mantle clinopyroxene from SW Greenland to
355 concomitant sulfide saturation during metasomatism. For olivine to exchange Ni with
356 sulfides, the latter must have initially been low in Ni. If the metasomatic agent was Ni-poor
357 then it is possible that a Fe-Cu immiscible liquid separated from the melt and precipitated
358 Fe-Cu base metal sulfides, as suggested by (Lorand and Grégoire, 2006) for the origin of Fe-
359 Cu-Base Metal Sulfides in phlogopite (\pm ilmenite \pm rutile) peridotites. In the case of BD3067,
360 the introduction of sulfides, a phase for which Ni is a stoichiometric component, shifted the
361 distribution coefficients for the evolving mineral assemblage. As a result local chemical
362 gradients changed depending on the resorbing or crystallising phases (e.g. Giuliani et al.,
363 2014); the late-stage crystallisation of Ni-poor sulfides catalysed Ni re-distribution and drove
364 the Ni concentration down in olivine (e.g. Barnes et al., 2013). The sulfides are associated
365 with olivine porphyroclasts, but there is no systematic spatial relationship between the
366 position of the sulfides and the specific zoned olivine porphyroclasts. The lack of a close
367 spatial relationship between the olivine and the sulfides suggests that there has been
368 migration of Ni to distant sulfides facilitated by open grain boundaries and cracks in the old
369 crystals; this sluggish solid-state diffusion is able to produce long wavelength, near-
370 symmetrical profiles like those observed in this study (e.g. Leitch et al., 1979). The same
371 systematic variation in Fe and Cu is not observed (other stoichiometric components of
372 sulfides) and we suggest this is because the melt was enriched in Fe and Cu relative to the
373 olivine. It is also possible that we do not see the same effect on Fe in olivine because Fe is a
374 stoichiometric component of olivine and the effect would be negligible relative to that on
375 Ni. If the sulfide melt was enriched in Cu (Lorand and Grégoire, 2006) there would be no
376 concentration gradient to cause any Cu loss from olivine.

377 It is plausible that a similar mechanism has affected the redistribution of Cr and Ti.
378 As aforementioned, we expect that the melt was not particularly enriched in Cr, and that
379 most of the Cr was partitioned into diopside and Cr-spinel during the early stages of
380 crystallisation. The later stage melt, interacting with the distal olivine porphyroclasts, was
381 therefore likely to be relatively Cr-poor. This reduced concentration gradient meant that re-
382 equilibration was slower. Ti has also been affected by the crystallisation of diopside and
383 ilmenite, with decreasing temperature (cooling post-metasomatism) Ti is preferentially
384 partitioned from olivine to clinopyroxene ((Witt-Eickschen and O'Neill, 2005) and during

385 subsolidus re-equilibration Ti will diffuse out of olivine and into adjacent clinopyroxene
 386 (Cherniak and Liang, 2014). Ti diffuses very slowly in olivine due to its high charge, therefore
 387 explaining the sluggish re-equilibration.

388

389 **6.3 Olivine diffusion anisotropy**

390 The subsolidus re-equilibration of the metasomatic mineral assemblage in BD3067 has
 391 been achieved for many elements but the most systematic variation is observed in Ni and
 392 Cr. These elements have been slower to equilibrate than others, but only in some crystals.
 393 Part of this is due to the influence of late-stage crystallizing phases but the preservation is
 394 also influenced by the anisotropic diffusion of Ni and Cr in olivine.

395 Most studies that have quantified diffusion of Ni in olivine have used the
 396 parameterisation given by Chakraborty (2010) which combines the results of Holzapfel et
 397 al. (2007) and Petry et al. (2004). This parameterisation states that both Ni and Fe-Mg
 398 diffusion are six times faster along the [001] axis than the [100] and [010] axes. Since Ni
 399 diffuses faster along the [001] axis than Fe-Mg, this assumption implies that Ni diffusion
 400 should also be faster in the [100] and [010] axes. This relationship was investigated by
 401 Spandler and O'Neill (2010) who experimentally determined the relationship between
 402 diffusion rate and crystallographic orientation in San Carlos olivine equilibrating with a
 403 silicate melt. They published the diffusion coefficients of 19 elements in each of the three
 404 principle crystallographic axes. Spandler and O'Neill (2010) showed that there is a strong
 405 anisotropy in Ni and Cr diffusion compared to other elements. This is represented in Figure
 406 10a by the higher gradient in logD for Ni and Cr relative to Fe-Mg and Mn. Ni and Cr diffuse
 407 faster along the [001] axis and slower along the [010] axis than Fe-Mg, which magnifies the
 408 order of anisotropy. (Ito and Ganguly, 2006) also found Cr to have anisotropic diffusion in
 409 olivine. The results of Spandler and O'Neill agree with the parameterisations in Chakraborty
 410 (2010) and Dohmen and Chakraborty (2007) for the [001] axis where Ni diffusion is 1.5 times
 411 faster than Fe-Mg, but they differ in the factor of anisotropy for Ni and Cr. The anisotropy
 412 observed by Spandler and O'Neill (2010) is as follows:

$$413 \quad D_{[001]}^{Ni} = 9D_{[100]}^{Ni} = 11D_{[010]}^{Ni}$$

$$414 \quad D_{[001]}^{Cr} = 6D_{[100]}^{Cr} = 17D_{[010]}^{Cr}$$

415 This study highlights the importance of crystallographic orientation as a control on
 416 multi-element diffusion in olivine. The observations of isolated Ni and Cr disequilibrium in
 417 our study support the implication that their diffusion in olivine is faster in the [001] axis and
 418 slower in the [100] and [010] axes, than Fe-Mg, Mn, V, Sc etc. This finding is consistent with
 419 other studies in natural volcanic systems which have also inferred that Ni diffusion can be
 420 slower than Fe-Mg interdiffusion (e.g. Ruprecht and Plank, 2013; Vinet and Higgins, 2010).

421 This greater anisotropy in Ni and Cr, relative to other elements, can be explained by
 422 their ordering in the olivine crystal structure. Ni is preferentially ordered into the M1 site in
 423 olivine, primarily due to its high electronegativity (Bish, 1981; Boström, 1989). The M1 sites
 424 form chains parallel to the [001] axis and the preferred diffusion pathway for Ni is along the
 425 M1 chains (Miyamoto and Takeda, 1983). As a result, Ni diffuses much faster along the [001]
 426 axis than the [100] and [010] axes. The diffusion mechanisms causing the enhanced
 427 anisotropy of Cr in olivine are poorly constrained (Ito and Ganguly, 2006) but it is
 428 understood that Cr³⁺ orders preferentially onto the M1 sites, and as a result has a stronger
 429 anisotropy. Cr²⁺ has an even distribution across M1 and M2 sites and therefore shows less
 430 anisotropy than Cr³⁺ (Jollands et al., 2017). Jollands et al. (2017) find a similar level of
 431 anisotropy between Cr³⁺ and Ni²⁺, hence the correlations observed made in this study imply
 432 that the majority of Cr in the olivine porphyroclasts is Cr³⁺.

433 The anisotropy of M1 ordered elements in olivine can be increased further by silica
 434 activity (Jollands et al., 2017; Zhukova et al., 2014). The silica activity (a_{SiO_2}) of the melt
 435 equilibrating with the olivine impacts the diffusion coefficients of M1 ordered cations
 436 because increased a_{SiO_2} creates more M1 vacancies, which increases the diffusion rates of
 437 M1 ordered cations along the [001] axis (Zhukova et al., 2014). Therefore, the low a_{SiO_2} of
 438 kimberlite and carbonatite magmas is anticipated to decrease diffusion rates of Ni and Cr in
 439 olivine by an order of magnitude (Zhukova et al., 2014; Jollands et al., 2017).

440 The anisotropic diffusion of Ni and Cr in olivine has meant that these elements have
 441 been slow to re-equilibrate in the [100] and [010] (and intermediate) slow axes. This
 442 interpretation is supported by the correlation between the level of Ni variation and the
 443 angle of the crystal relative to the [001] axis (Figure 7).

444

445 **7. Timescales of metasomatism**

446 Timescales of metasomatism are generally estimated by dating minerals or modelling
 447 diffusion profiles. We have determined a U-Pb age of 84 ± 11 Ma for the large zircon shown
 448 in Figure 1e (see Supplementary file E for details). This indicates that the metasomatism
 449 occurred concomitantly with the Late Cretaceous kimberlite activity in Kimberley, including
 450 the eruption of the host Bultfontein kimberlite. To improve on this precision we have used
 451 the preservation of Ni and Cr disequilibrium in the olivine porphyroclasts to estimate the
 452 timing of metasomatism relative to the emplacement of the Bultfontein kimberlite. Since
 453 much more work has been done on Ni diffusion in olivine than Cr we have opted to only
 454 model the Ni diffusion profiles, to estimate the timescales over which the large
 455 porphyroclasts would equilibrate with their local mineral assemblage. Our calculations
 456 assume that Ni diffusion in the xenolith ceased during kimberlite emplacement and cooling.
 457 We have calculated timescales using both the simplistic relationship that the diffusion time
 458 (t) is proportional to the distance (x) squared over diffusion rate (D):

$$459 \quad t \propto x^2/D$$

460 and we have also modelled the diffusion profiles using the following relationship:

$$461 \quad C = (C_0 - C_e) \operatorname{erf}\left(\frac{X}{2\sqrt{Dt}}\right) + C_e \quad [1]$$

462 (Crank, 1956)

463 Where C_0 is the concentration at the core, C_e is the concentration at the rim, X is the
 464 distance from the edge of the crystal, and C is the concentration at position X . In both cases
 465 the diffusion coefficient for Ni in the [001] axis was calculated using the following
 466 parameterisation:

$$467 \quad D_{[001]Ni} = 3.84 \times 10^{-9} \left(\frac{fO_2}{10^{-6}}\right)^{1/4.25} 10^{1.5(X_{Fe}-0.1)} \exp\left(-\frac{220000+(P-10^5)(7 \times 10^{-6})}{RT}\right) \quad [2]$$

468 (Chakraborty, 2010; Petry et al., 2004; Holzapfel et al., 2007)

469 We considered the anisotropy to be close to the factors established by Spandler and O'Neill
 470 (2010), i.e. 10 times slower in the [100] and [001] axes. Another factor we had to account

471 for is that the low silica activity of the metasomatic melt can decrease the diffusion rates.
472 There is no formal parameterisation to take this into consideration for kimberlite melts but
473 based on the conclusions of Zhukova et al. (2014) we have decreased diffusion rates by a
474 factor of 10.

475 The modelling has been carried out at a range of temperature and pressure conditions
476 appropriate to the Kaapvaal geotherm because the mineral assemblage and chemistry of
477 this xenolith is unsuitable for the application of any standard thermo-barometers. Figure 11
478 demonstrates the several orders of magnitude change in equilibration timescales for a 2
479 mm crystal over the temperature interval 900-1250°C. The simple calculations show that
480 homogeneity can be achieved within 1 million years even at the low temperature of 900°C.
481 Our calculations are in good agreement with the diffusion profile modelling, the results of
482 which are shown in Figure 11b-d. The model suggests that at 1000°C metasomatism
483 occurred 200,000-300,000 years prior to kimberlite emplacement, and at 1200°C this
484 decreases to 10,000-20,000 years. This is important because it gives an idea of the extended
485 period of metasomatism by 'failed' kimberlite melts prior to the ascent of the final host
486 kimberlite magma (e.g. Bussweiler et al., 2016; Fitzpayne et al., 2018; Giuliani et al., 2013a;
487 Soltys et al., 2018).

488

489 **8. Reactive infiltration of proto-kimberlite melts**

490 The sub-continental lithospheric mantle has a complex metasomatic history. We
491 suggest that xenolith BD3067 represents an important period of metasomatism prior to
492 eruption of the kimberlite at the surface. Kimberlite eruptions are enigmatic but it is
493 generally agreed that they consist of pulses of magmatism (e.g. Dawson and Smith, 1977;
494 Field et al., 2009; Giuliani et al., 2016, 2014a; Mitchell, 2008, 1991). Polymict breccias from
495 Bultfontein are interpreted to represent 'failed' kimberlite melts that did not make it to the
496 surface (Giuliani et al., 2013a) and recently Jollands et al. (2018) identified two stages of
497 metasomatism preserved in garnet, the second of which was interpreted as reactive
498 infiltration of a silico-carbonate-rich melt, assumed to be a proto-kimberlite melt.
499 (Woodhead et al., 2017) have identified isotopic homogeneity amongst mantle-derived
500 zircon megacrysts from kimberlites in the Kaapvaal craton, they attribute this to a continent-

501 wide metasomatic event that occurred between 114 Ma and several hundred million years
502 ago. Our work provides evidence for a more localised metasomatic event, associated with
503 the Bultfontein kimberlite, that occurred within half a million years of emplacement.

504 We suggest that the clinopyroxene vein in BD3067 crystallized during reactive percolation of
505 a very early pulse of proto-kimberlite melt that infiltrated and reacted with harzburgite wall
506 rock. This involved assimilation of orthopyroxene and diffusive exchange with the olivines in
507 the mantle wall rock, together with fractionation of clinopyroxene from the melt (e.g.
508 (Aulbach et al., 2017; Lim et al., 2018; Simon et al., 2003; Tollan et al., 2015). We suggest
509 that the clinopyroxene crystallised during an early stage of fractionation and the changing
510 mineral assemblage created localised concentration gradients that initiated elemental re-
511 distribution during subsolidus re-equilibration. At a later stage of fractionation, the residual
512 melt became saturated in sulfur (e.g. Aulbach et al., 2017; Giuliani et al., 2013a). As sulfides
513 precipitated the equilibrium conditions changed again and the introduction of a phase with
514 Ni as a stoichiometric component drove the equilibrating olivine Ni concentration even
515 lower. The Cr concentration of the olivines was affected by the initial fractionation of
516 clinopyroxene and Cr-spinel. Many elements experienced diffusive exchange during
517 metasomatism, however the lag in Ni and Cr diffusion that we observe in some olivine
518 porphyroclasts has also been influenced by their anisotropic diffusion in olivine.

519 In this scenario, the newly formed mineral assemblage achieves local equilibrium
520 over time, until another kimberlite pulse utilises the pathway generated by thousands of
521 years of metasomatism to ascend to the surface (Giuliani et al., 2016, 2014a). The rapid,
522 violent ascent of the kimberlite fractures the wall-rock and the metasomatised mantle
523 becomes entrained in the kimberlite. Figure 12 shows a summary of the metasomatic
524 history of xenolith BD3067.

525

526 **9. Comparison with previous estimates for the duration of pre-emplacment** 527 **metasomatism**

528 A number of attempts have been made to constrain the precise timescales of the
529 metasomatism that accompanies kimberlite activity. Most studies have applied diffusion
530 modelling to major- and trace-element disequilibrium in pyrope garnet (Griffin et al., 1996;

531 Smith et al., 1991; Smith and Ehrenberg, 1984), although Konzett et al. (2013) have adopted
532 the same approach to Sr zoning in K-richterite. These studies concluded that metasomatic
533 processes occur on timescales of up to 10^4 years before the eruption of the kimberlite host.
534 Recent modelling of Ni diffusion in garnets by Jollands et al. (2018) suggests that
535 metasomatism by a “failed” kimberlite occurred between 25 days and 400 years prior to
536 host kimberlite entrainment. Giuliani et al. (2013) modelled Ni-rich metasomatism
537 associated with the Bultfontein kimberlite, they present similar profiles to Figure 5 but show
538 that Ni has diffused into the olivine from Ni-rich melts. Their modelling implies that
539 metasomatism occurred 100 years (at 1000°C) to 1.7 Myr (at 700°C) prior to emplacement
540 of the Bultfontein kimberlite.

541 The ten to hundred thousand year timescales of equilibration estimated in this study
542 are comparable to, or slightly longer than, those derived from previous studies (Cordier et
543 al., 2015; Andrea Giuliani et al., 2013; Griffin et al., 1996; Jollands et al., 2018; Smith and
544 Ehrenberg, 1984). The olivine disequilibrium related to crystallographic orientation that we
545 observed in our detailed study of BD3067 adds an extra constraint (c.f. garnets that have a
546 more uniform compositional zonation due to their isotropic structure). The range of
547 timescales provided both here and in previous studies supports the hypothesis that a period
548 of metasomatism by ‘failed’ kimberlite melts precedes the final kimberlite eruption, and
549 that this period is necessary to provide a lubricated pathway for the ‘successful’ kimberlite
550 to exploit (Giuliani et al., 2016, 2014a).

551

552 **10. Conclusions**

553 Our investigation of the processes associated with the infiltration of metasomatic melts
554 in sub-cratonic lithospheric mantle is focused on a peridotite xenolith that represents a
555 relatively rare, incomplete stage of melt-rock reaction. The xenolith (BD3067), which was
556 brought to the surface by the Late Cretaceous Bultfontein kimberlite (South Africa), contains
557 a spectacular metasomatic vein of diopside, sulfide, phlogopite, spinel and zircon set in a
558 dunite host that we interpret as a relict melt channel. From our detailed petrographic study
559 and systematic *in-situ* micro-analyses of major, minor and trace elements we conclude that
560 the vein assemblage in BD3067 crystallised from a percolating proto-kimberlite (silico-

561 carbonate) melt. Reactive infiltration of this small-fraction silica-poor melt appears to have
562 caused the dissolution of orthopyroxene and precipitation of olivine neoblasts (Fo_{88}).

563 Unique information on the variable rates of diffusion of major, minor and trace
564 elements during sub-solidus re-equilibration that follows mantle metasomatism is preserved
565 in relict olivine porphyroclasts (Fo_{88}). These have homogeneous concentrations of Mg, Fe
566 and Mn but exhibit significant core to rim decreases in both NiO (0.32 – 0.18 wt.%), and Cr
567 (60 – 35 ppm). We propose that this heterogeneity is strongly influenced by concentration
568 gradients in Ni and Cr away from the main melt channel and also the localised crystallisation
569 of sulfides and spinel. Our EBSD study of the olivine porphyroclasts shows that Ni and Cr
570 zoning is only evident in grains that contain large components of the [100] and [010] slow
571 diffusion axes.

572 Diffusion models suggest the preservation of the observed core-to-rim variation of
573 Ni in olivine requires that mantle metasomatism must have occurred within the order of 10^5
574 years prior to emplacement of the host Bultfontein kimberlite. This metasomatism may
575 have been fundamental to the subsequent ascent and emplacement of the host kimberlite.
576 The time constraint on preservation of core-to-rim variations in Ni and Cr combined with the
577 crystallographic control on diffusion may explain why Ni and Cr heterogeneity in otherwise
578 homogeneous mantle olivine has been so rarely encountered.

579 A broader implication of our study is that the magnitude of anisotropy in diffusion
580 for Fe-Mg in olivine is different to highly-ordered elements such as Ni, which appear to
581 diffuse much more slowly along the [100] and [010] axes. This is of significance to both
582 mantle and magmatic olivines and highlights the importance for diffusion studies that
583 combine chemical zonation with crystallographic information.

584

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593

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893 **13. Figure Captions**

894 **Figure 1:** Images of the veined xenolith BD3067. (a) Image of the hand specimen of mantle
895 xenolith BD3067 together with locations of (b) the proximal and (c) distal sections as
896 described in the main text. (b) Photograph of a large, 13 by 9 cm, thin section containing a
897 trichotomously branching vein; (c) Photomicrographs of olivine porphyroclasts (ol pc) and
898 neoblasts (ol nb) (cross polarised light); (d) Complex deformation in clinopyroxene (cross
899 polarised light); (e) Cathodoluminescence image of large zircon crystal within the main vein.

900

901 **Figure 2:** Clinopyroxene variation diagrams (a) CaO vs. SiO₂; (b) Na₂O vs. SiO₂; (c) FeO vs.
902 SiO₂; (d) Cr₂O₃ vs. SiO₂; (e) Ti vs. Al; (f) Ni vs. Al. Clinopyroxene from the main vein are
903 plotted as red circles, and those away from the main vein in blue diamonds.

904

905 **Figure 3:** Primitive mantle normalised trace-element plot of clinopyroxenes in BD3067. The
906 composition of clinopyroxenes found in xenoliths from Finsch Mine are shown for comparison
907 (Gibson et al., 2008).

908

909 **Figure 4:** Olivine variation diagrams (a) NiO vs. Fo content; (b) MnO vs. Fo content; (c) Al vs.
910 Ni; (d) Cr vs. Ni; (e) Cu vs. Ni; (f) Ti vs. Ni. The legend in (a) reflects the textural and location
911 (proximal or distal) of each measurement; pc stands for porphyroclast. NiO, MnO and Fo
912 content have been measured using EPMA; Ni, Al, Cr, Cu, Ti have been measured using LA-
913 ICP-MS. These are representative measurements that can be found in Supplementary File B.

914

915 **Figure 5:** Olivine profiles taken of rim to rim transects in four olivine crystals. Three distal
916 olivines are shown (a) Olivine A; (b) Olivine B; (c) Olivine D, and one proximal olivine (d)
917 Olivine M. The black circles represent the measurements for the y axis on the left-hand side
918 (e.g. Mg# in column 1), and the red triangles are associated with the y axis on the right-hand
919 side (e.g. MnO in column 1). Mg# (MgO/(MgO+FeO) mol) and MnO are measured using

920 EPMA and Ni, Cu, Cr and Ti are LA-ICP-MS measurements. Olivines A and D show strong
921 variation in Ni, Cr and Ti across the crystal whereas Olivines B and M are homogeneous.

922

923 **Figure 6:** Back-Scatter Electron (BSE) image (6a, e) and EDS maps showing the Fe (6b, f), Ni
924 (6c, g) and S (6d, h) distribution in two selected sulfides from BD3067.

925

926 **Figure 7:** (a) The range in NiO (Δ NiO) across the profile in each crystal plotted against the
927 orientation of the [001] axis. The position of the [001] is also shown in the stereonet
928 visualisation for (b) Olivine A and (c) Olivine B. Each stereonet shows the plunge and trend of
929 the <100> axes relative to the crystal face and the strike and dip of the (001) plane. The star
930 shows the orientation of the profile analysed.

931 **Figure 8.** Forsterite content vs. (a) NiO and (b) MnO for olivines in BD3067 compared to
932 xenocrystic olivine cores and magmatic olivine rims from Kimberlites in the Kimberley region
933 (Giuliani, 2018)

934

935 **Figure 9.** Primitive mantle normalised trace-element plot comparing the melt in equilibrium
936 with the clinopyroxenes with melt compositions of the host Bultfontein kimberlite (Roex et
937 al., 2003) and a global average carbonatite composition (Bizimis et al., 2003). The four
938 equilibrium melts plotted use (a) cpx-carbonatite melt partition coefficients from Dasgupta
939 et al. (2009) and cpx-silicate melt partition coefficients from Suzuki et al. (2012); and (b) cpx-
940 silicate-carbonate melt partition coefficients from Giris et al. (2013) and cpx-CO₂-rich
941 kimberlite melt partition coefficients from Keshav et al. (2005). Details of these can be found
942 in supplementary file D..

943

944 **Figure 10. (a)** Diffusion coefficients for a selection of major, minor and trace elements in
945 olivine, between the [010] axis (at 0°) and the [001] axis (at 90°) as determined by Spandler
946 and O'Neill (2010). The elements are labelled on the diagram, including Fe-Mg
947 interdiffusion. The level anisotropy is depicted in the gradient of the line for each element.

948 (b) The olivine structure, showing the position of the M1 and M2 sites relative to the
949 tetrahedra (pink triangle) in the (001) and (010) orientations. The M1 sites form chains
950 parallel to (001), the preferred diffusion pathway for M1 ordered cations, such as Ni and
951 Cr^{3+} , is along these chains.

952

953 **Figure 11.** (a) Bultfontein geotherm (Mather et al., 2011) with annotations to show the
954 relative times required for Ni to diffuse in olivine over 2 mm at various points in the
955 lithosphere, controlled by the temperature. (b-d) Diffusion model showing estimates of
956 times required to achieve the NiO zoning observed in the steepest limbs (hence slowest –
957 gives minimum time) of the profiles in (b) Olivine A at 1000°C; (c) Olivine D 1000°C and (d)
958 Olivine A at 1200°C. The diffusion is modelled using the equations of Chakraborty et al.
959 (2010) and Crank et al. (1956) and the method is described in the main text.

960

961 **Figure 12.** Schematic diagram summarising the history of the veined xenolith from
962 Bultfontein (BD3067). The emplacement of the Bultfontein kimberlite is preceded by a
963 period of melt infiltration and metasomatism. The metasomatism changes the equilibrium
964 conditions by altering the mineral assemblage. The subsequent re-equilibration takes place
965 over a short timescale, 10^3 - 10^5 years, and the disequilibrium can be preserved if the xenolith
966 is entrained before equilibration has been achieved.

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975 **Table 1: Representative major, minor and trace element composition of clinopyroxene in**
 976 **BD3067.**

Crystal Location *	cpx_10 / 1	cpx_18 / 1	cpx_19 / 1	cpx_31 / 1	BD3067CPX 1	BD3067CPX 2	BD3067CPX4 a	BD3067CPX 5
	NV	NV	NV	NV	V	V	V	V
EPMA (wt.%)								
SiO ₂	54.67	55.52	55.01	55.39	54.46	54.57	54.13	54.37
TiO ₂	0.22	0.20	0.19	0.20	0.20	0.20	0.20	0.19
Al ₂ O ₃	0.62	0.66	0.67	0.60	0.62	0.66	0.60	0.63
Cr ₂ O ₃	1.85	1.13	0.80	0.83	0.74	0.73	0.85	0.59
FeO	2.85	3.02	3.01	2.62	3.04	3.00	3.11	2.98
MnO	0.06	0.09	0.09	0.10	-	-	-	-
MgO	15.90	16.23	16.68	16.69	16.55	16.65	16.61	16.71
NiO	b.d.l	b.d.l	b.d.l	b.d.l	0.02	0.02	0.02	0.02
CaO	21.60	22.17	21.97	22.52	22.19	22.21	22.00	22.17
Na ₂ O	1.49	1.25	1.10	1.11	1.22	1.09	1.22	1.15
Mg#	90.9	90.5	90.8	91.9	90.6	90.8	90.5	90.9
LA-ICP-MS (ppm)								
Al	3264	3179	3571	3448	3478	3438	3470	3516
Sc	38.8	43.4	54.5	53.9	57.0	54.3	52.0	54.4
Ti	1046	1023	1061	1144	1239	1256	1398	1356
V	261.3	328.3	432.7	418.8	373.3	359.2	353.0	391.5
Co	18.70	21.50	19.83	22.06	18.65	17.96	18.38	19.55
Ni	350.8	231.8	248.9	250.6	177.5	176.3	197.8	163.8
Sr	171.2	219.6	154.9	203.6	153.3	161.2	158.6	146.2
Y	4.96	4.70	3.93	5.30	3.97	3.89	4.37	4.25
Zr	139.6	110.4	109.8	121.9	99.7	85.8	109.6	92.9
Nb	0.38	0.91	0.88	1.62	0.209	0.187	0.374	0.209
La	2.44	3.84	3.24	4.89	2.21	2.45	2.65	2.16
Ce	12.23	17.15	13.67	19.88	10.22	11.46	11.76	10.20
Pr	2.12	2.89	2.34	3.18	2.002	2.233	2.233	1.947
Nd	11.24	14.61	11.61	15.57	10.98	11.97	12.34	10.81
Sm	2.93	3.67	2.91	3.66	2.816	3.036	3.256	2.926
Eu	0.98	1.13	0.87	1.20	0.891	0.902	0.924	0.913
Gd	2.46	2.89	1.97	2.83	2.398	2.42	2.563	2.486
Tb	0.327	0.340	0.263	0.367	0.264	0.297	0.308	0.297
Dy	1.59	1.59	1.35	1.61	1.34	1.34	1.46	1.40
Ho	0.225	0.202	0.176	0.238	0.176	0.187	0.198	0.22
Er	0.486	0.435	0.335	0.387	0.341	0.33	0.363	0.396
Tm	0.046	0.036	0.038	0.040	0.033	0.033	0.033	0.033
Yb	0.204	0.160	0.207	0.196	0.143	0.143	0.187	0.165
Lu	0.019	0.016	0.014	0.018	0.011	0.011	0.022	0.022
Hf	7.56	6.16	6.85	6.86	6.28	5.25	6.70	5.83
Ta	0.017	0.045	0.052	0.083	0.011	0.011	0.022	0.022
Pb	0.272	0.416	0.403	0.507	0.308	0.264	0.308	0.264

977 *NV = not in main vein; V = in main vein; b.d.l = below detection limit; - not measured

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980 **Table 2: Representative major, minor and trace elements of olivine porphyroclasts and**
 981 **neoblasts in BD3067.**

Crystal	olivine_a	olivine_a	olivine_b	olivine_b	olivine_d	olivine_d	olivine_g	olivine_g
Location	distal							
pc/nb*	pc							
position	core	rim	core	rim	core	rim	core	rim
EPMA								
(wt.%)								
MgO	47.16	50.00	47.12	47.26	48.11	47.82	47.45	47.46
SiO2	40.62	43.25	40.64	40.73	40.66	40.22	39.92	39.78
CaO	b.d.l	0.09	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
NiO	0.32	0.21	0.21	0.18	0.30	0.20	0.28	0.21
MnO	0.14	0.15	0.18	0.16	0.19	0.16	0.17	0.15
FeO	11.36	11.31	11.57	11.61	11.65	11.62	11.60	11.47
Fo	88.1	88.7	87.9	87.9	88.0	88.0	87.9	88.1
LA-ICP-MS (ppm)								
Na	88.6	56.3	58.5	89.3	85.2	65.1	67.4	47.5
Al	7.99	7.87	3.70	7.00	10.54	6.35	9.07	7.71
P	87.6	77.6	100.4	116.8	104.7	85.7	116.1	56.5
Sc	1.55	1.68	1.40	1.54	1.31	1.28	1.80	1.83
Ti	168.9	145.6	135.3	148.0	143.9	125.0	154.0	141.3
V	3.72	3.81	3.47	3.67	3.44	3.51	3.71	4.09
Cr	52.7	37.8	34.1	35.4	49.0	32.6	42.4	37.9
Mn	1038	1038	1042	1117	1053	966	1059	1060
Co	129.6	136.3	128.6	140.0	133.7	120.0	136.9	134.6
Ni	2157	1526	1438	1538	2089	1381	1905	1523
Cu	1.07	1.14	1.12	1.07	1.12	1.14	0.93	1.19
Zn	104.1	105.3	96.8	109.2	102.5	91.1	122.2	112.5
Y	0.0331	0.0196	b.d.l	0.0291	0.046	b.d.l	b.d.l	b.d.l
Zr	0.592	0.288	0.412	0.629	0.549	0.345	0.398	0.466
Nb	0.521	0.289	0.267	0.442	0.584	0.298	0.33	0.257
Sn	0.74	0.737	0.642	0.834	0.577	0.627	0.979	0.766
Pb	0.013	b.d.l	b.d.l	0.012	b.d.l	0.008	b.d.l	0.021
EPMA								
(wt.%)								
MgO	47.69	47.41	47.50	47.52	48.08	47.49	47.69	47.83
SiO2	41.10	40.32	40.44	40.63	40.24	40.86	40.45	40.53
CaO	b.d.l	b.d.l	b.d.l	b.d.l	0.03	b.d.l	b.d.l	0.04
NiO	0.22	0.20	0.19	0.23	0.20	0.20	0.20	0.20
MnO	0.14	0.17	0.11	0.16	0.15	0.14	0.16	0.15
FeO	11.59	11.33	11.43	11.55	11.62	11.49	11.39	11.89
Fo	88.0	88.2	88.1	88.0	88.1	88.0	88.2	87.8

982 *pc = porphyroclast; nb = neoblast; b.d.l = below detection limit

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985 **Table 3:** Major element compositions of Ni-rich regions and exsolution lamellae in sulfides.

Section Sulfide/ Point	Tk5	Tk4	Tk4	Tk4	Tk4	Tk4
	45 / 1	28 / 1.	32 / 1	27 / 21.	27 / 25	27 / 27.
S (wt.%)	18.93	12.45	15.75	16.65	20.20	27.13
Ni	24.60	11.63	15.53	21.26	28.19	34.33
Fe	53.19	74.36	60.71	55.69	50.51	32.69
Si	2.46	0.94	4.07	5.95	0.43	5.15
Cu	0.12	0.38	3.67	0.09	0.22	0.10
Cr	0.71	0.23	0.28	0.36	0.45	0.60
NiS (mol. %)	30.38	13.24	18.61	25.70	35.20	43.71
FeS	66.34	85.20	73.37	67.90	63.70	42.08

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987 **Supplementary File A:** Analytical Techniques988 **Supplementary File B:** All mineral chemistry, olivine and clinopyroxene EPMA and LA-ICP-MS
989 analyses and olivine profiles.990 **Supplementary File C:** Graphical plots of each olivine profile. Supplementary to Figure 5.991 **Supplementary File D:** Calculated composition of the melt in equilibrium with BD3067
992 clinopyroxenes.993 **Supplementary File E:** Zircon dating

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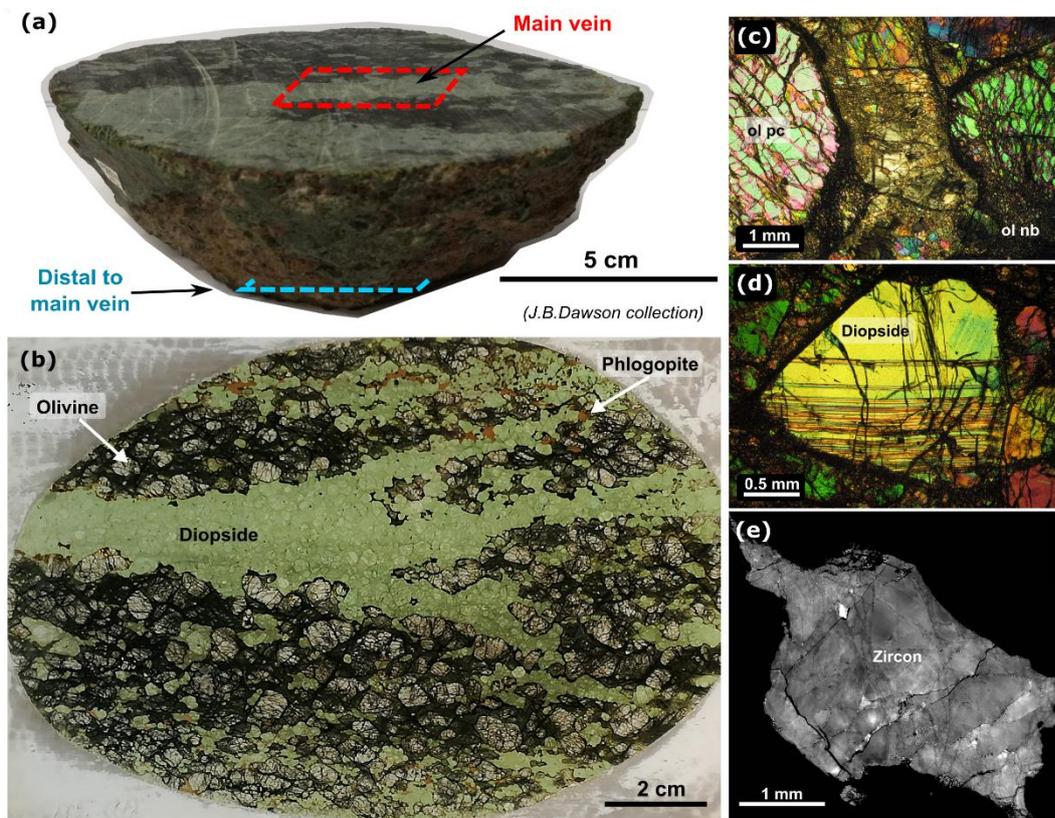
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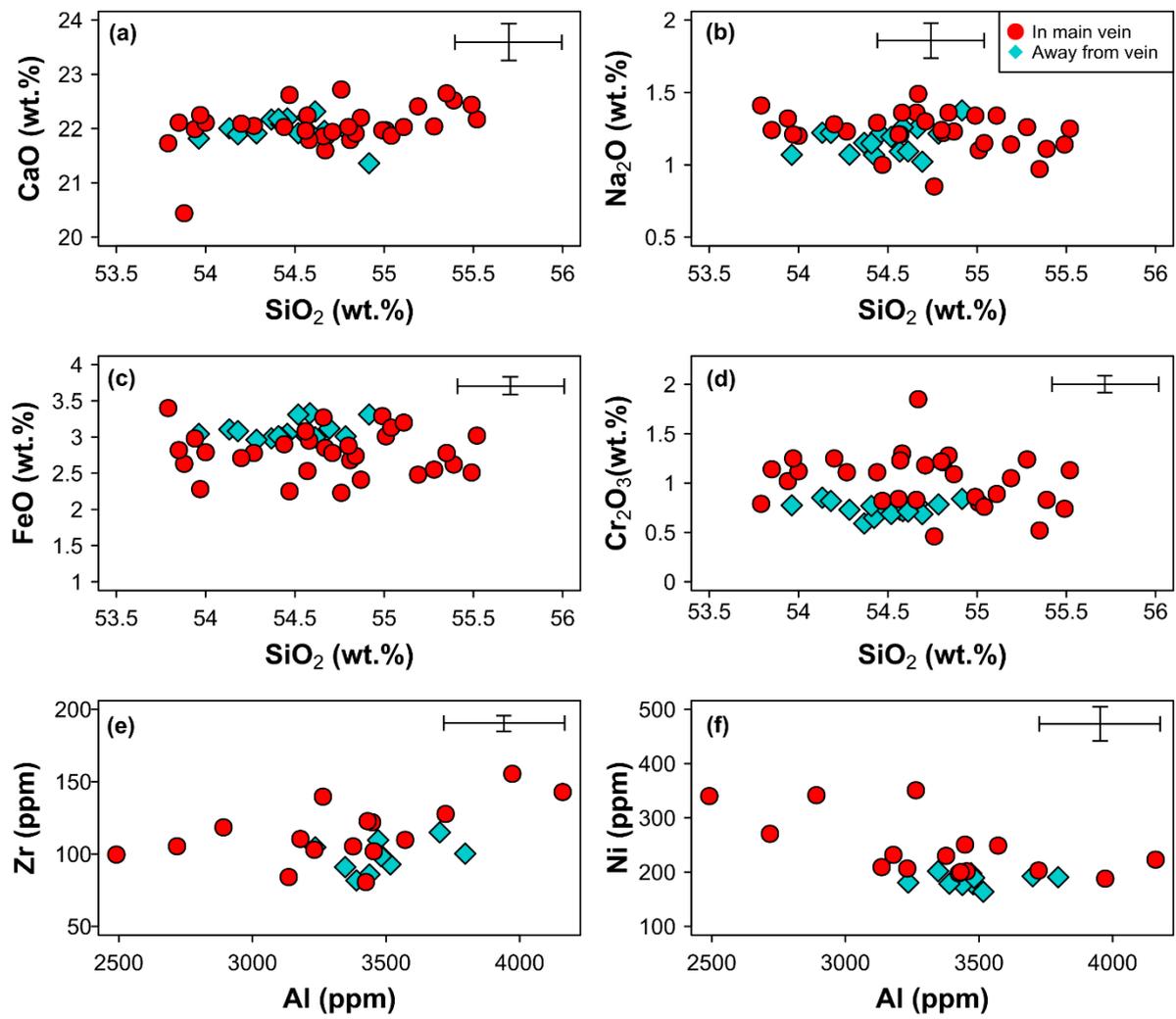
1003 Figure 1:



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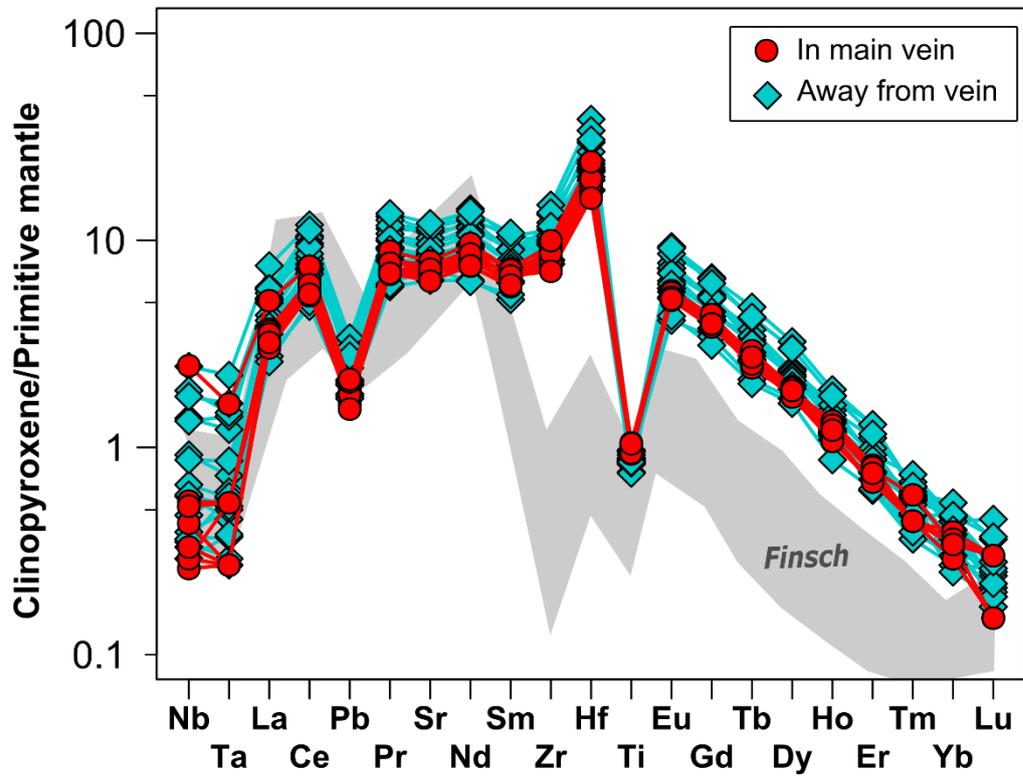
1007 Figure 2:



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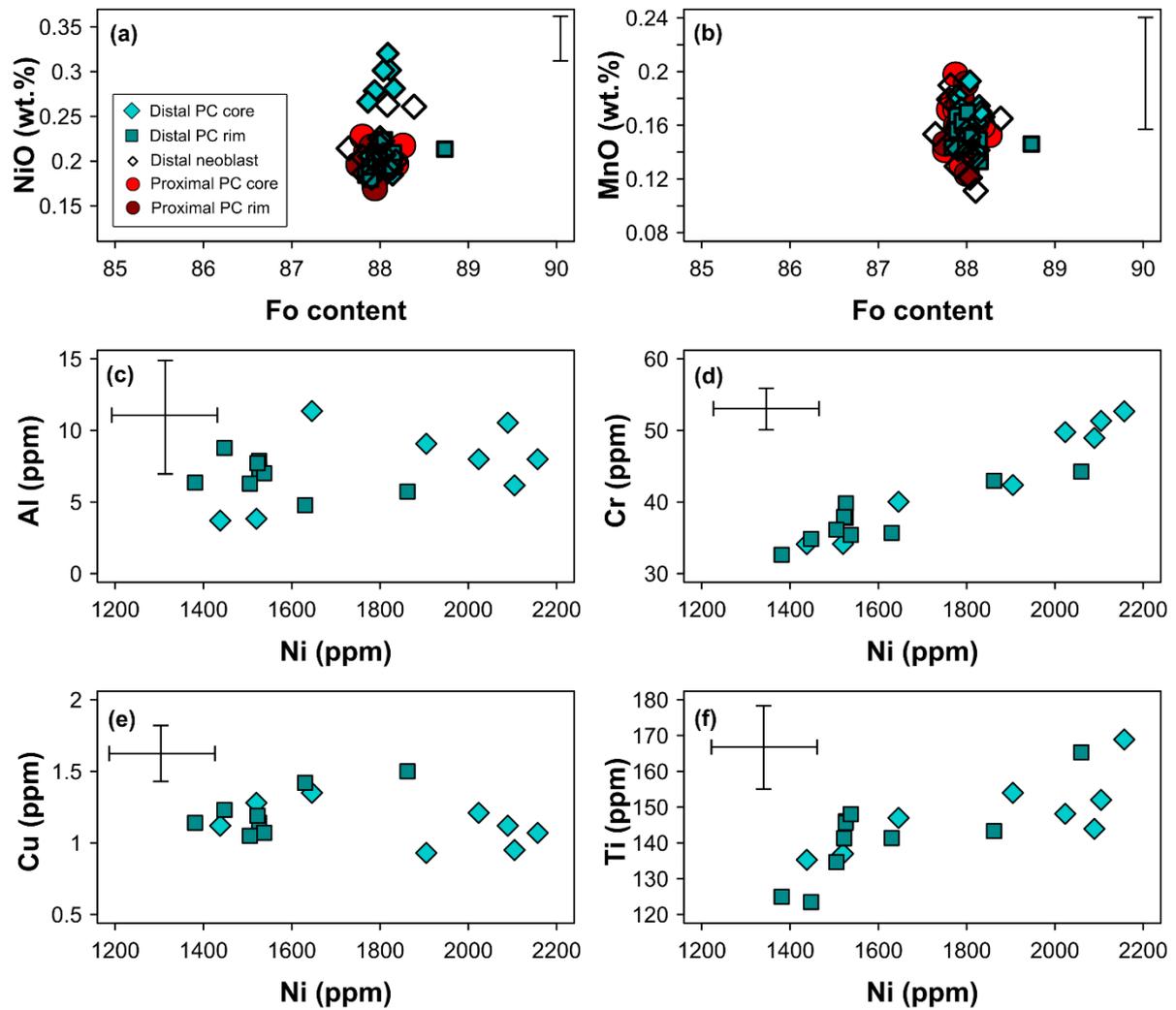
1010 Figure 3:



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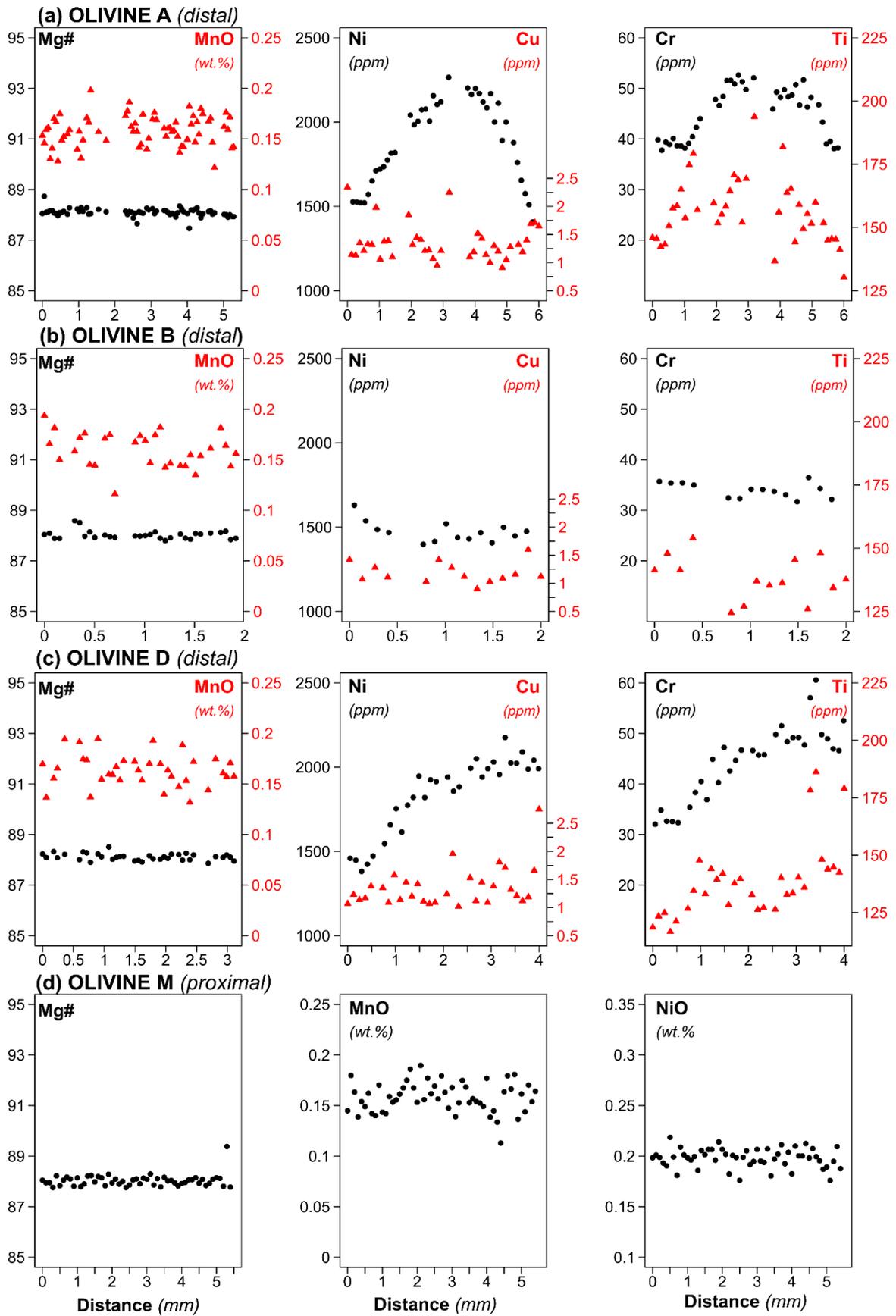
1013 Figure 4:



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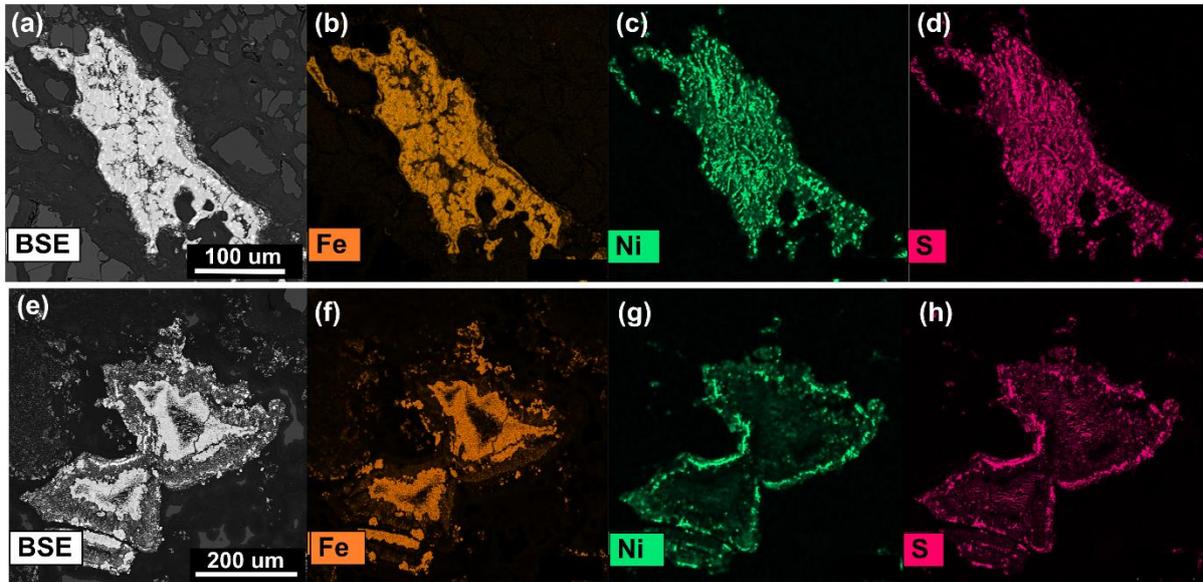
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1016 Figure 5:



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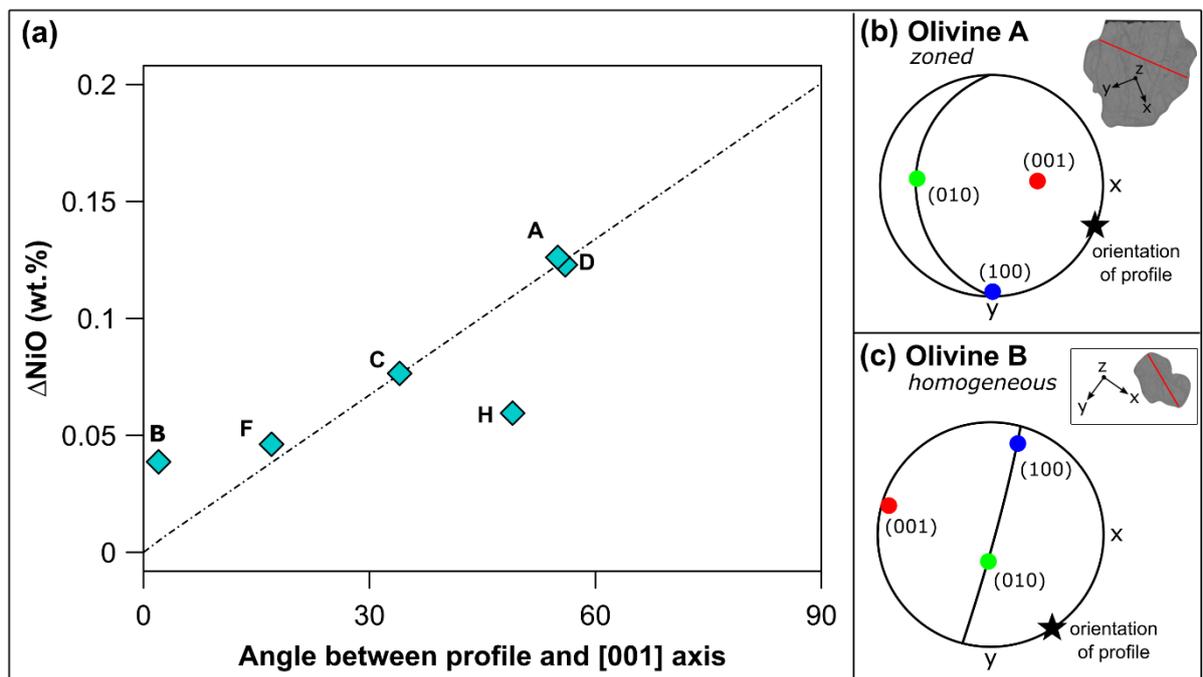
1018 Figure 6:



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1021 Figure 7:

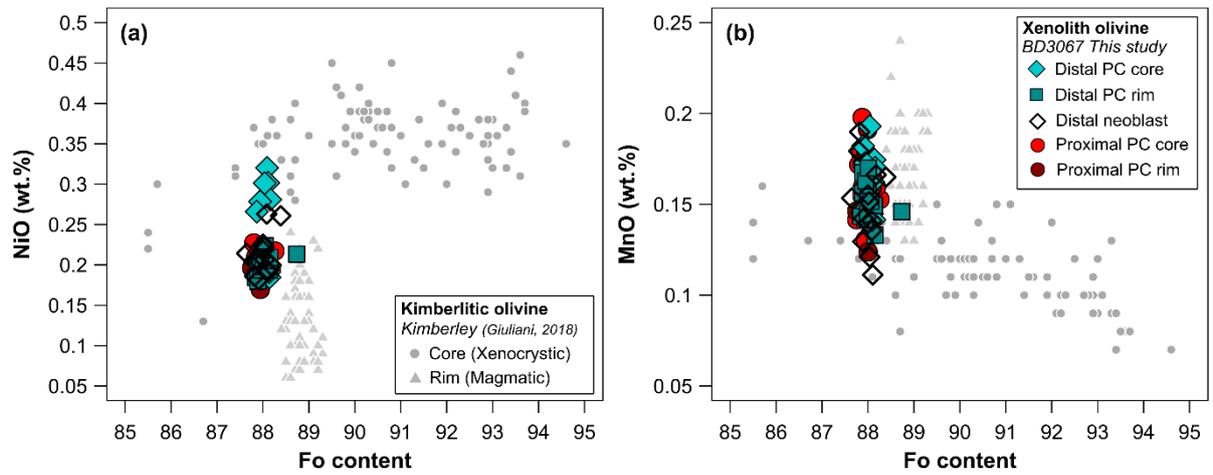


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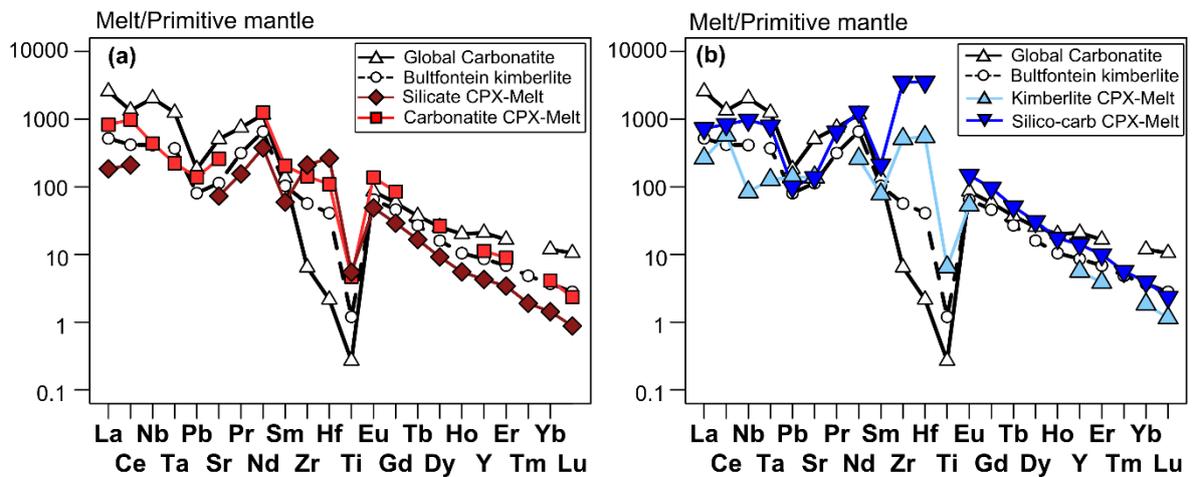
1025 Figure 8:



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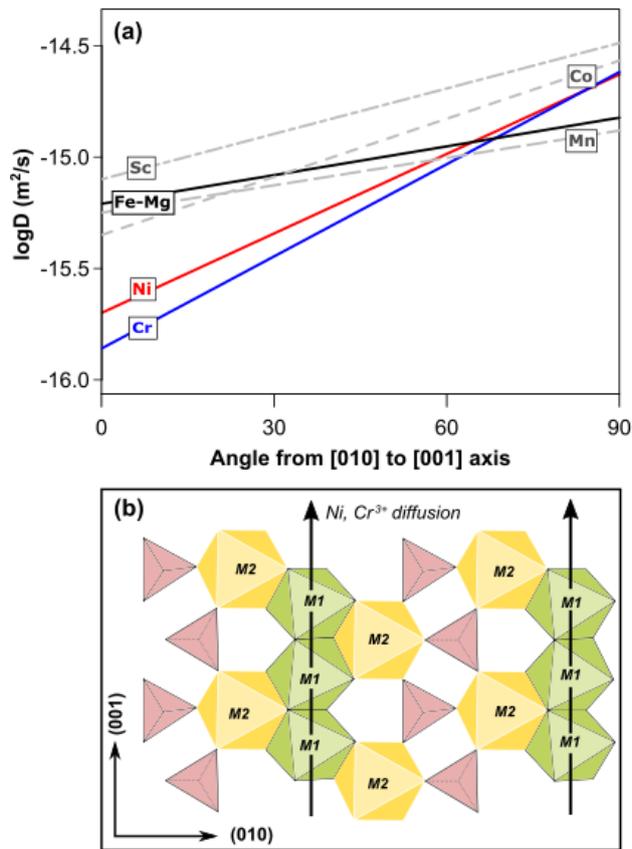
1028 Figure 9:



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1031 Figure 10:

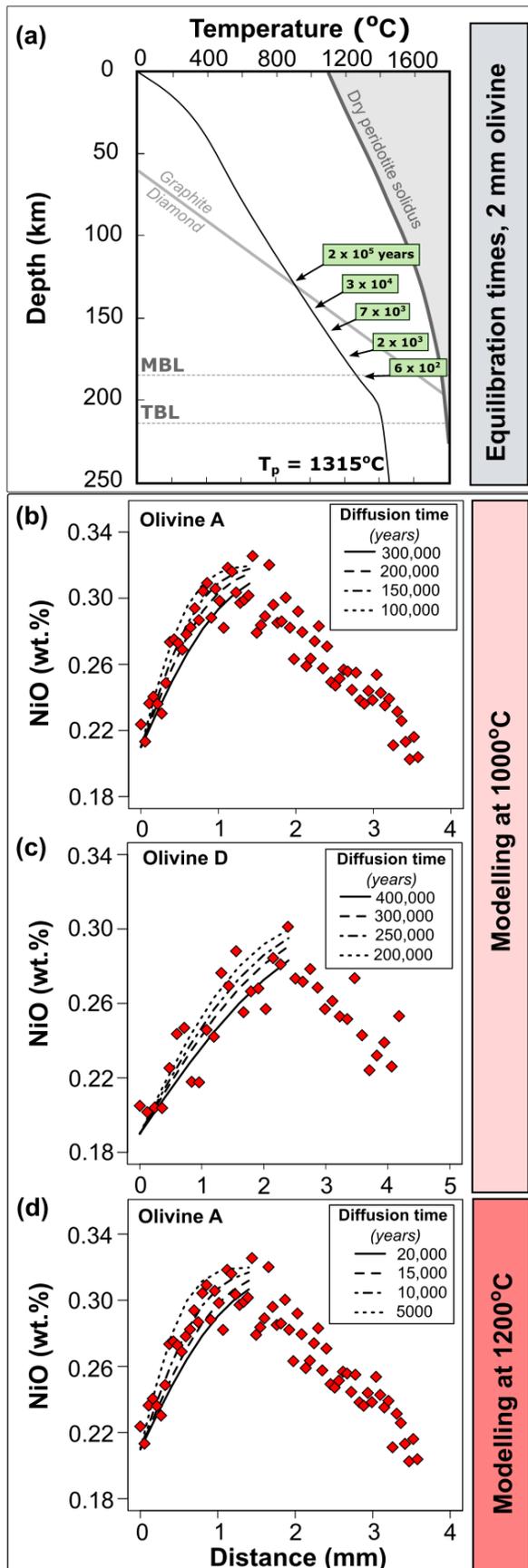


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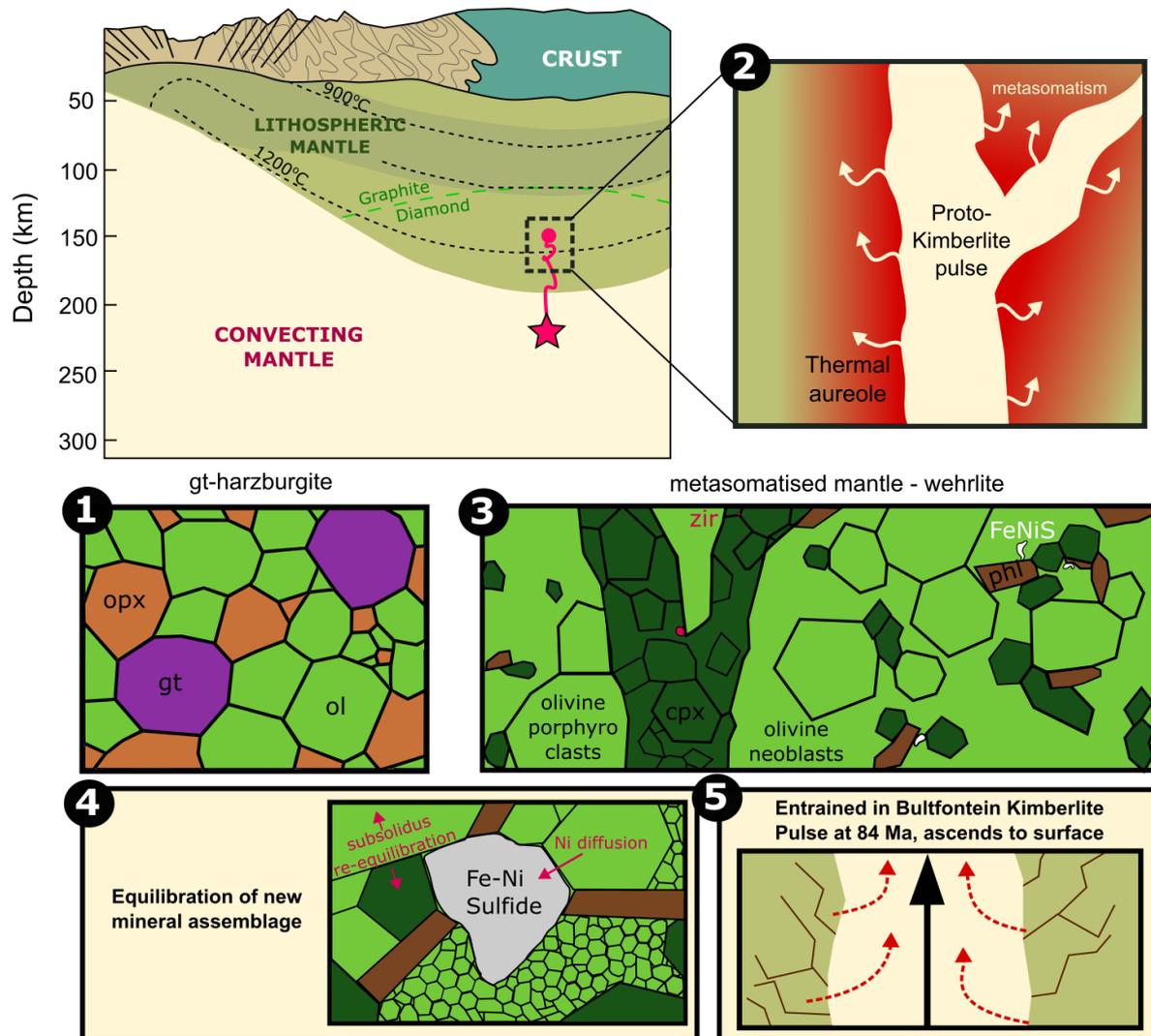
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1035 Figure 11:



1037 Figure 12:



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